



$U(1)'$ seesaw and RHsN DM

KIAS

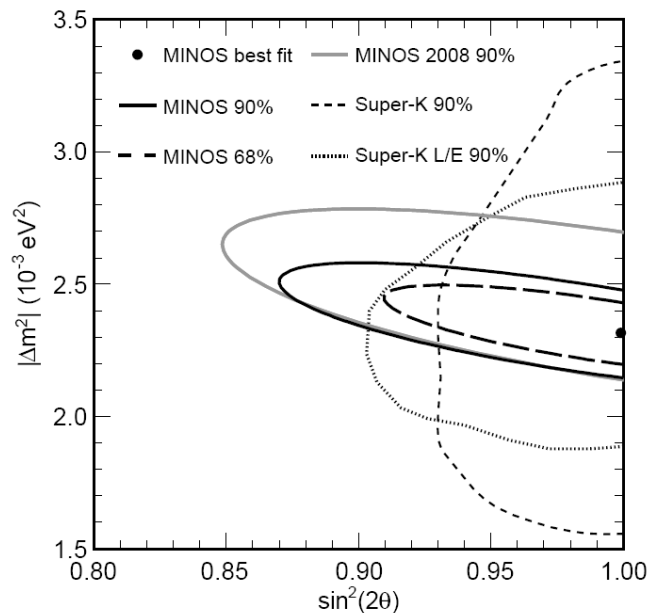
Korea Institute for Advanced Study

Eung Jin Chun

P. Bandyopadhyay, EJC, J.C. Park, arXiv:1105.1652 [hep-ph]

Neutrinos

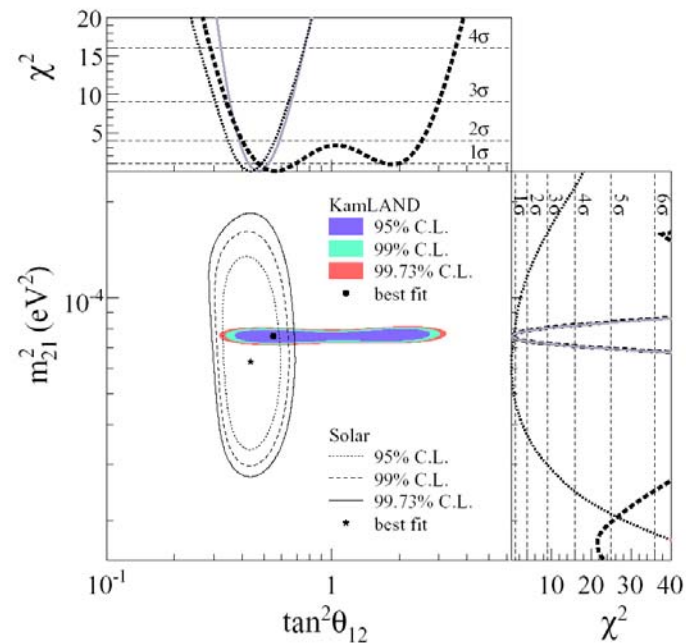
MINOS (2011)



$$|\Delta m_{31}^2| = 2.32^{+0.12}_{-0.08} \times 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta_{23} > 0.90$$

KAMLAND (2008)



$$\Delta m_{21}^2 = 7.59^{+0.21}_{-0.21} \times 10^{-5} \text{ eV}^2$$

$$\tan^2 \theta_{12} = 0.47^{+0.06}_{-0.05}$$



First results from the Double Chooz experiment

H. De Kerret

LowNu, Seoul, Nov. 11

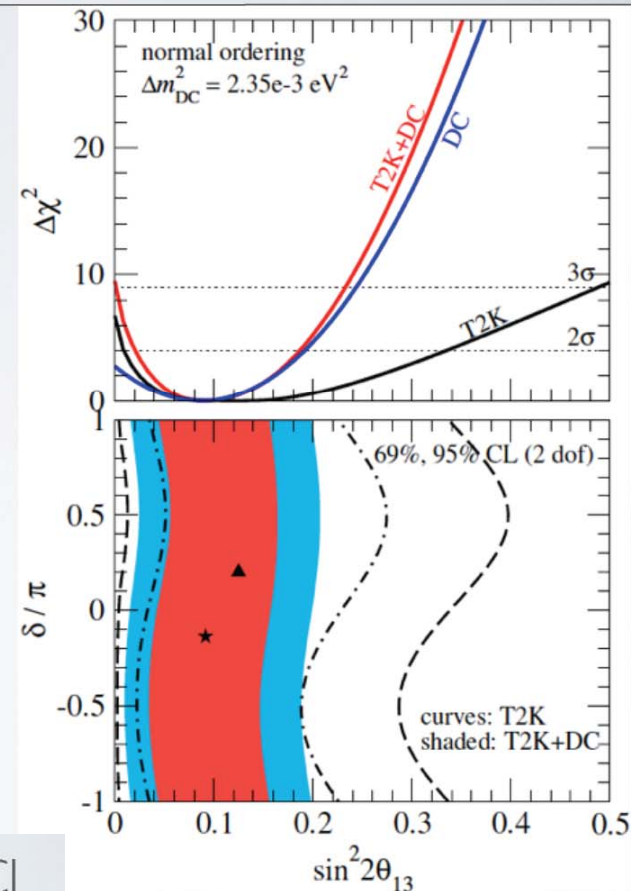
θ_{13}

Global θ_{13} Hunt

DC results consistent with
T2K results

Major impact in θ_{13}
knowledge today by DC

Combined effect: $\sin^2(2\theta_{13})$
>0 @ 3σ



$\sin^2(2\theta_{13}) = 0.085 \pm 0.029(\text{stat}) \pm 0.042(\text{syst})$ @ 68%CL

Status of the **RENO** Reactor Neutrino Experiment

RENO = **R**eactor **E**xperiment for **N**eutrino **O**scillation

(For RENO Collaboration)

(Inauguration for RENO Experiment)



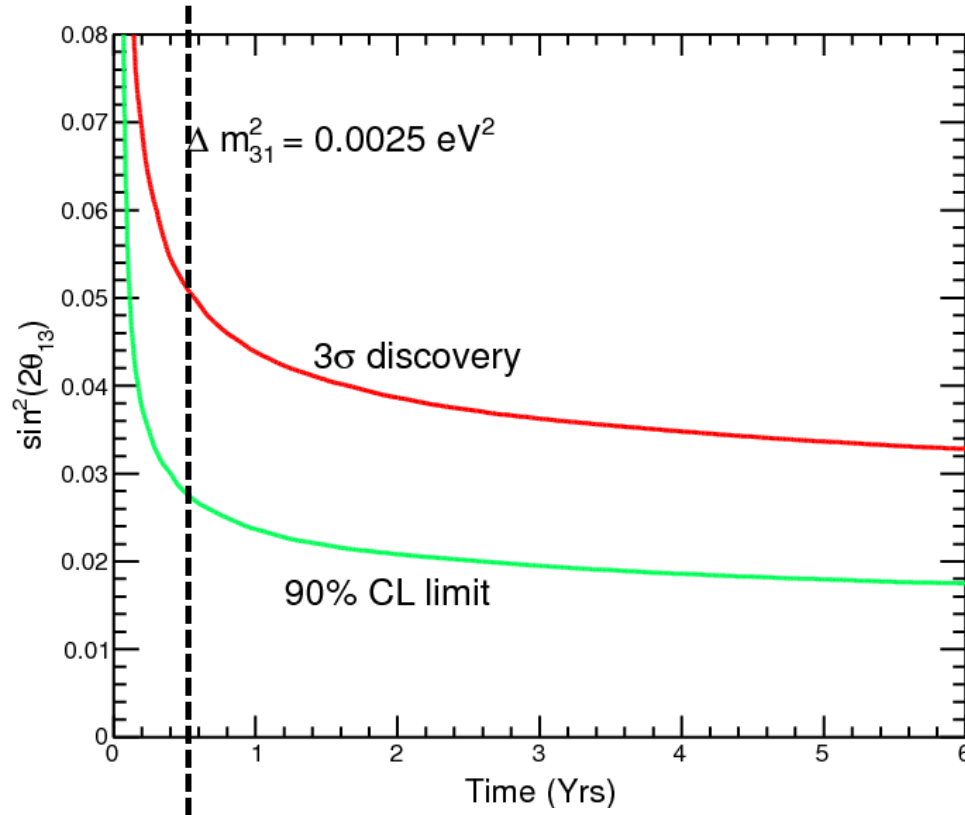
K.K. Joo

Chonnam National University

November 09, 2011

6th Low Energy Neutrino Physics, November 09-12 2011, SNU

- First results on $\sin^2(2\theta_{13}) \sim 0.05$ are expected to be available within a half year
 - Goal: ~ 0.07 (end of this yr)
 - ~ 0.05 (March, 2012)



- RENO group hope to tell the value of $\sin^2(2\theta_{13})$ at the anticipated time
 - Goal: Neutrino 2012 @ Kyoto (June, 2012)

Dark Matter

► Unidentified dark matter 23%

$$\Omega_{DE} h^2 = 0.361 \pm 0.008$$

$$\Omega_m h^2 = 0.1349 \pm 0.004$$

$$\Omega_B h^2 = 0.0226 \pm 0.0008$$

$$\Omega_\nu h^2 = 0.0005 \sim 0.0047$$

$$\Omega_{DM} h^2 = 0.1123 \pm 0.0035$$

$$\Omega_r h^2 = 2.469 \times 10^{-5}$$

WMAP7 + BAO + H_0
Komatsu et al. 2010

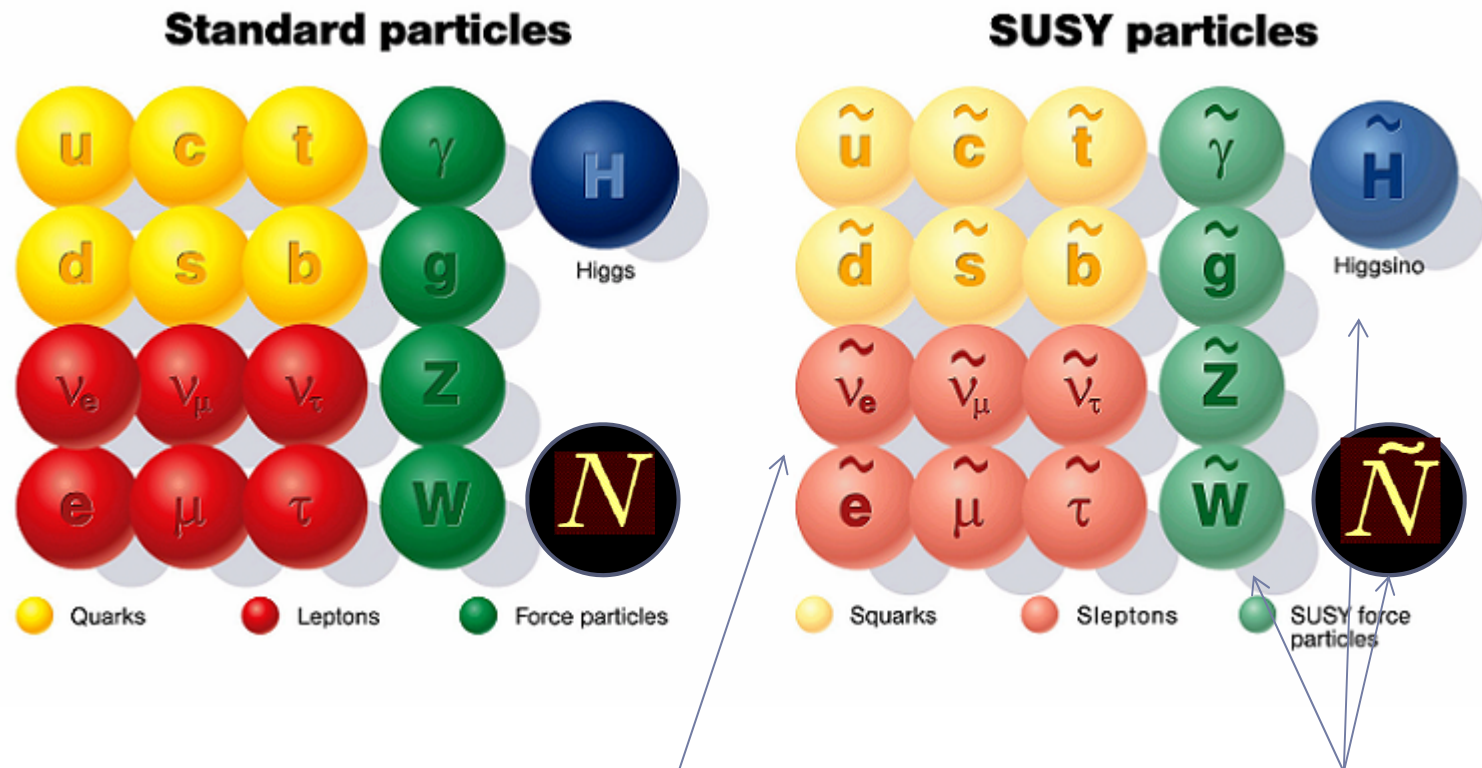
Why RHsN DM?

- ▶ Combines two evidences for physics BSM.
- ▶ It may be the LSP.
- ▶ It may be tested at LHC.

clean seesaw signatures:

same-sign dileptons, displaced vertices + MET

SUSY DM



Large Z^0 interaction excluded by direct detection

Good DM candidates

The Model: SUSY $U(1)'$ seesaw at TeV

- ▶ $U(1)' = U(1)_{B-L}, U(1)_\chi$, or $U(1)_\psi$ from GUT.

$$W = y_\nu L H_2 N + \frac{1}{2} \lambda S N N$$

$$m_N = \lambda \langle S \rangle$$

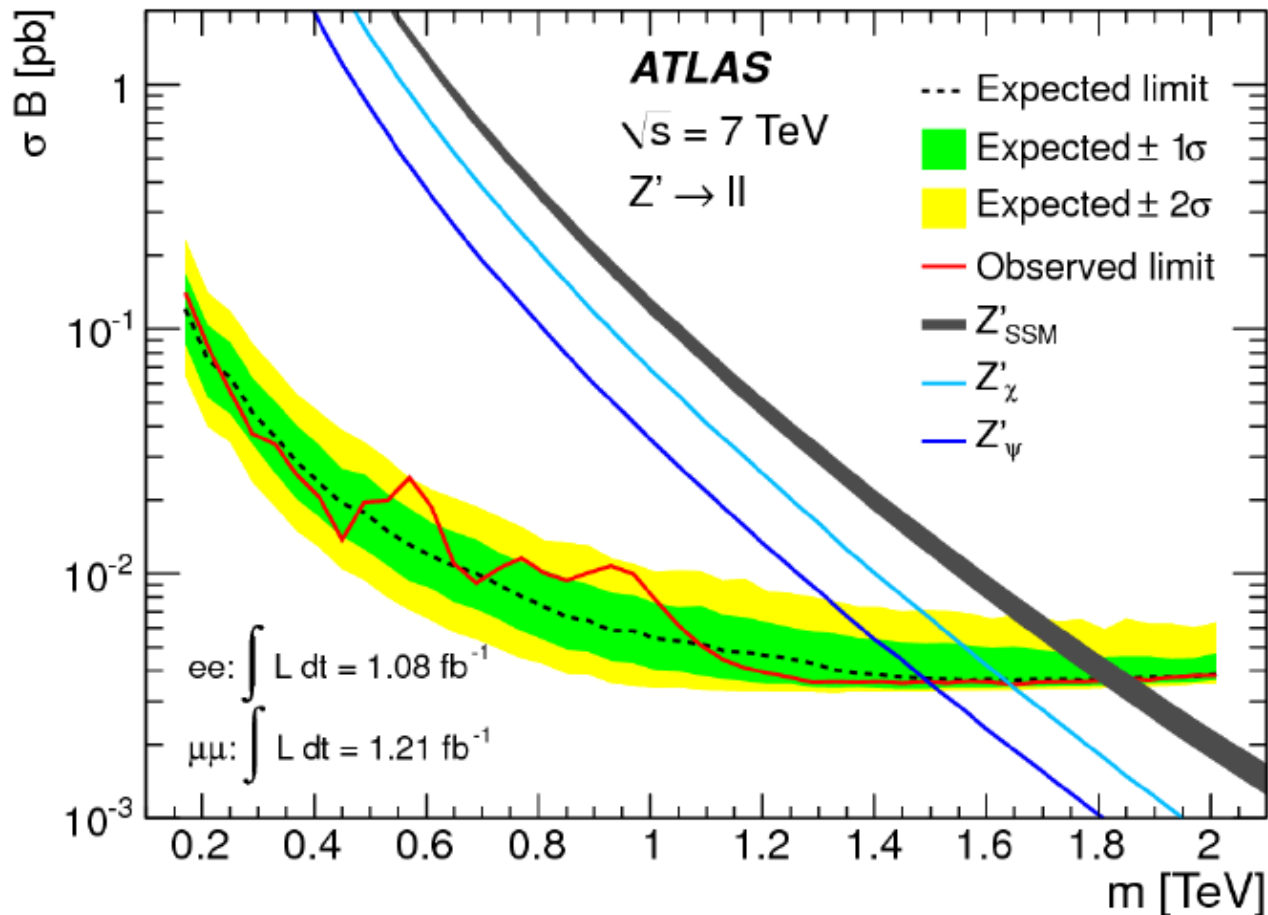
- ▶ TeV seesaw: $M_N \sim 10^{2:3} \text{ GeV}$.

$$W = y_\nu L H_2 N + \frac{1}{2} m_N N N$$

$$m_\nu = \frac{y_\nu^2 \langle H_2^0 \rangle^2}{m_N} \Rightarrow y_\nu \sim 10^{-6:7}$$

- ▶ DM: Right-Handed sNeutrino $\tilde{N}$
- ▶ LHC: $Z' / N, \tilde{Z}' / \tilde{N}$

LHC pushes up Z'



arXiv:1108.1582

EWPD and Tevatron

Z'	$M_{Z'} \text{ [GeV]}$				$\sin \theta_{ZZ'}$			$\chi^2_{\min}$
	electroweak	CDF	DØ	LEP 2	$\sin \theta_{ZZ'}$	$\sin \theta_{ZZ'}^{\min}$	$\sin \theta_{ZZ'}^{\max}$	
Z_χ	1,141	892	800	673	-0.0004	-0.0016	0.0006	47.3
Z_ψ	147	878	763	481	-0.0005	-0.0018	0.0009	46.5
Z_η	427	982	810	434	-0.0015	-0.0047	0.0021	47.7
Z_I	1,204	789	692		0.0003	-0.0005	0.0012	47.4
Z_S	1,257	821	719		-0.0003	-0.0013	0.0005	47.3
Z_N	623	861	744		-0.0004	-0.0015	0.0007	47.4
Z_R	442				-0.0003	-0.0015	0.0009	46.1
Z_{LR}	998	630		804	-0.0004	-0.0013	0.0006	47.3
Z_Ψ	(803)	(740)			-0.0015	-0.0094	0.0081	47.7
Z_{SM}	1,403	1,030	950	1,787	-0.0008	-0.0026	0.0006	47.2
Z_{string}	1,362				0.0002	-0.0005	0.0009	47.7
SM	∞				0			48.5

Langacker, 0801.1345

How to produce? Dirac neutrino (I)

- ▶ Nonthermal production.
- ▶ Thermal regeneration.

$$W = y_\nu L H_2 N$$

$$m_\nu = y_\nu \langle H_2^0 \rangle \Rightarrow y_\nu \sim 10^{-13}$$

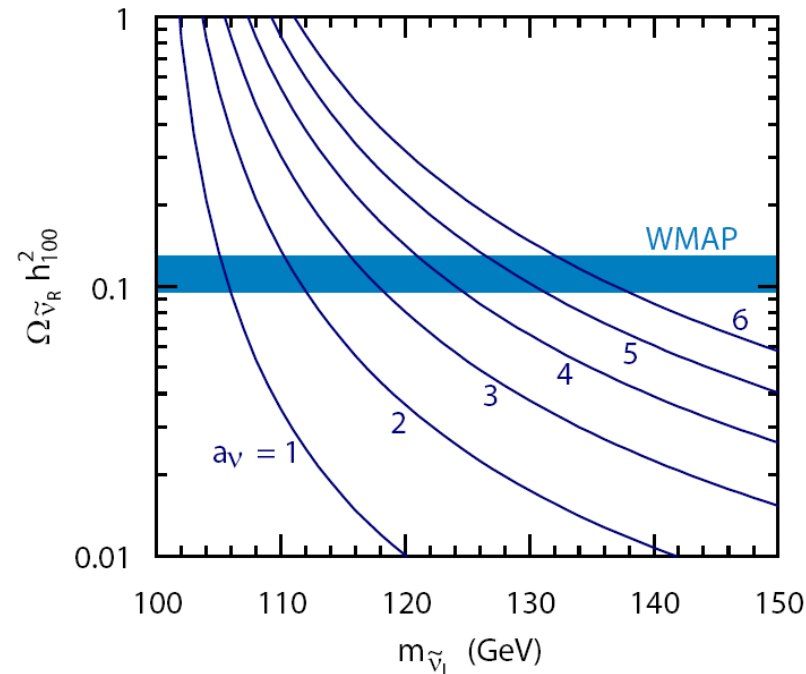
$$\mathcal{L}_A = A_\nu H_u \tilde{L} \tilde{\nu}_R^c + \text{h.c.}$$

$$A_\nu = a_\nu y_\nu m_{\tilde{L}}$$

$$\Gamma_{\tilde{W}^0 \rightarrow \tilde{\nu}_R \bar{\nu}_L} = \frac{\beta_f^2 g_2^2}{64\pi} \left[\frac{A_\nu v_T}{m_{\tilde{\nu}_L}^2 - m_{\tilde{\nu}_R}^2} \right]^2 m_{\tilde{W}}$$

$$m_{\tilde{\nu}_R} = 100 \text{ GeV}, m_{\tilde{W}} = 300 \text{ GeV}, \text{ and } \mu_H = 150 \text{ GeV}$$

Asaka, et.al., 0512118

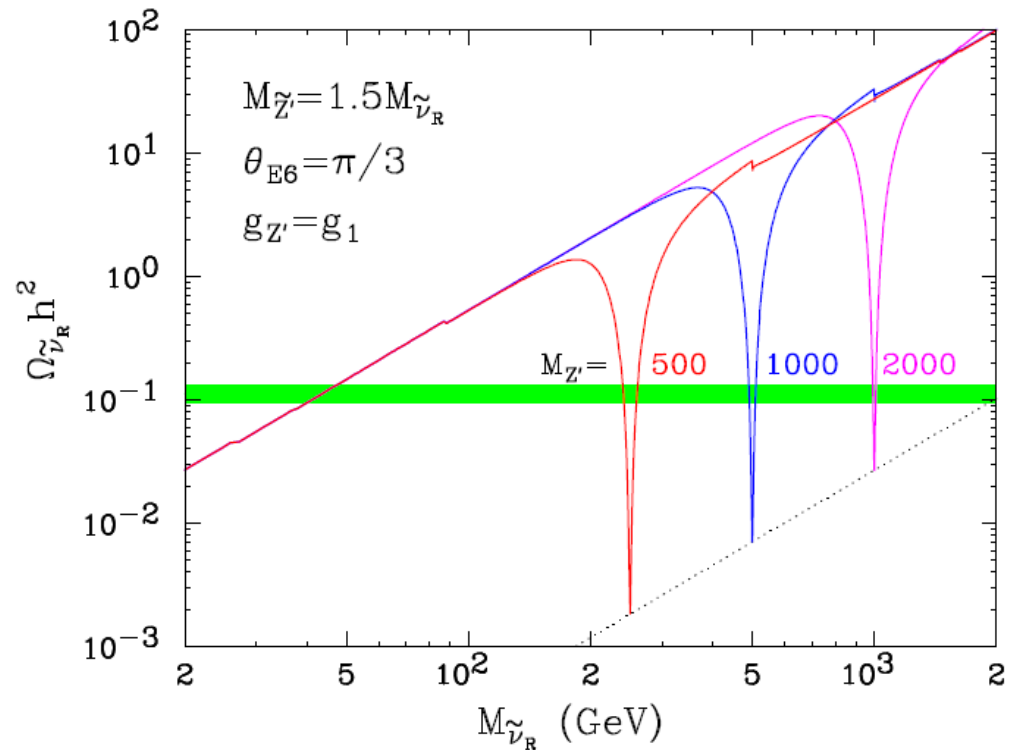
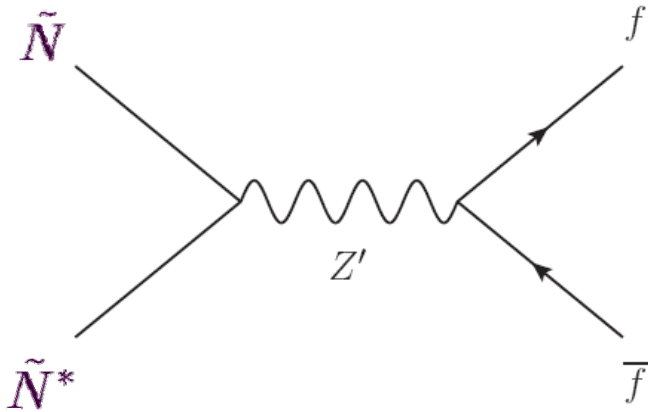


How to produce? Dirac neutrino (II)

► U(1)': $Q' = Q_\chi \cos \theta_{E6} + Q_\psi \sin \theta_{E6}$

Lee, et.al., 0702223

$$\tilde{N}\tilde{N}^* \rightarrow Z' \rightarrow f\bar{f}$$



How to produce? Majorana neutrino (I)

- ▶ Too huge thermal regeneration:

$$y_\nu \sim 10^{-7} \rightarrow \Omega_{\text{DM}} \gg 1.$$

- ▶ Need thermalization: large Yukawa or gauge.

- ▶ Large A term:

Arkani-Hamed, et.al., 0006312

Borzumati-Nomura, 0007018

$$V_{\text{soft}} = A_\nu \tilde{l} H_u \tilde{N} \text{ with } A_\nu \sim 100 \text{ GeV}$$

$$\text{Usually, } A_\nu \sim y_\nu m_{3/2}$$

How to produce? Majorana neutrino (II)

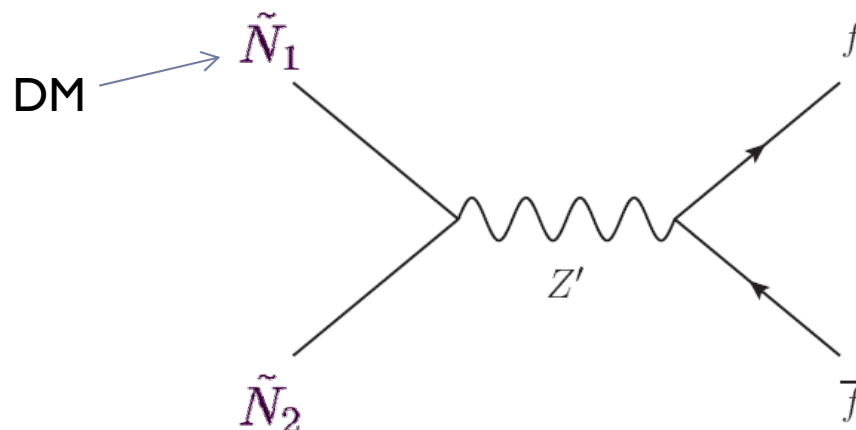
► Z'?

► RHsN mass splitting: $\tilde{N} = \tilde{N}_1 + i\tilde{N}_2$

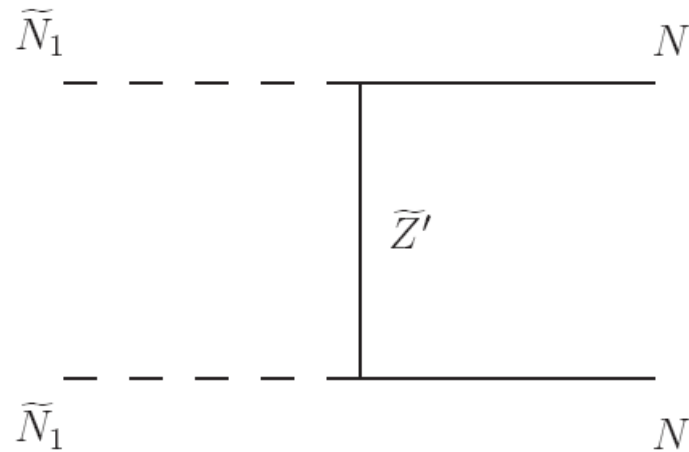
$$V_{mass} = (m_N^2 + m_{\tilde{N}}^2 - \frac{1}{4}m_{Z'}^2 c_{2\beta'}) |\tilde{N}|^2 - \frac{1}{2}B_N m_N (\tilde{N}\tilde{N} + \tilde{N}^*\tilde{N}^*)$$

$$m_{\tilde{m}_{1,2}}^2 = m_N^2 + m_{\tilde{N}}^2 - \frac{1}{4}m_{Z'}^2 \mp B_N m_N.$$

► Z': too heavy, off-diagonal coupling.



Can $U(1)'$ gaugino help?



► Yes! How?

$U(1)_X$ model

- ▶ New fields and anomaly-free charge assignment:

$SU(5)$	10_F	$\bar{5}_F$	$1(N)$	5_H	$\bar{5}_H$	$1(X)$	$1(S_1)$	$1(S_2)$
$2\sqrt{10}Q'$	-1	3	-5	2	-2	0	10	-10

- ▶ Gauge invariant Yukawas:

$$W_{seesaw} = y_{ij} L_i H_u N_j + \frac{\lambda_{N_i}}{2} S_1 N_i N_i$$

$$W' = \mu' S_1 S_2 \quad \text{or} \quad W' = \lambda X S_1 S_2 + \frac{\kappa}{3} X^3$$

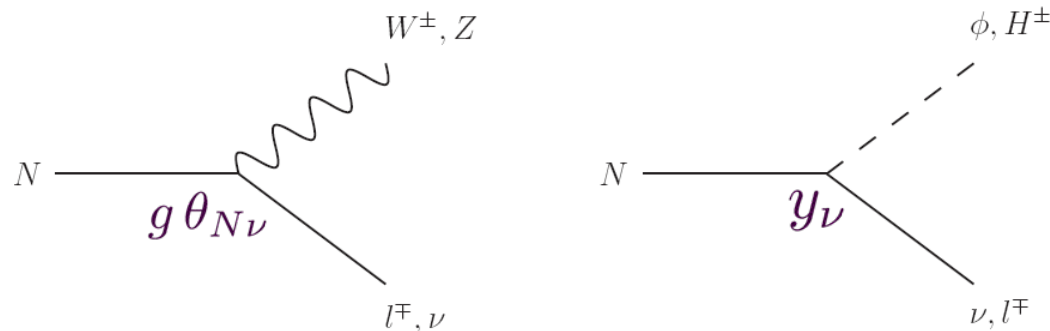
Left-Right Mixing

- ▶ L-R mixing (Dirac) mass leads to a small mixing between LH neutrino/sneutrino and RH neutrino/sneutrino:

$$\mathcal{L}_{mixing} = m_D \nu N + \tilde{m} m_D \tilde{\nu} \tilde{N}$$

$$\theta_{N\nu} \approx \frac{y_\nu v_u}{\sqrt{2} m_N} \quad \theta_{\tilde{N}_{1,2} \tilde{\nu}_{R,I}} = y_\nu \frac{(m_N \pm A_L \mp \mu \cot \beta) v_u}{\sqrt{2} (m_{\tilde{N}_{1,2}}^2 - m_{\tilde{\nu}_{R,I}}^2)}$$

- ▶ RH neutrino/sneutrino couples not only to Higgs bosons but also to Z/W bosons through mixing:



RHN decays

$$\Gamma(N \rightarrow \nu h) = \Gamma(N \rightarrow \bar{\nu} h) = \frac{1}{8} \frac{y_\nu^2 m_N}{8\pi} \left(1 - \frac{m_h^2}{m_N^2}\right)^2 \cos^2 \alpha ,$$

$$\Gamma(N \rightarrow \nu H) = \Gamma(N \rightarrow \bar{\nu} H) = \frac{1}{8} \frac{y_\nu^2 m_N}{8\pi} \left(1 - \frac{m_H^2}{m_N^2}\right)^2 \sin^2 \alpha ,$$

$$\Gamma(N \rightarrow \nu A) = \Gamma(N \rightarrow \bar{\nu} A) = \frac{1}{8} \frac{y_\nu^2 m_N}{8\pi} \left(1 - \frac{m_A^2}{m_N^2}\right)^2 \cos^2 \beta ,$$

$$\Gamma(N \rightarrow l H^+) = \Gamma(N \rightarrow \bar{l} H^-) = \frac{1}{4} \frac{y_\nu^2 m_N}{8\pi} \left(1 - \frac{m_{H^\pm}^2}{m_N^2}\right)^2 \cos^2 \beta ,$$

$$\Gamma(N \rightarrow \nu Z) = \Gamma(N \rightarrow \bar{\nu} Z) = \frac{1}{8} \frac{y_\nu^2 m_N}{8\pi} \left(1 - \frac{M_Z^2}{m_N^2}\right)^2 \left(1 + 2 \frac{M_Z^2}{m_N^2}\right) \sin^2 \beta ,$$

$$\Gamma(N \rightarrow l W^+) = \Gamma(N \rightarrow \bar{l} W^-) = \frac{1}{8} \frac{y_\nu^2 m_N}{8\pi} \left(1 - \frac{M_W^2}{m_N^2}\right)^2 \left(1 + 2 \frac{M_W^2}{m_N^2}\right) \sin^2 \beta .$$

U(1)' gaugino $\tilde{Z}'$

- ▶ U(1)' gaugino-Higgsino mass matrix:

$$(\tilde{Z}', \tilde{S}'_1, \tilde{S}'_2)$$

$$\mathcal{M} = \begin{bmatrix} m_M & M_{Z'} c_{\beta'} & -M_{Z'} s_{\beta'} \\ M_{Z'} c_{\beta'} & 0 & \mu' \\ -M_{Z'} s_{\beta'} & \mu' & 0 \end{bmatrix}$$

- ▶ Gaugino lighter than Z' : $m_M, M_{Z'} \ll \mu'$

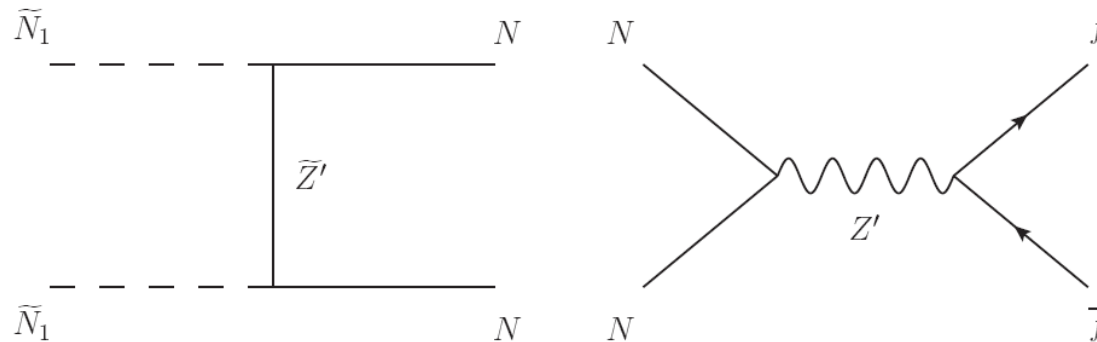
$$m_{\tilde{Z}'} \approx m_M - \frac{M_{Z'}^2}{\mu'} \ll 1\text{TeV}$$

U(1)' gauge interactions

$\bar{f} f Z'$	$ig'(Q'_{f_L} P_L - Q'_{f_R^\dagger} P_R) \gamma^\mu$	$h A Z'$	$-ig' Q'_{H_u} \sin(\alpha + \beta) q^\mu$
$\tilde{f} \tilde{f}^* Z'$	$ig' Q'_f q^\mu$	$H A Z'$	$ig' Q'_{H_u} \cos(\alpha + \beta) q^\mu$
$\tilde{f}_L \bar{f} \tilde{Z}'$	$-i\sqrt{2} g' Q'_{f_L} P_R$	$H^+ H^- Z'$	$ig' Q'_{H_u} q^\mu$
$\tilde{f}_R \bar{f} \tilde{Z}'$	$-i\sqrt{2} g' Q'_{f_R^\dagger} P_L$	$h Z Z'$	$-i\sqrt{2} g' M_Z Q'_{H_u} \cos(\alpha + \beta) g^{\mu\nu}$
$\tilde{N}_1 \tilde{N}_2 Z'$	$ig' Q'_N q^\mu$	$H Z Z'$	$-i\sqrt{2} g' M_Z Q'_{H_u} \sin(\alpha + \beta) g^{\mu\nu}$
$\bar{N} \tilde{N}_1 \tilde{Z}'$	$-ig' Q'_N P_R$	$h \bar{\chi}_i^0 \tilde{Z}'$	$-ig' Q'_{H_u} (N_{4i} \cos \alpha + N_{3i} \sin \alpha) P_R$
$\bar{N} \tilde{N}_2 \tilde{Z}'$	$g' Q'_N P_R$	$H \bar{\chi}_i^0 \tilde{Z}'$	$-ig' Q'_{H_u} (N_{4i} \sin \alpha - N_{3i} \cos \alpha) P_R$
$S_{1,2} S_{1,2}^* Z'$	$ig' Q'_{S_{1,2}} q^\mu$	$A \bar{\chi}_i^0 \tilde{Z}'$	$g' Q'_{H_u} (N_{4i} \cos \beta - N_{3i} \sin \beta) P_R$
$\bar{\tilde{S}}_{1,2} \tilde{S}_{1,2} Z'$	$ig' Q'_{S_{1,2}} \gamma^\mu$	$H^+ \bar{\chi}_i^+ \tilde{Z}'$	$-i\sqrt{2} g' Q'_{H_u} (V_{2i}^* \cos \beta P_R - U_{2i} \sin \beta P_L)$
$S_{1,2} \bar{\tilde{S}}_{1,2} \tilde{Z}'$	$-i\sqrt{2} g' Q'_{S_{1,2}} P_R$		

Thermal abundance of RHsN (I)

- ▶ RHsN annihilation to RHN; RHN to the usual thermal fermions:

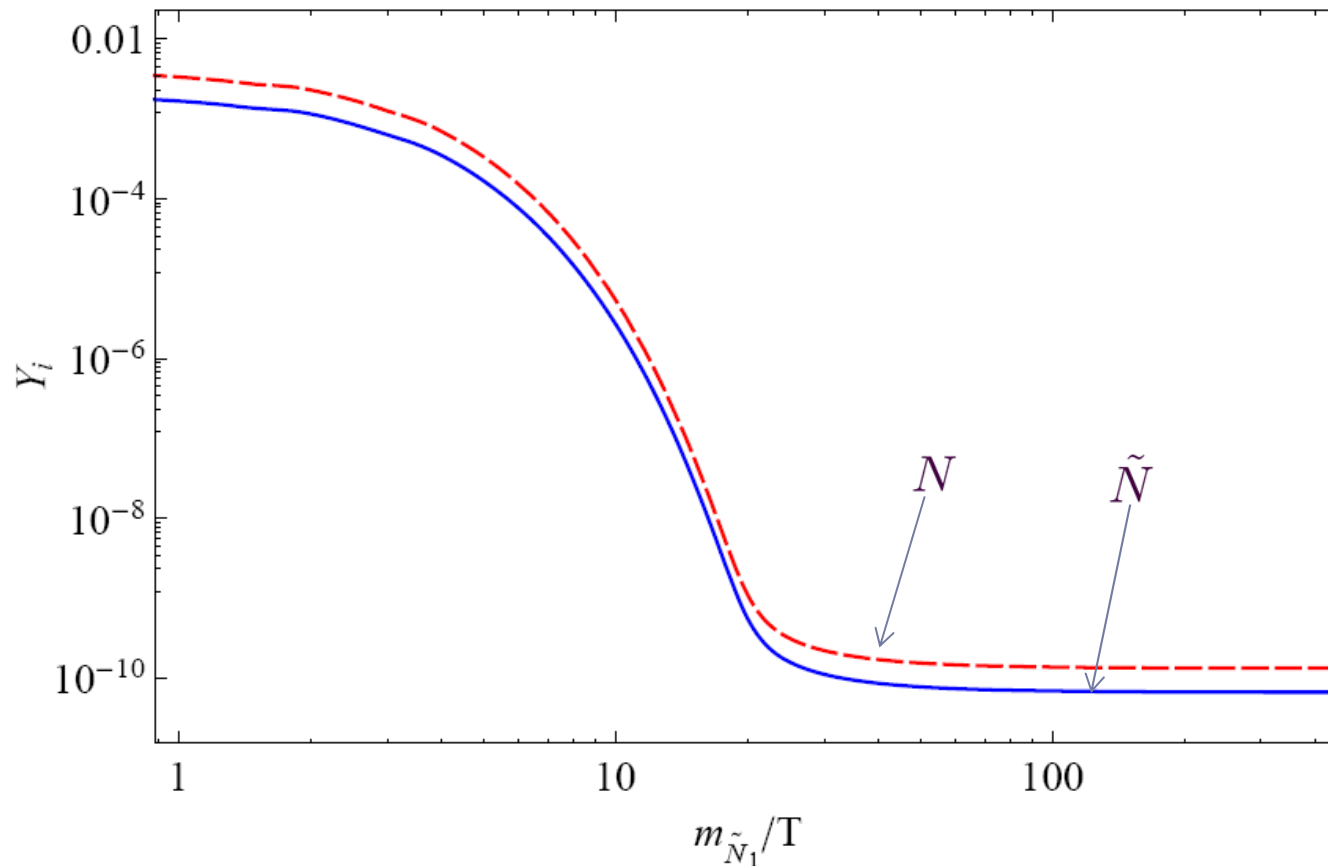


- ▶ Boltzmann equations to solve:

$$\frac{dn_{\tilde{N}_1}}{dt} = -3Hn_{\tilde{N}_1} - \langle \sigma_{\tilde{N}_1} v_{\tilde{N}_1} \rangle \left[(n_{\tilde{N}_1})^2 - \left(\frac{g_{\tilde{N}_1}}{g_N} n_N \right)^2 \right]$$

$$\frac{dn_N}{dt} = -3Hn_N - \langle \sigma_N v_N \rangle [(n_N)^2 - (n_N^{\text{eq}})^2] + \langle \sigma_{\tilde{N}_1} v_{\tilde{N}_1} \rangle \left[(n_{\tilde{N}_1})^2 - \left(\frac{g_{\tilde{N}_1}}{g_N} n_N \right)^2 \right]$$

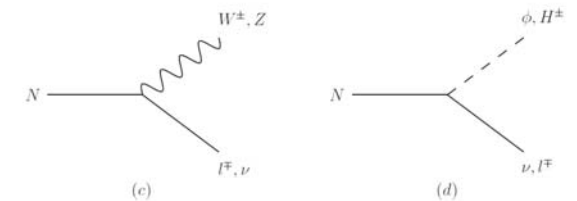
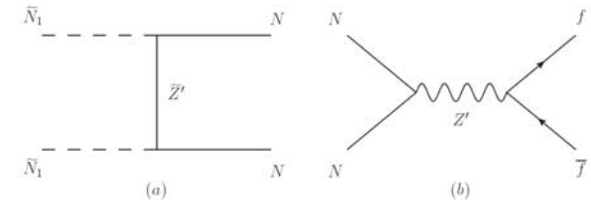
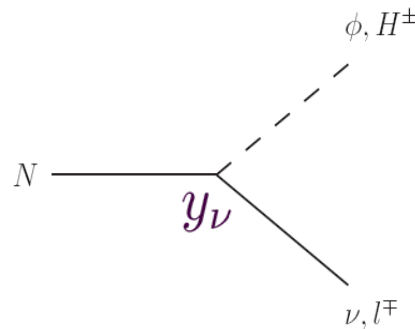
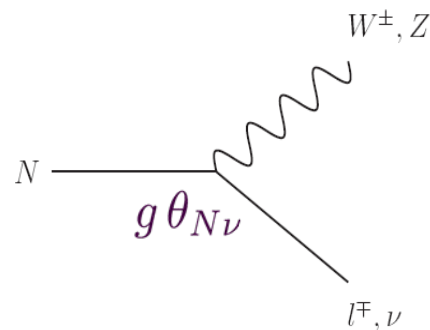
Solution (I): too large!



$$m_{\tilde{Z}'} = 600\text{GeV}, m_{Z'} = 1200\text{GeV}$$

Thermal abundance of RHsN (II)

- Include the decay and inverse-decay of RHN:



- Boltzmann equations:

$$\frac{dn_{\tilde{N}_1}}{dt} = -3Hn_{\tilde{N}_1} - \langle \sigma_{\tilde{N}_1} v_{\tilde{N}_1} \rangle \left[(n_{\tilde{N}_1})^2 - \left(\frac{g_{\tilde{N}_1}}{g_N} n_N \right)^2 \right]$$

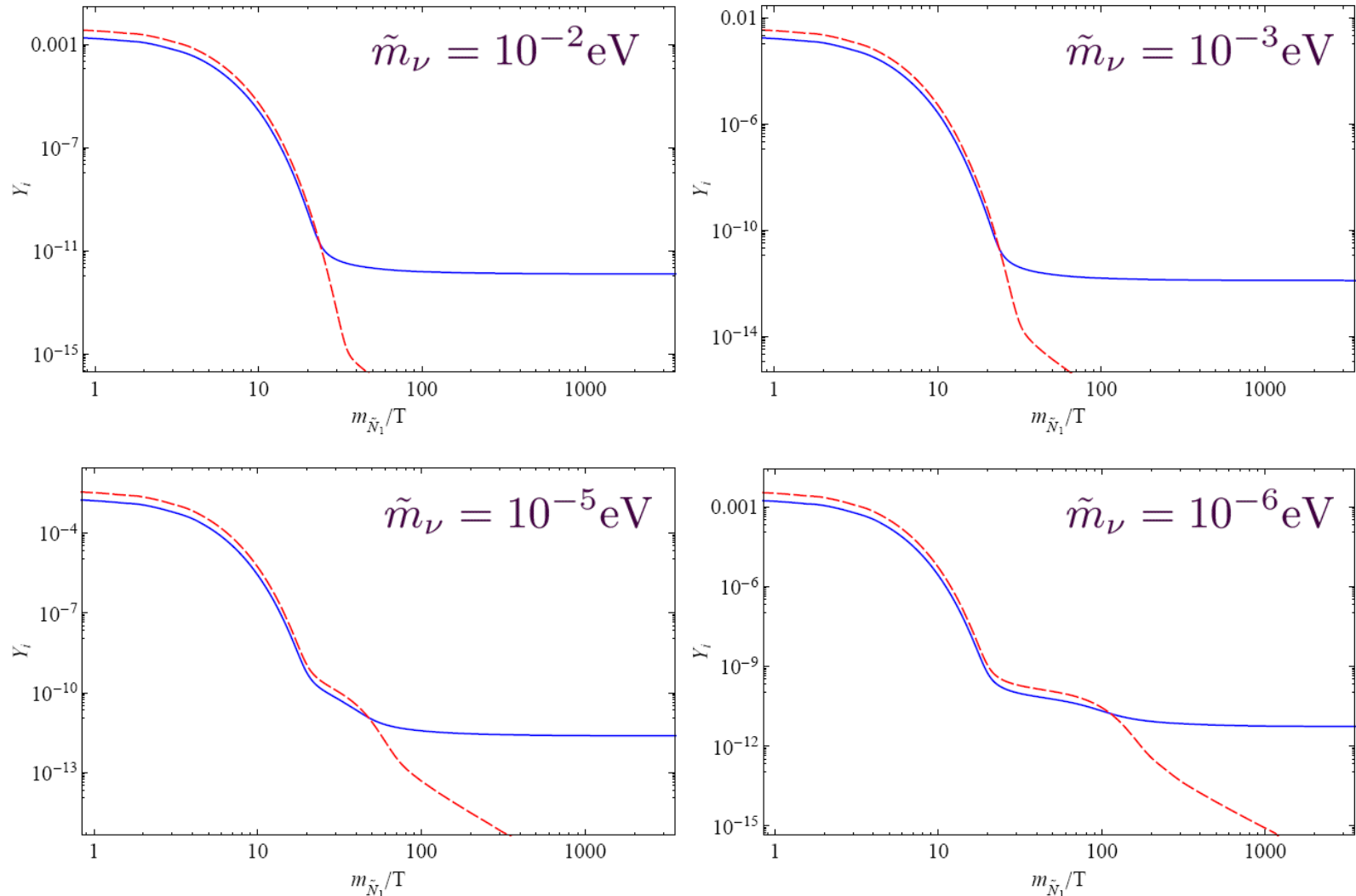
$$\begin{aligned} \frac{dn_N}{dt} = & -3Hn_N - \langle \sigma_N v_N \rangle [(n_N)^2 - (n_N^{\text{eq}})^2] + \langle \sigma_{\tilde{N}_1} v_{\tilde{N}_1} \rangle \left[(n_{\tilde{N}_1})^2 - \left(\frac{g_{\tilde{N}_1}}{g_N} n_N \right)^2 \right] \\ & - \Gamma_N (n_N - n_N^{\text{eq}}) , \end{aligned}$$

Solution (II)

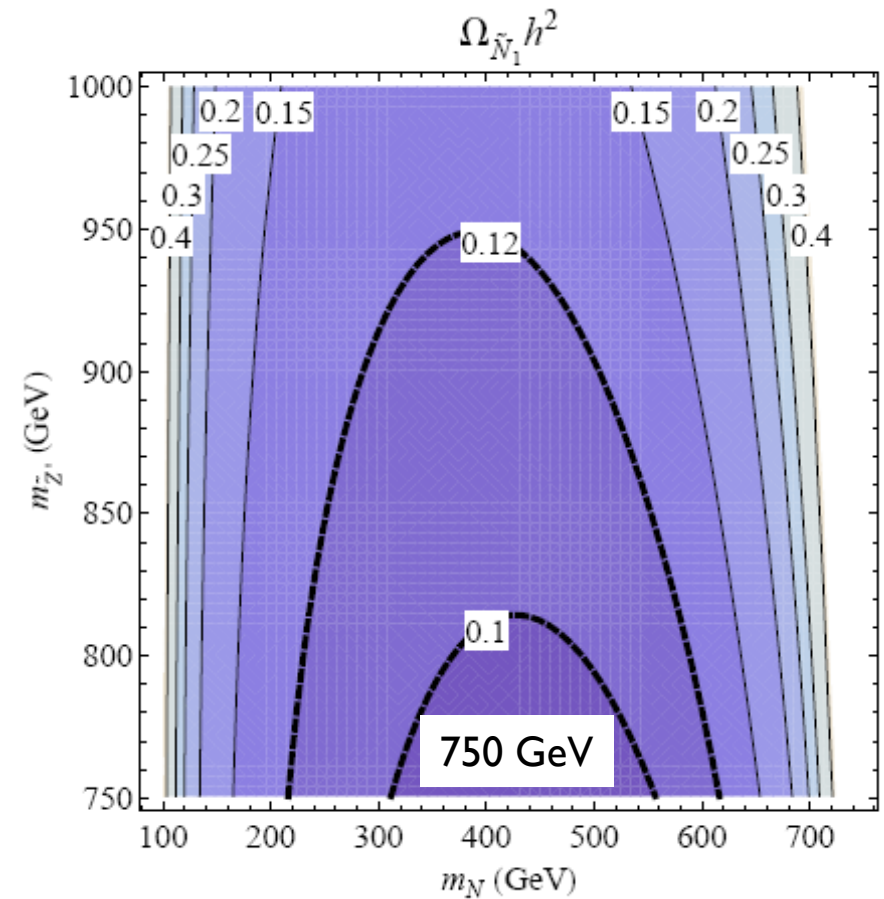
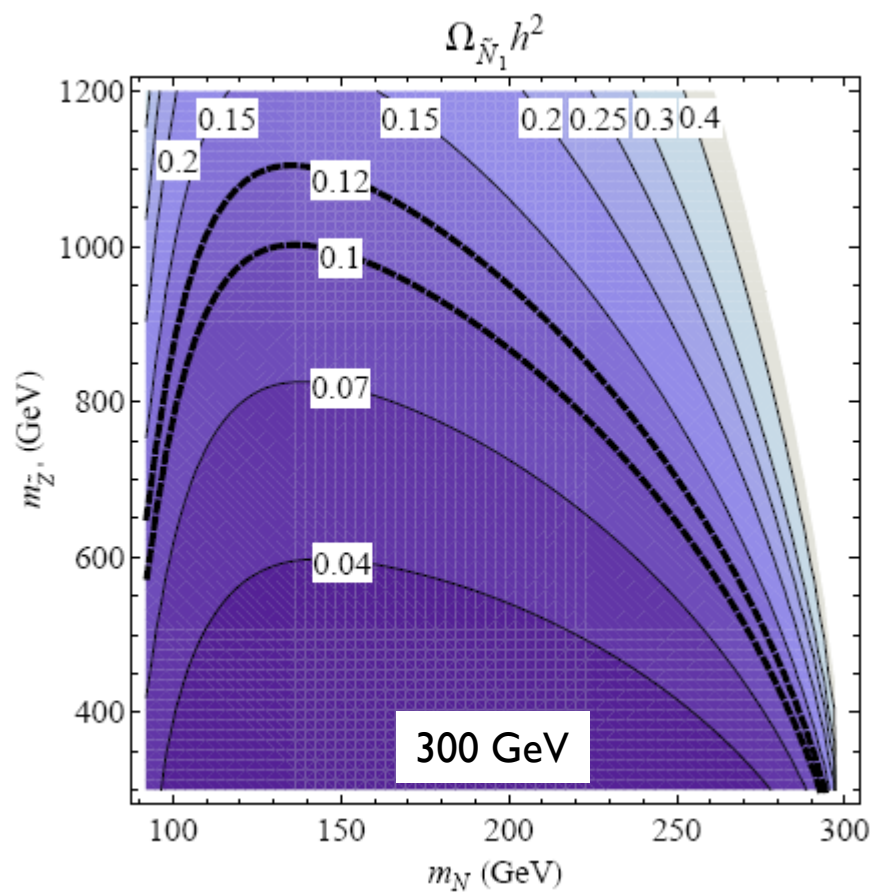
- ▶ Although suppressed by small Yukawas $y_\nu \sim 10^{-6}$, the decay and inverse-decay processes maintain in thermal equilibrium longer the RHN and thus also the RHsN.
- ▶ The RHsN DM relic density can be in the right range unless y_ν is small.
- ▶ To see a peculiar dependence on y_ν , define an effective neutrino mass:

$$\tilde{m}_\nu \equiv \frac{|y_\nu|^2 v^2}{M}$$

Decay/inverse-decay effect

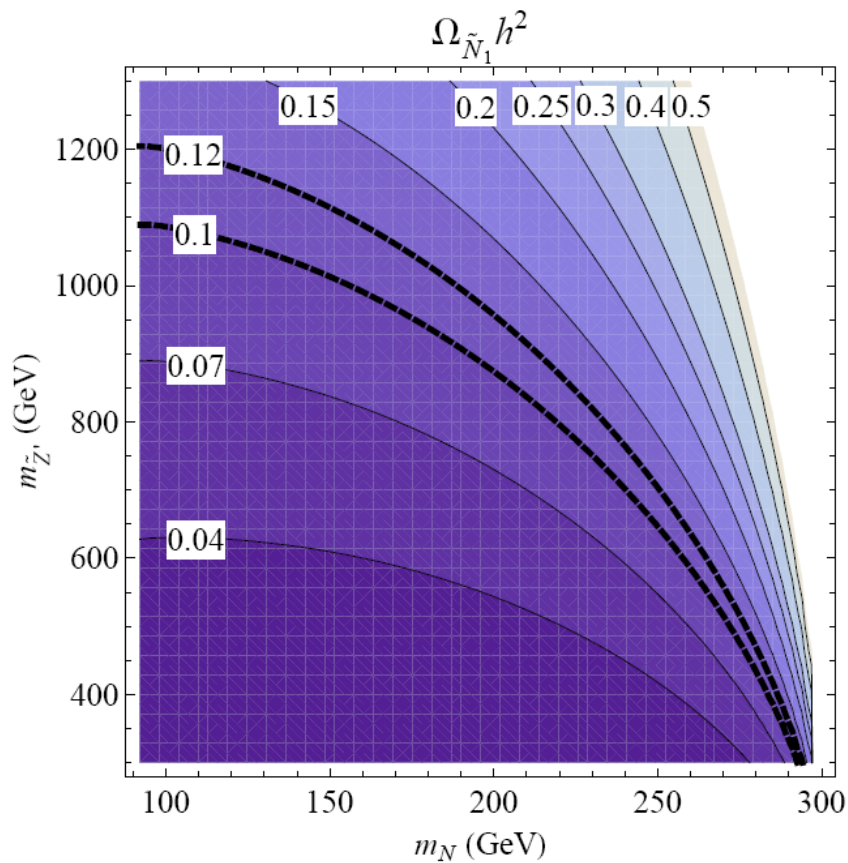


DM mass dependence $\tilde{m}_\nu = 10^{-3}\text{eV}, m_{Z'} = 1.2\text{TeV}$

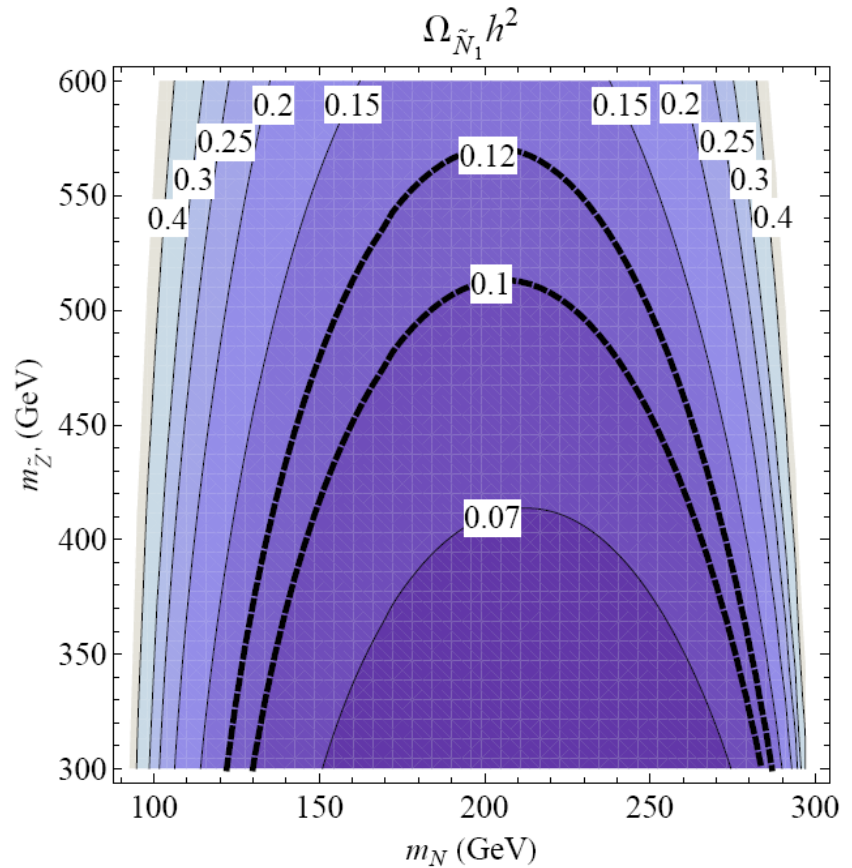


Dependency on y_ν

$$m_{\tilde{N}_1} = 300\text{GeV}, m_{Z'} = 1.2\text{TeV}$$

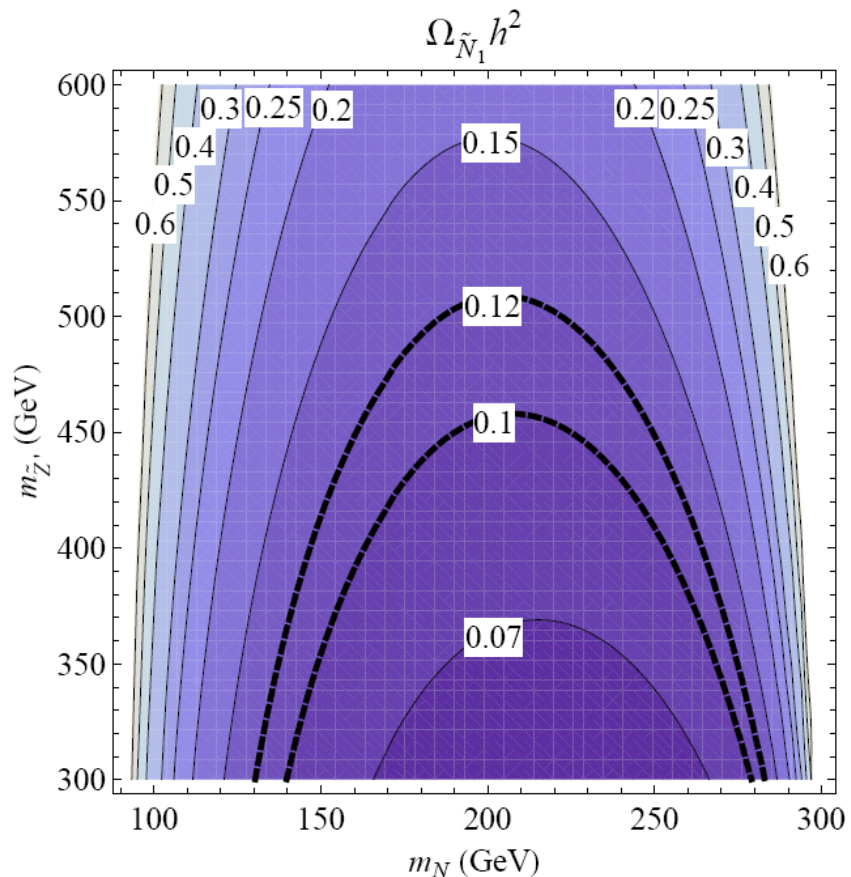


$$\tilde{m}_\nu = 10^{-1}\text{eV}$$

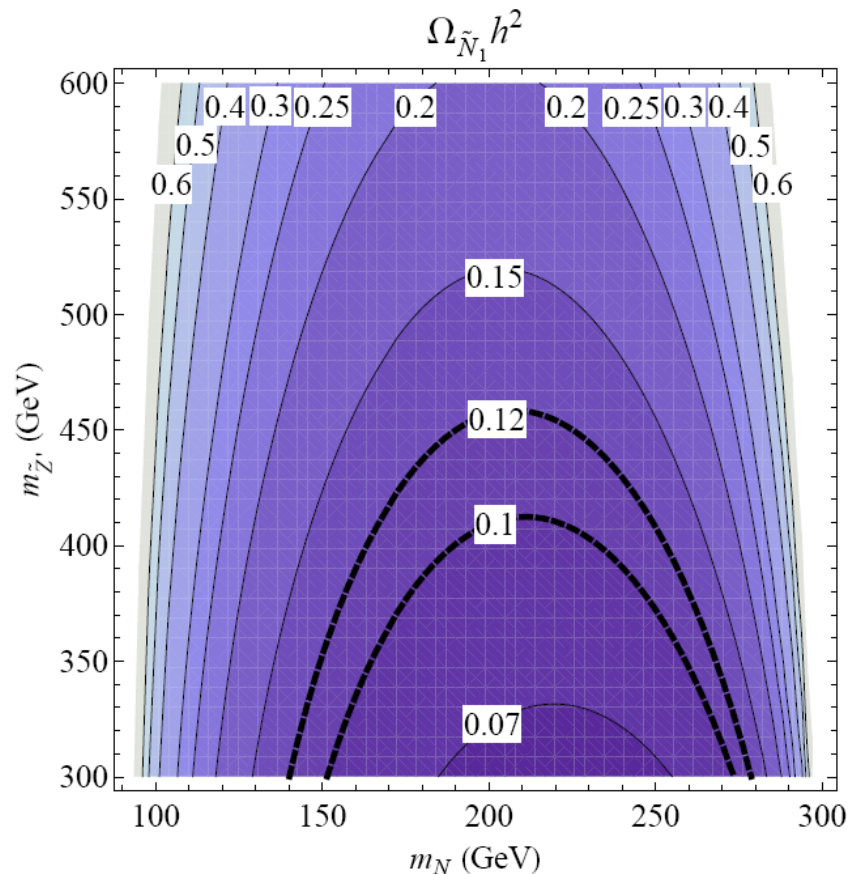


$$\tilde{m}_\nu = 10^{-5}\text{eV}$$

Dependency on Z' mass $m_{\tilde{N}_1} = 300\text{GeV}$, $\tilde{m}_\nu = 10^{-3}\text{eV}$



$m_{Z'} = 2\text{TeV}$

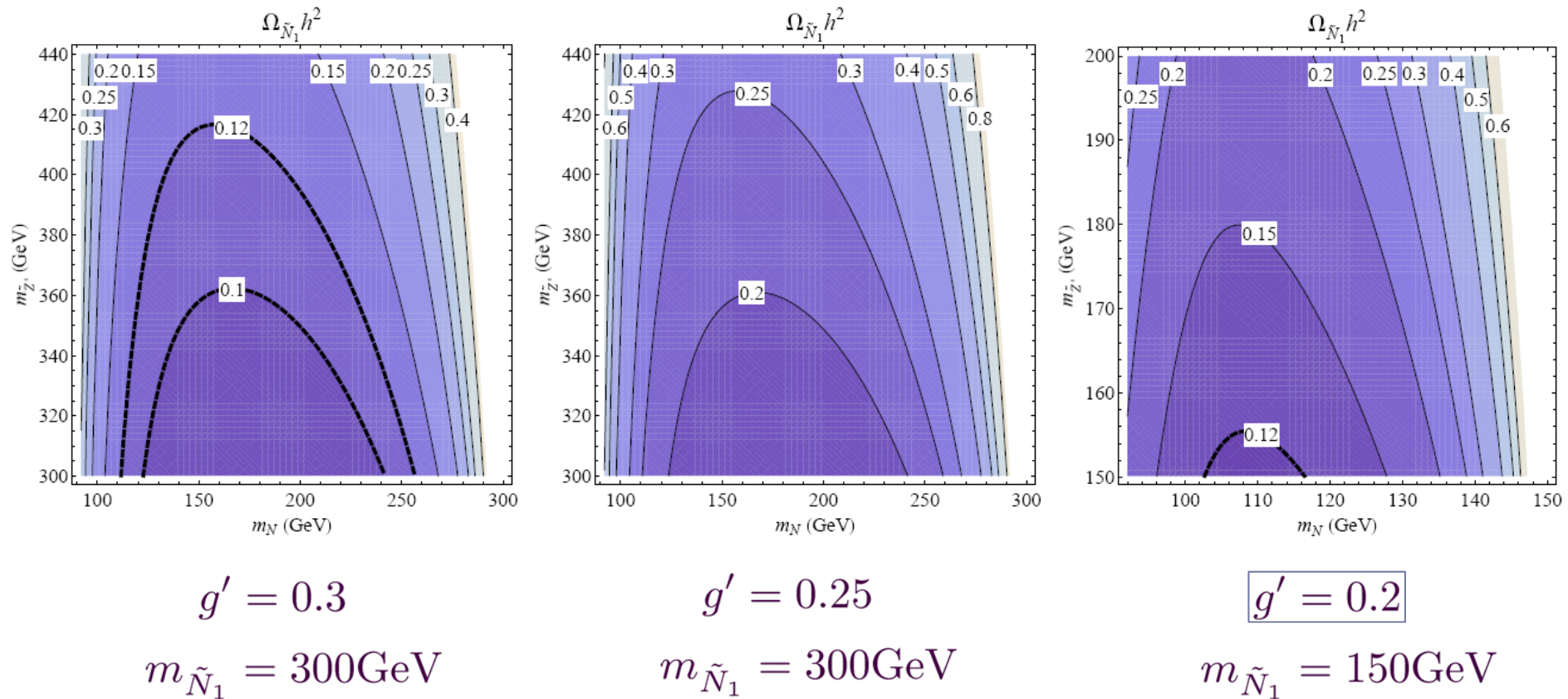


$m_{Z'} = 4\text{TeV}$

Dependency on g'

$$m_{Z'} = 1.2\text{TeV}, \tilde{m}_\nu = 10^{-3}\text{eV}$$

$$g'_{\text{standard}} = \sqrt{5/3}g_2 \tan \theta_W = 0.46$$



LHC signatures

- ▶ Discover RHN: **SSD** and/or **displaced vertices**

$$pp \rightarrow Z' \rightarrow NN \rightarrow l^\pm l^\pm W^\mp W^\mp (4j)$$

$$N \rightarrow lW, \nu Z, \nu h, \nu H, \nu A, lH^+$$

Probe Majorana nature of neutrino

- ▶ Similar with large **missing** E_T :

$$pp \rightarrow \tilde{Z}' \tilde{Z}' \rightarrow NN\tilde{N}_1\tilde{N}_1 \rightarrow l^\pm l^\pm 4j + \cancel{E}_T$$

- ▶ Probe the **Higgs** channel(s) (with MET)?

$$\begin{aligned} pp \rightarrow Z'(\tilde{Z}'\tilde{Z}') &\rightarrow l^\pm W^\mp \nu h (+\cancel{E}_T) \\ &\rightarrow l^\pm W^\mp l^\pm H^\mp (+\cancel{E}_T) \end{aligned}$$

Discovery channels

$$\tilde{q} \rightarrow q \tilde{Z}_{B-L}$$

Dijet, same-sign dilepton, MET
Dijet, displaced b-jet(s), MET

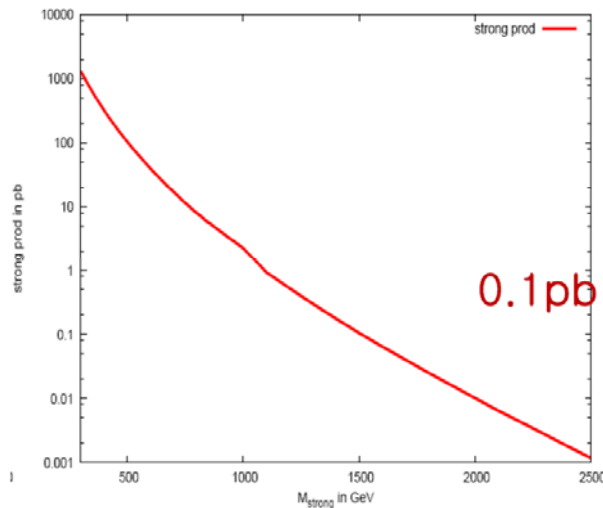
$$N \tilde{N}$$

$$\nu h, l^{\pm} W^{\mp}, \nu Z$$

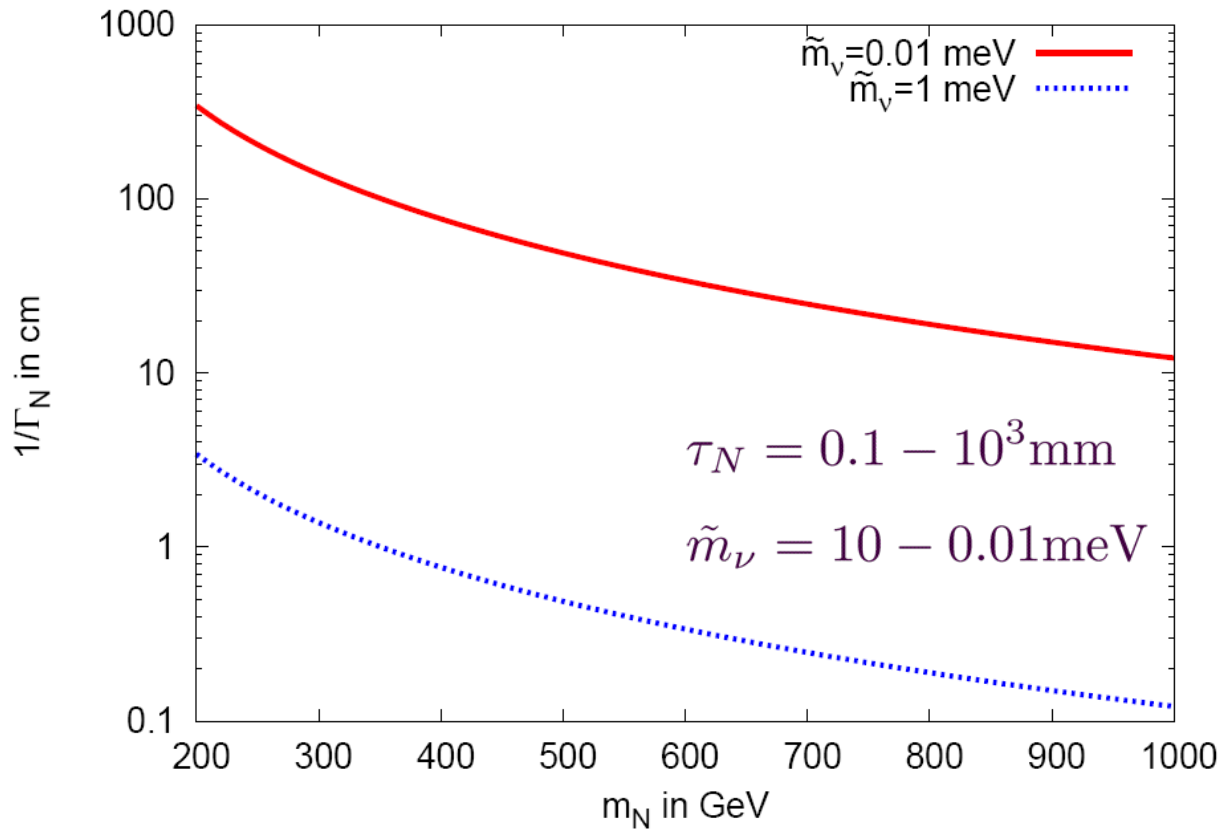
0.1 pb for $m_{\text{squark}} = 1.5 \text{ TeV}$

$$b \bar{b}$$

$\tau = 0.2 \text{ cm}$
for $m_{\nu} = 10^{-3} \text{ eV}$ and $m_N = 200 \text{ GeV}$

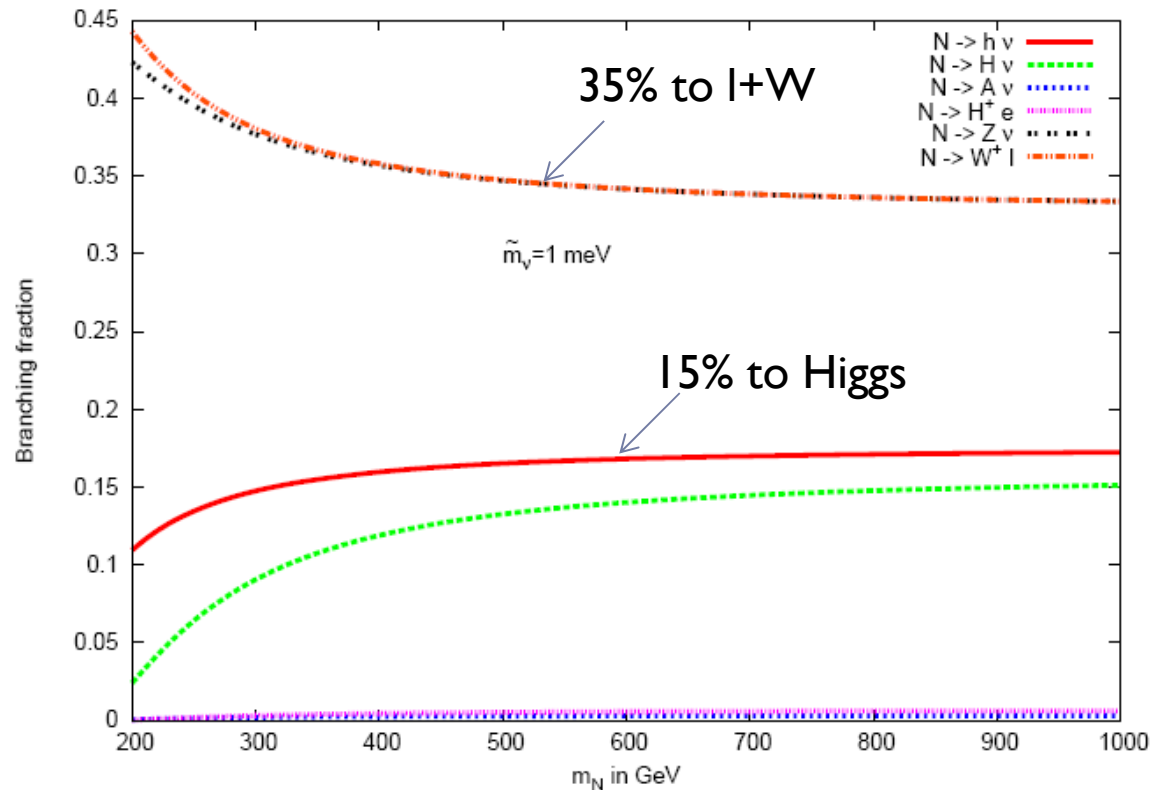


Displaced vertex



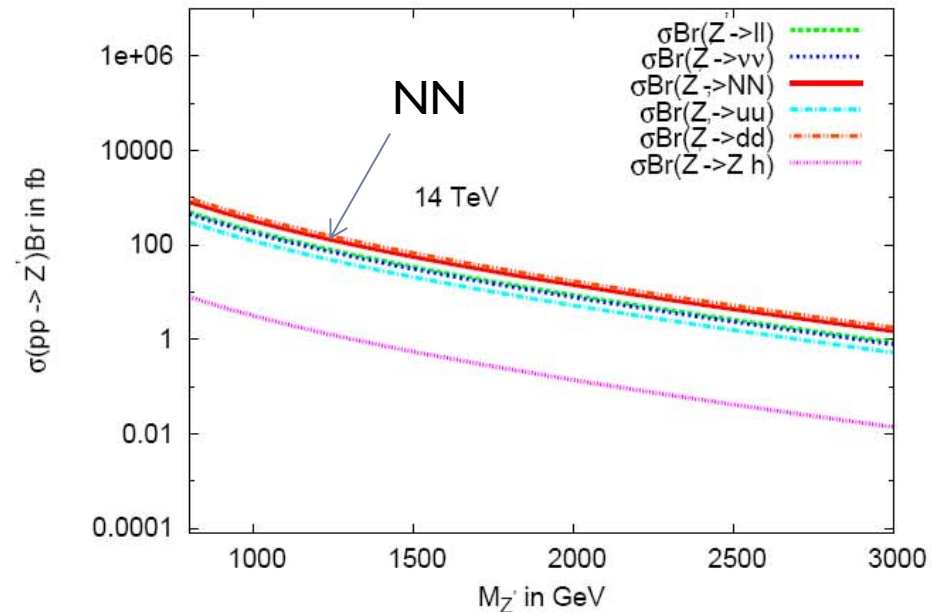
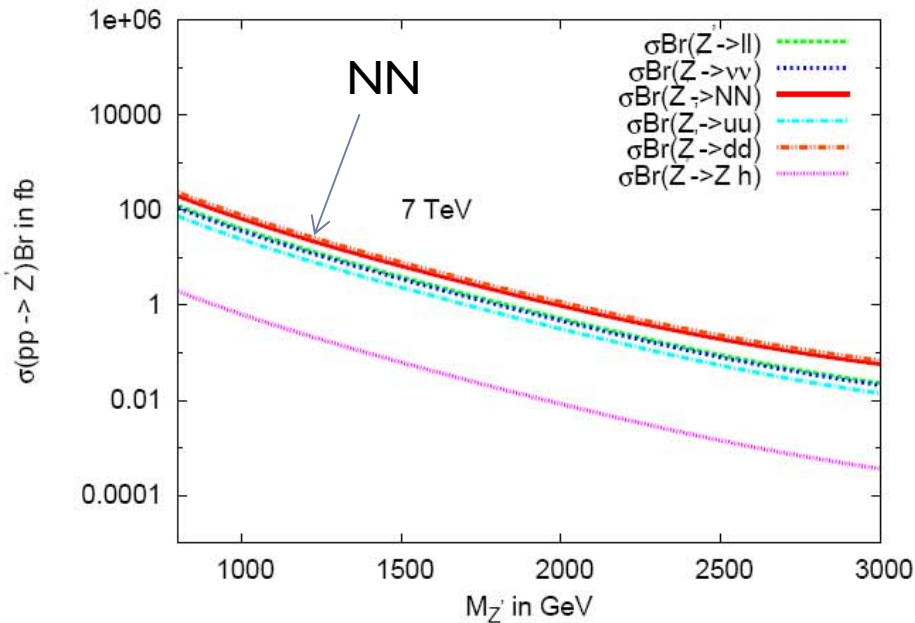
$$\Gamma_N \sim y_\nu^2 m_N \sim \tilde{m}_\nu m_N^2 / v$$

N decay branching fraction



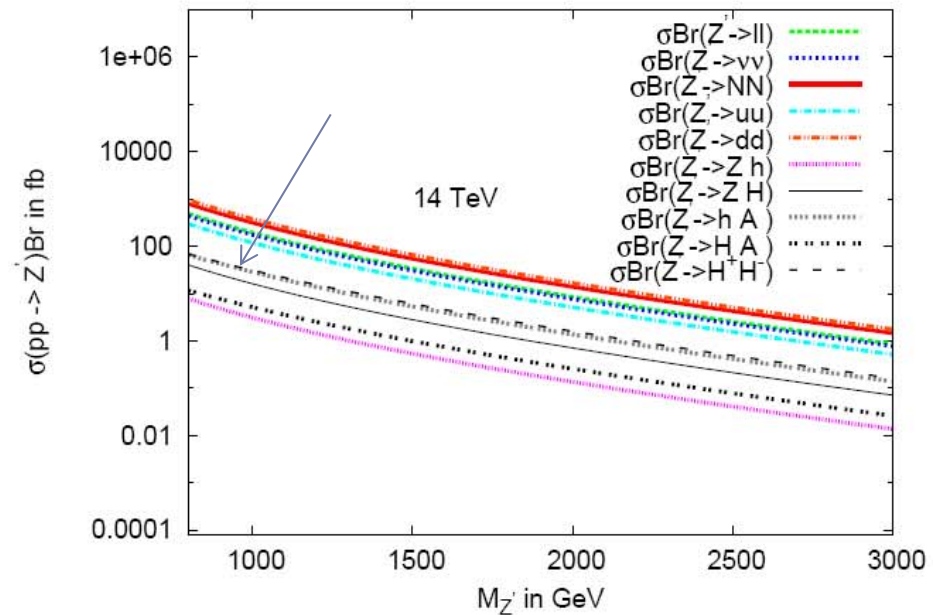
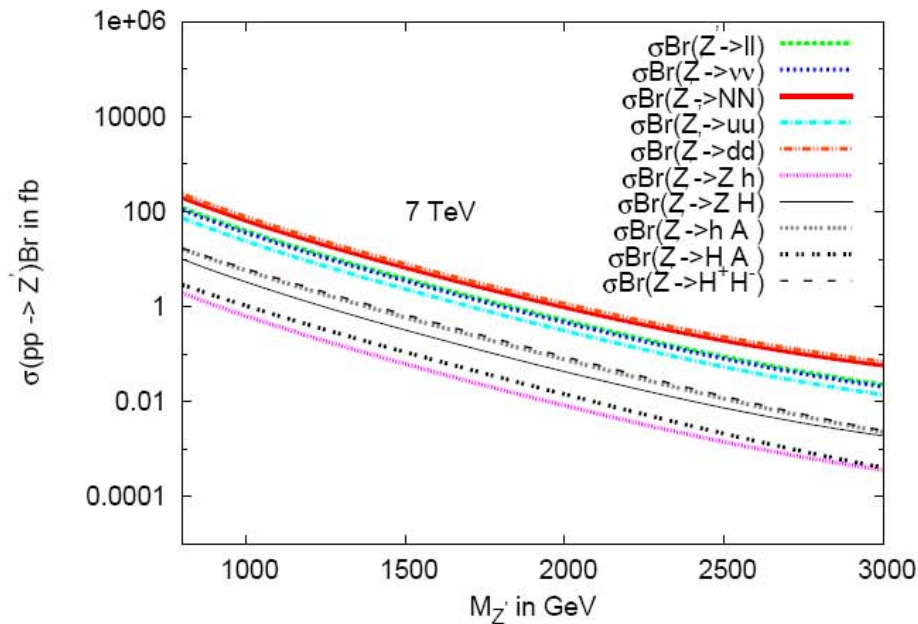
Z' production and decay

Heavy Higgs decoupled



Z' production and decay at LHC

Heavy Higgses non-decoupled



$$m_A = 162\text{GeV}, m_{H^\pm} = 180\text{GeV}, m_H = 164\text{GeV}$$

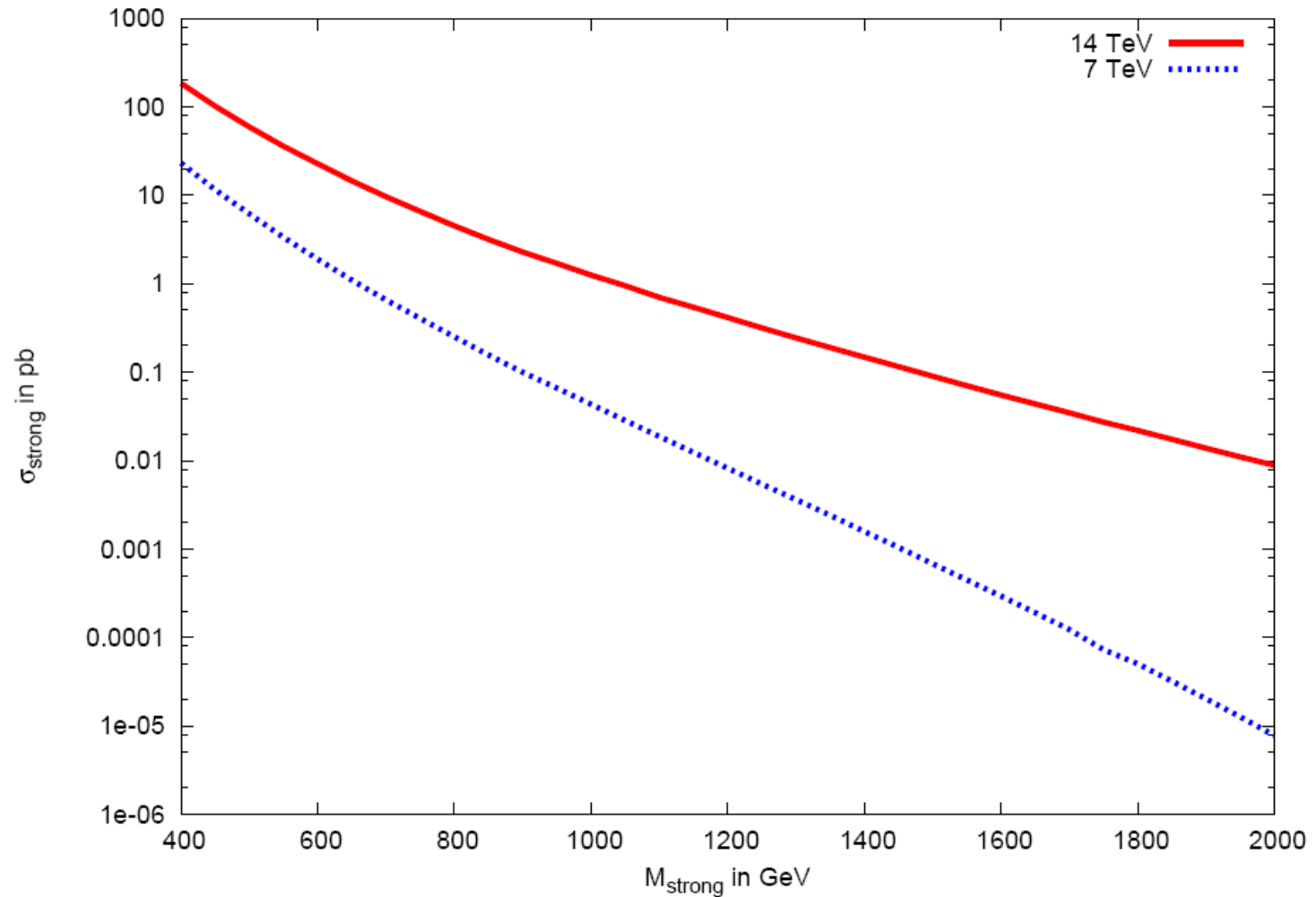
U(1)' gaugino production

- ▶ Cascade decay from gluinos/squarks:

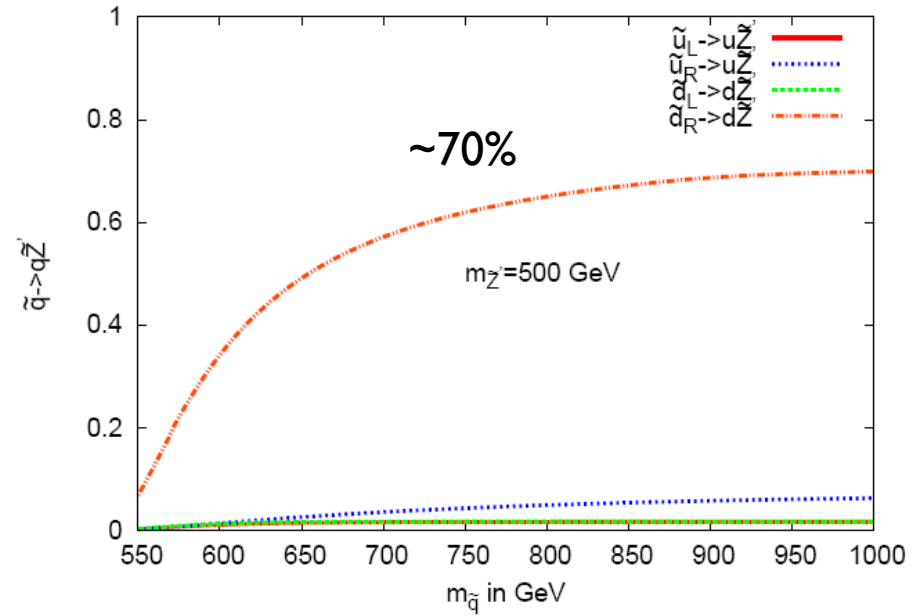
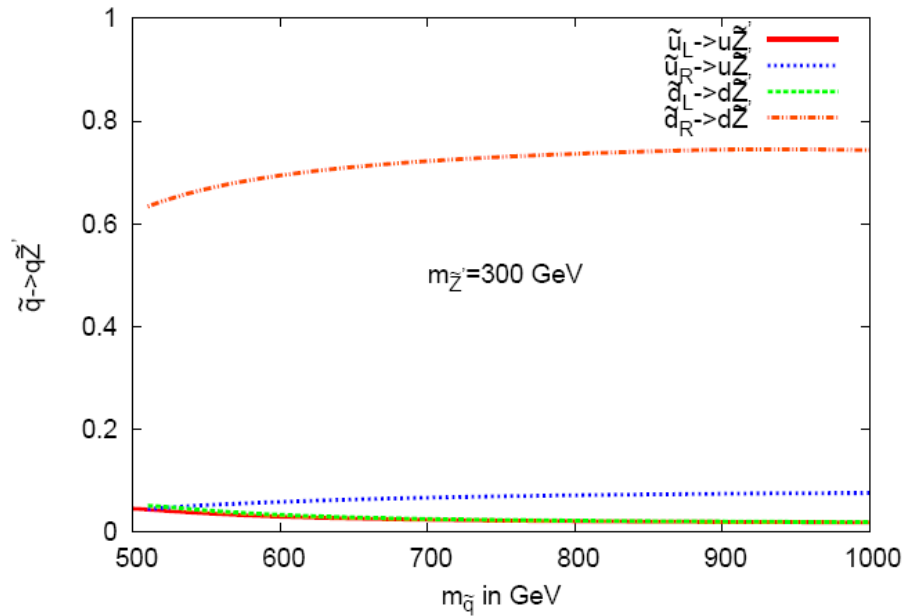
$$pp \rightarrow \tilde{q}\tilde{q} \rightarrow qq\tilde{Z}'\tilde{Z}' \rightarrow qqNN\tilde{N}_1\tilde{N}_1$$

- ▶ Maximised if U(1)' gaugino is the NLSP.

Cascade production of $U(1)'$ gaugino



Squark decay to $U(1)'$ gaugino



Discovery prospect: rough numbers

- ▶ SSD channel: $\text{BR}(\text{NN} \rightarrow \ell \ell \text{WW}) \sim 12\%$
- ▶ Higgs channel: $\text{BR}(\text{NN} \rightarrow h\nu \ell \text{W}) \sim 10\%$.
- ▶ For $\text{U}(1)'$ gaugino as NLSP:

$$\begin{aligned} \text{LHC7} \quad & \sigma(pp \rightarrow Z' \rightarrow \text{NN}) \approx 10 \text{ fb for } m_{Z'} = 1.5 \text{ TeV} \\ & \sigma(pp \rightarrow \tilde{Z}' \tilde{Z}' \rightarrow \text{NN}) \approx 1 \text{ fb for } m_{\text{strong}} = 1.5 \text{ TeV} \\ & \Rightarrow \sigma(h\ell \text{W} (+\text{MET})) \approx 1 \text{ (0.1) fb} \\ \text{LHC14} \quad & \Rightarrow \sigma(h\ell \text{W} (+\text{MET})) \approx 10 \text{ (10) fb} \end{aligned}$$

- ▶ Needs a careful background study.

Conclusion

- ▶ **RHsN: a good dark matter candidate in $U(1)'$ seesaw.**
- ▶ Right freeze-out relic density from $U(1)'$ gaugino exchange + Z' scattering, decay and inverse-decay of the RHN.
- ▶ **Unique LHC signatures: SSD w/ and w/o MET.**
- ▶ Probe the Higgs channel, $h \rightarrow b\bar{b}$, associated with lepton+W, in particular, at displaced vertices.