

CRESST Dark Matter Search

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Bethe Forum 2011

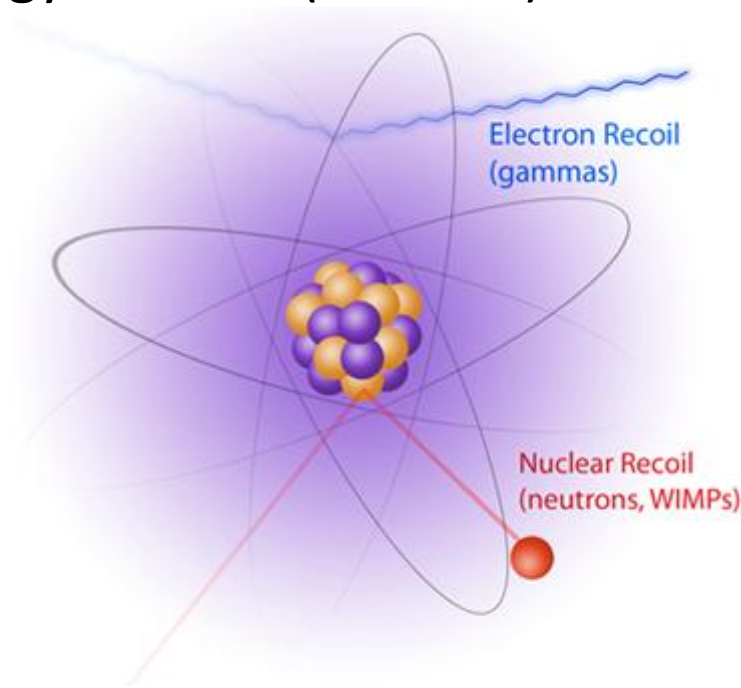
CRESST Collaboration
MPI für Physik, Oxford University,
TU München, Universität Tübingen
Laboratori Nazionali del Gran Sasso

Outline

- WIMP Search with CRESST
- Background Estimation
- Full Analysis Procedure and Results

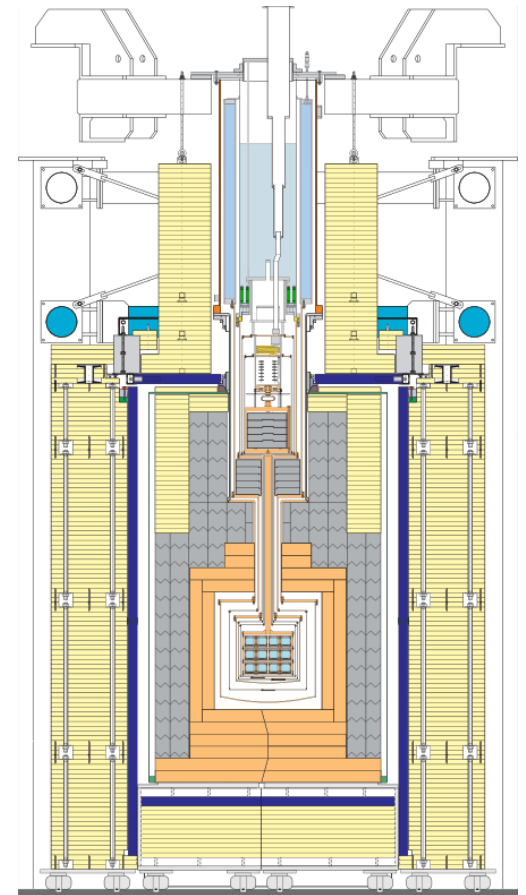
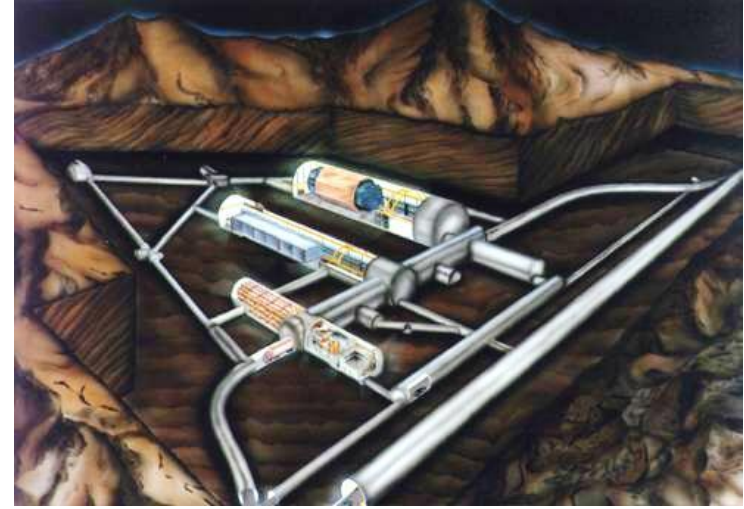
WIMP Detection

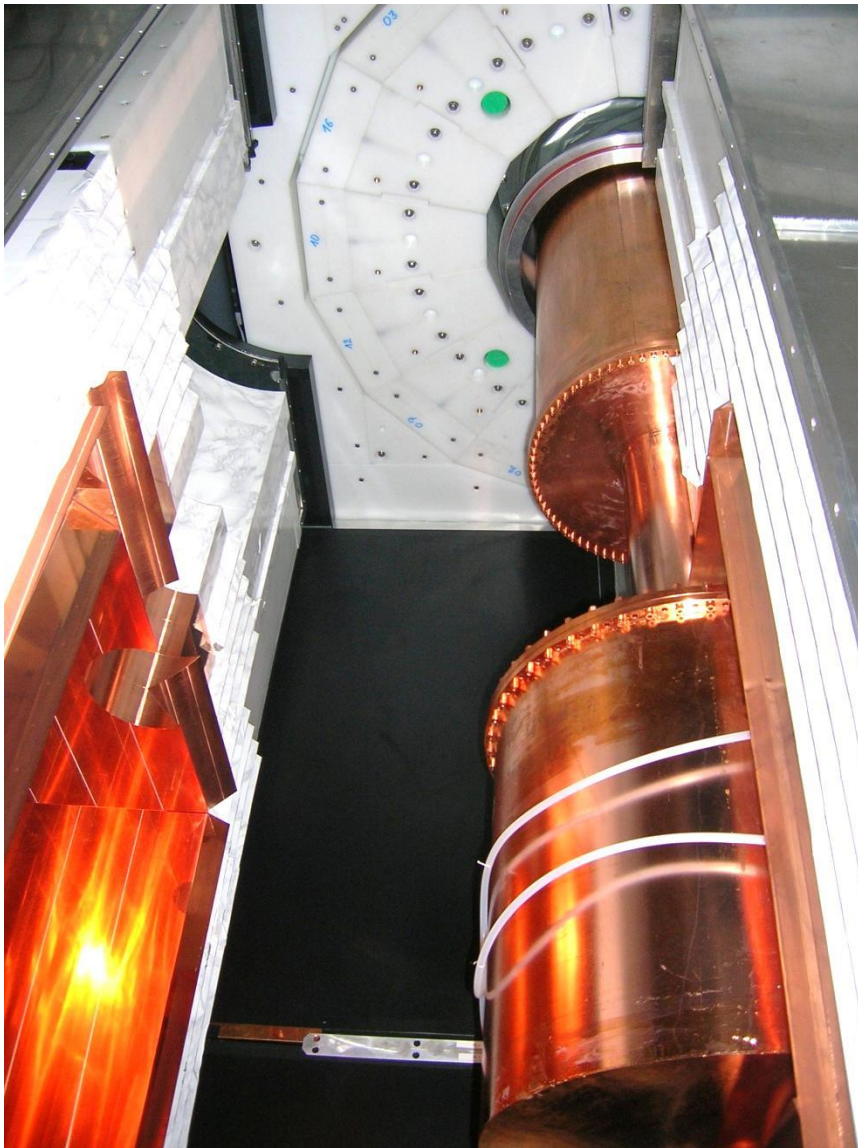
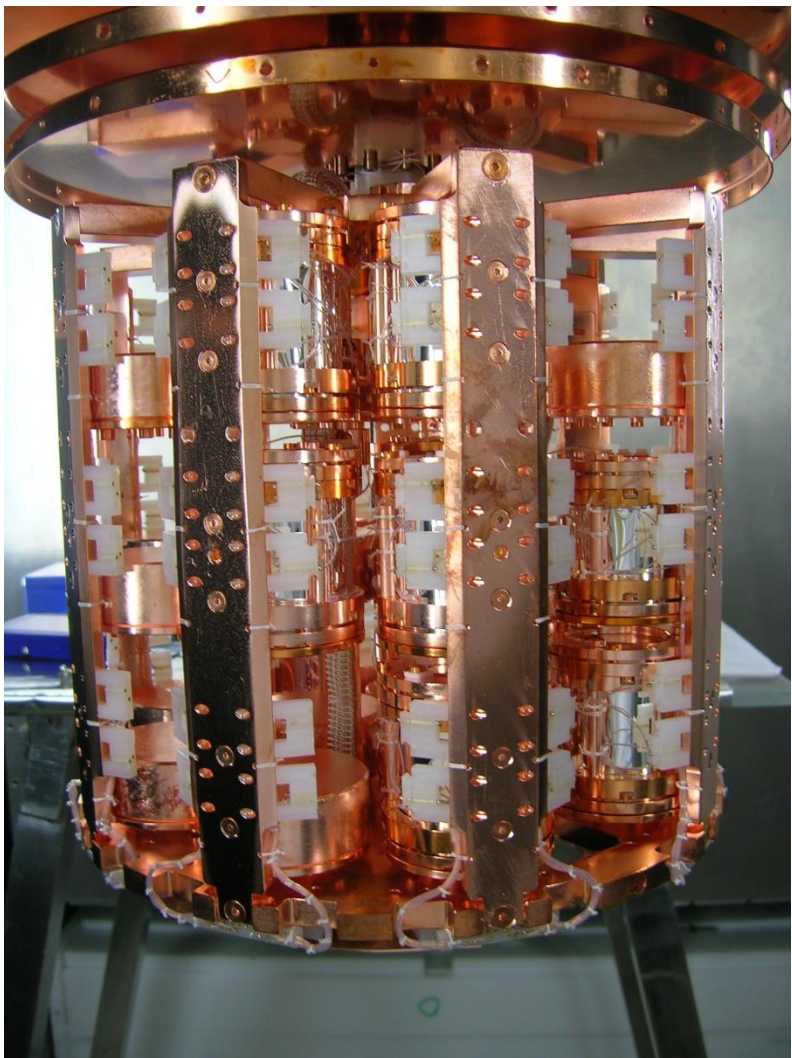
- Expected WIMP mass: GeV – TeV range
- Expected WIMP-nucleon cross section: $< 10^{-5}$ pb
- Coherent scattering off nuclei, $\sigma \sim A^2$
- Signal: nuclear recoils
- Challenges:
 - low rate (few counts per kg year)
 - Small energy transfer (few keV)



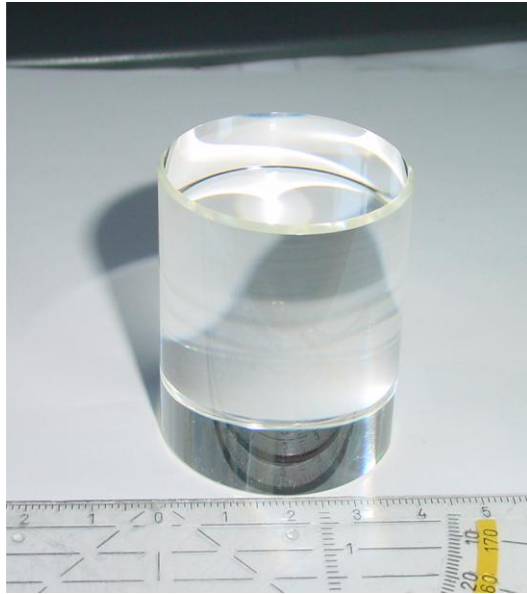
CRESST at LNGS

- Located in Hall A of the LNGS deep underground laboratory
- Shielding of setup:
 - PE (12 tons)
 - Active muon veto
 - Radon box
 - Lead shielding (24 tons)
 - Copper (10 tons)
- Detector carrousel
 - Can currently house up to 33 detector modules at ~ 10 mK (total target mass for CaWO_4 : ~ 10 kg)

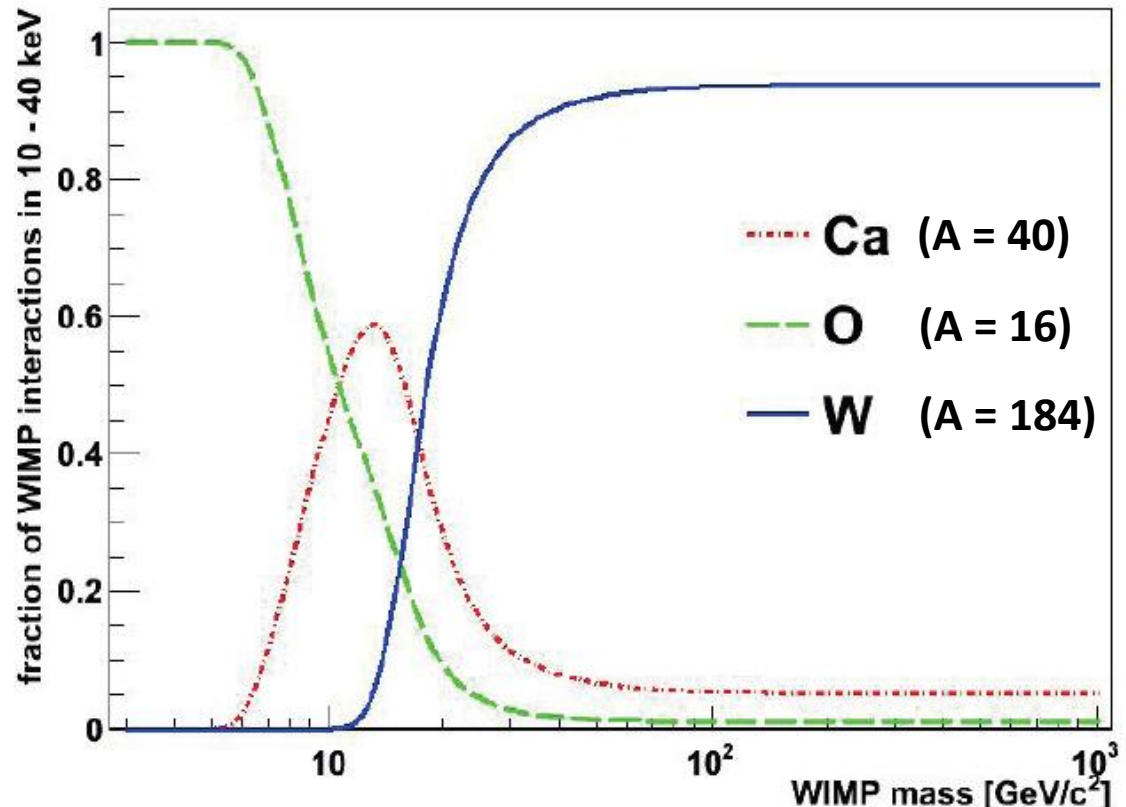




CaWO₄ Single Crystals as Target Material

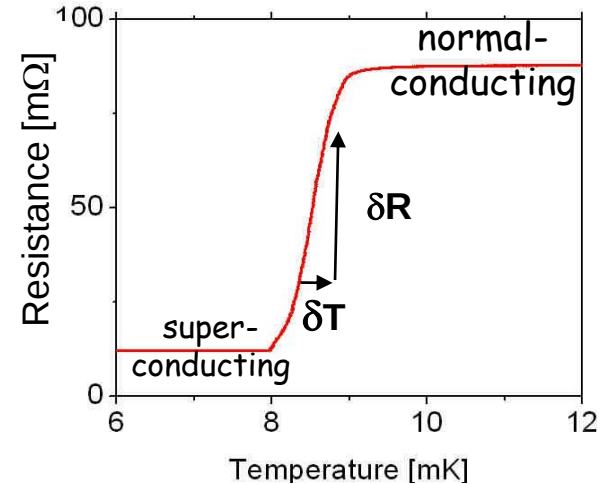
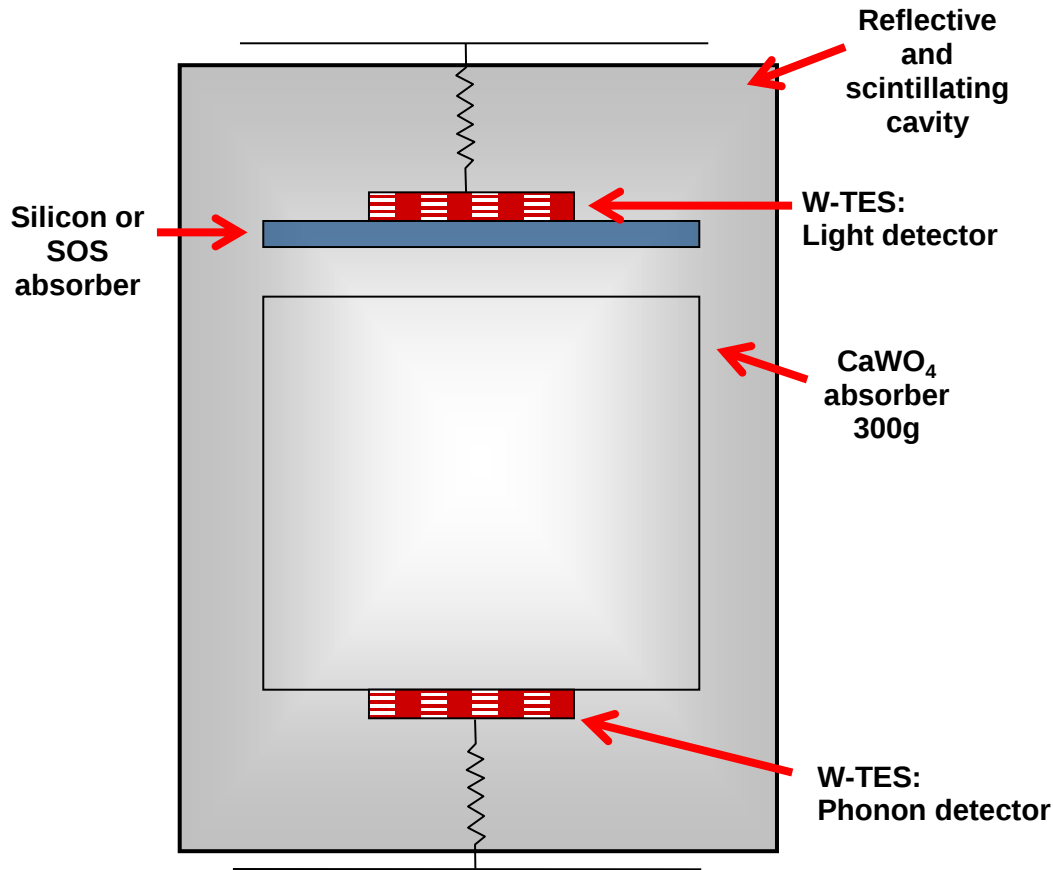


300 g CaWO₄ crystal
Diameter: 40 mm
Height: 40 mm

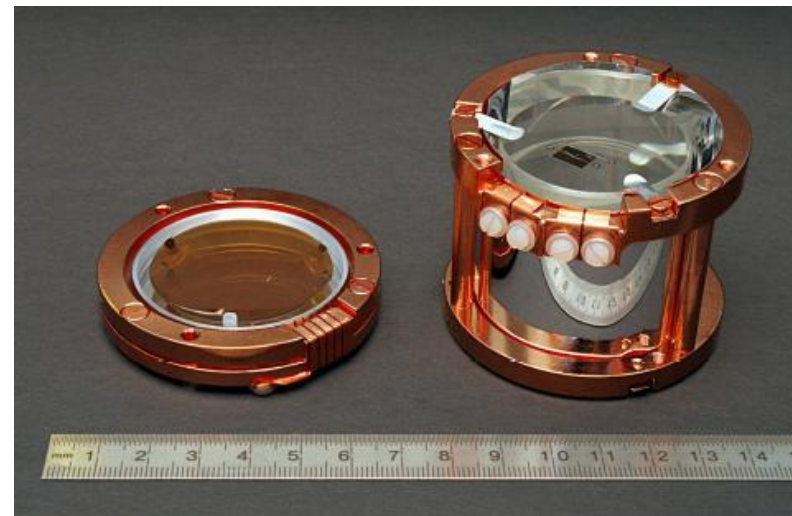


- multi-material target CaWO₄: **unique feature of CRESST**
 - probe for light and heavy WIMPs

CRESST Detector Design

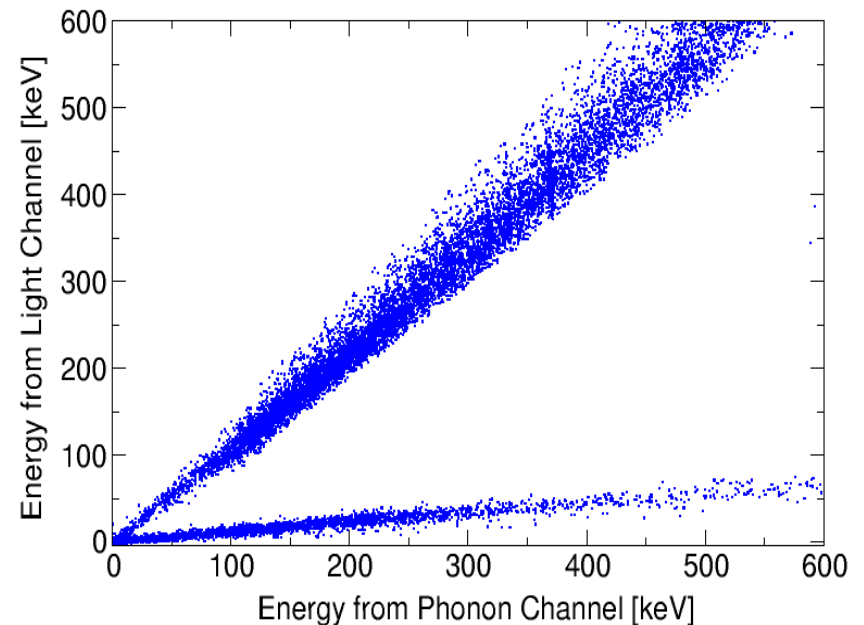


- Independent measurement of two signals of a particle interaction:
- Phonon signal ($\approx$ complete energy deposition)
 - Scintillation light signal



Detector Response

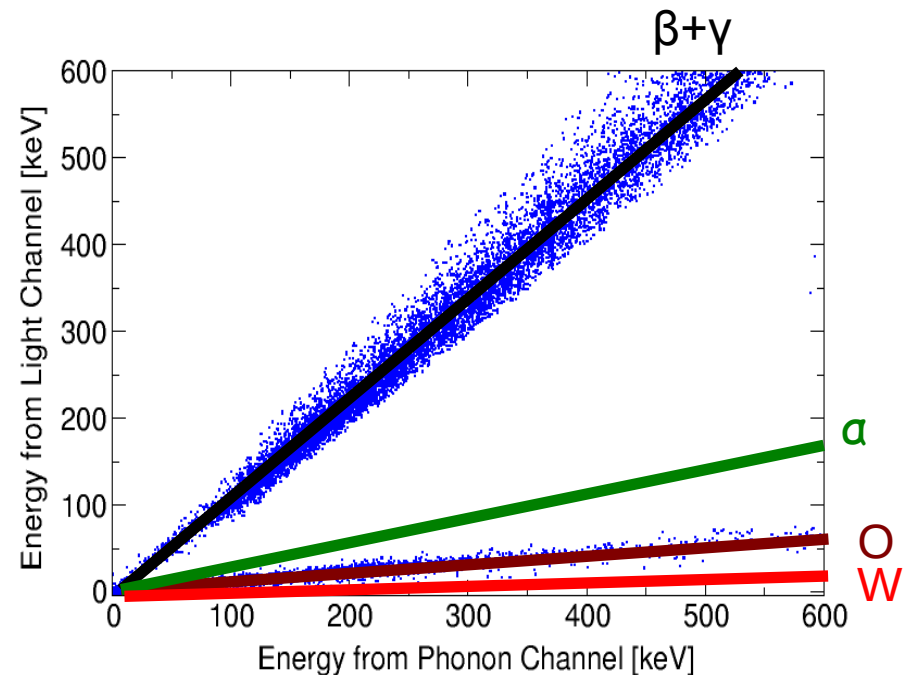
- Energy deposition of a particle interaction visible in both channels:
 - Phonon channel (>95% of total energy)
 - Light channel (~5% for electron recoils)
- Scintillation light output for nuclear recoils quenched compared to electron recoils (er)



Calibration measurement with
neutron source

Detector Response

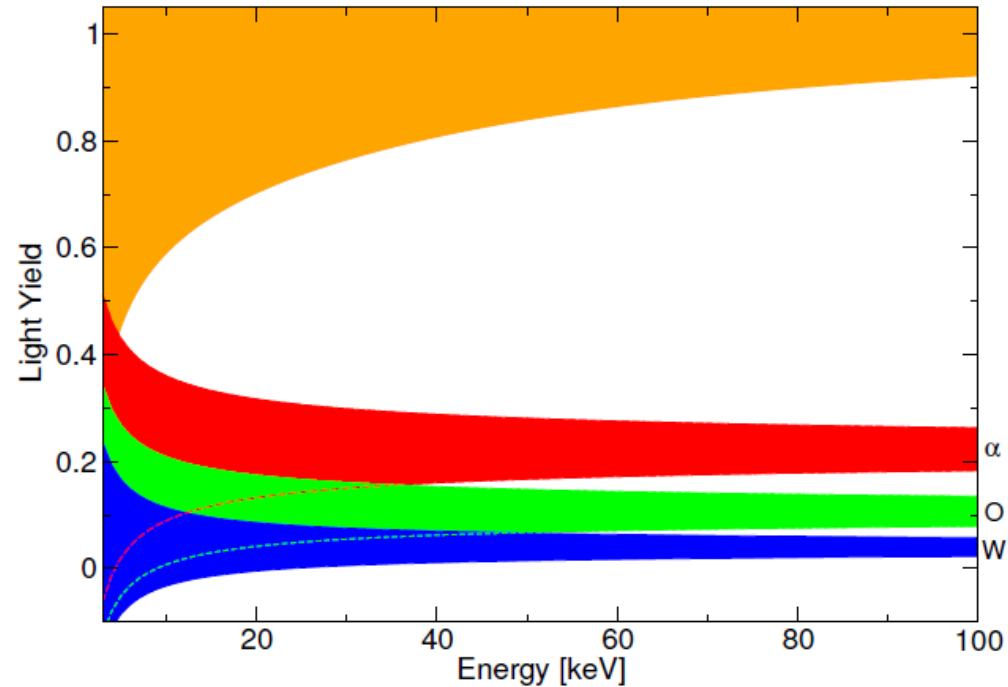
- Energy deposition of a particle interaction visible in both channels:
 - Phonon channel (>95% of total energy)
 - Light channel (~5% for electron recoils)
- Scintillation light output for nuclear recoils quenched compared to electron recoils (er)
- *Quenching Factors:*
 $QF(W) < QF(Ca) < QF(O) < QF(\alpha) < QF(er) = 1$



Calibration measurement with
neutron source

Event Type Discrimination

- Light yield defined as $E_{\text{light}} / E_{\text{phonon}}$
- $QF(er) = 1$ (by def)
- $QF(\alpha) \approx 0.22$
- $QF(O) \approx 0.1$
- $QF(Ca) \approx 0.06$
- $QF(W) \approx 0.04$
- In the signal region, the nuclear recoil bands overlap due to the finite resolution of the light detector

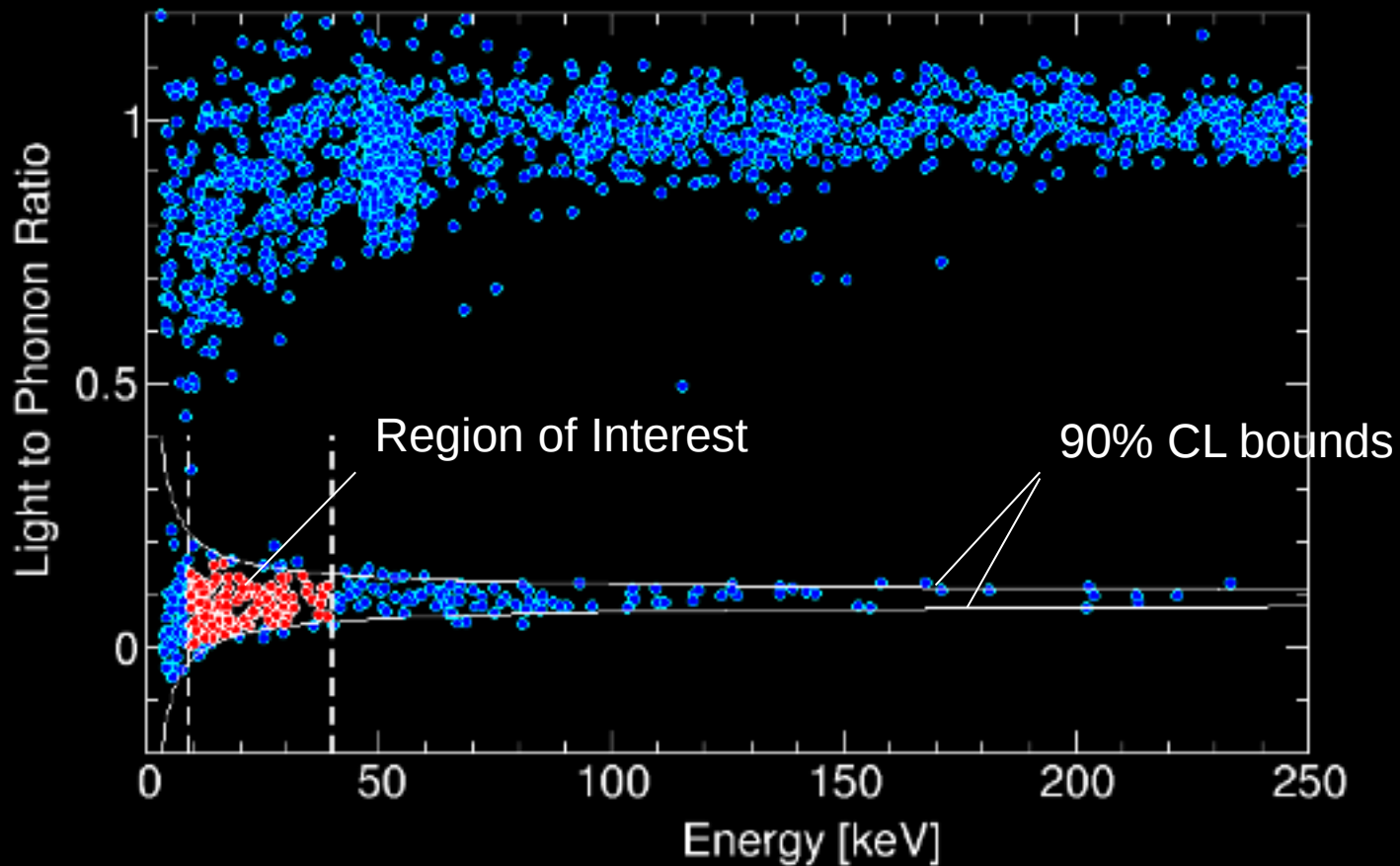


Excellent rejection of
electron recoils

Neutron Calibration

Calibration campaign with neutron source (Am-Be)

Ch51/52: neutron calibration

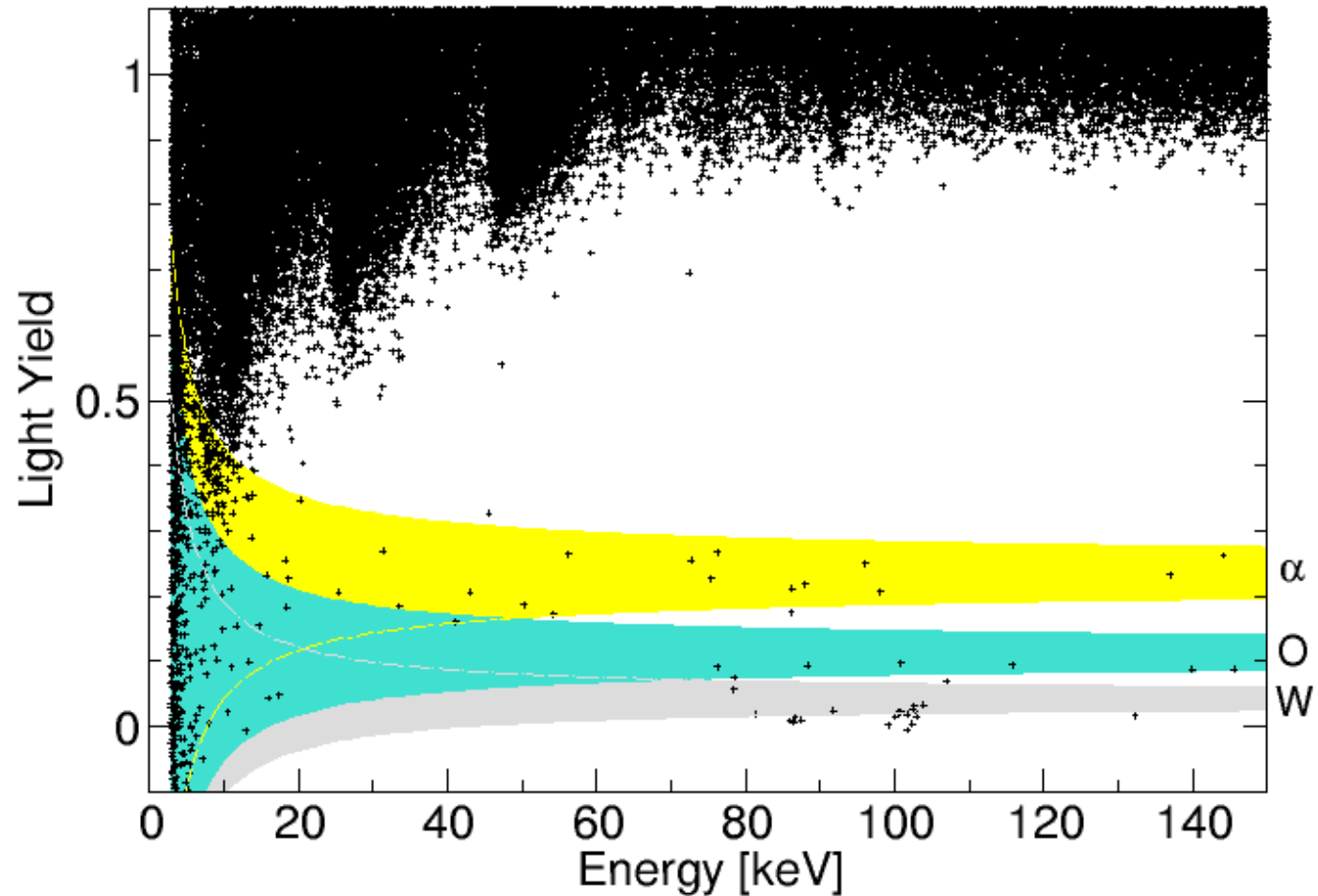


Excellent nuclear recoil discrimination seen in calibration run

The Latest CRESST Run

- June 2009 to April 2011
- 18 detector modules installed
 - 10 complete detector modules (9 CaWO_4 and 1 ZnWO_4)
 - 8 modules used for analysis
 - 7 additional modules for vetoing of coincident events (i.e. events with multiplicity > 1)
- Total exposure after cuts: 730 kg d

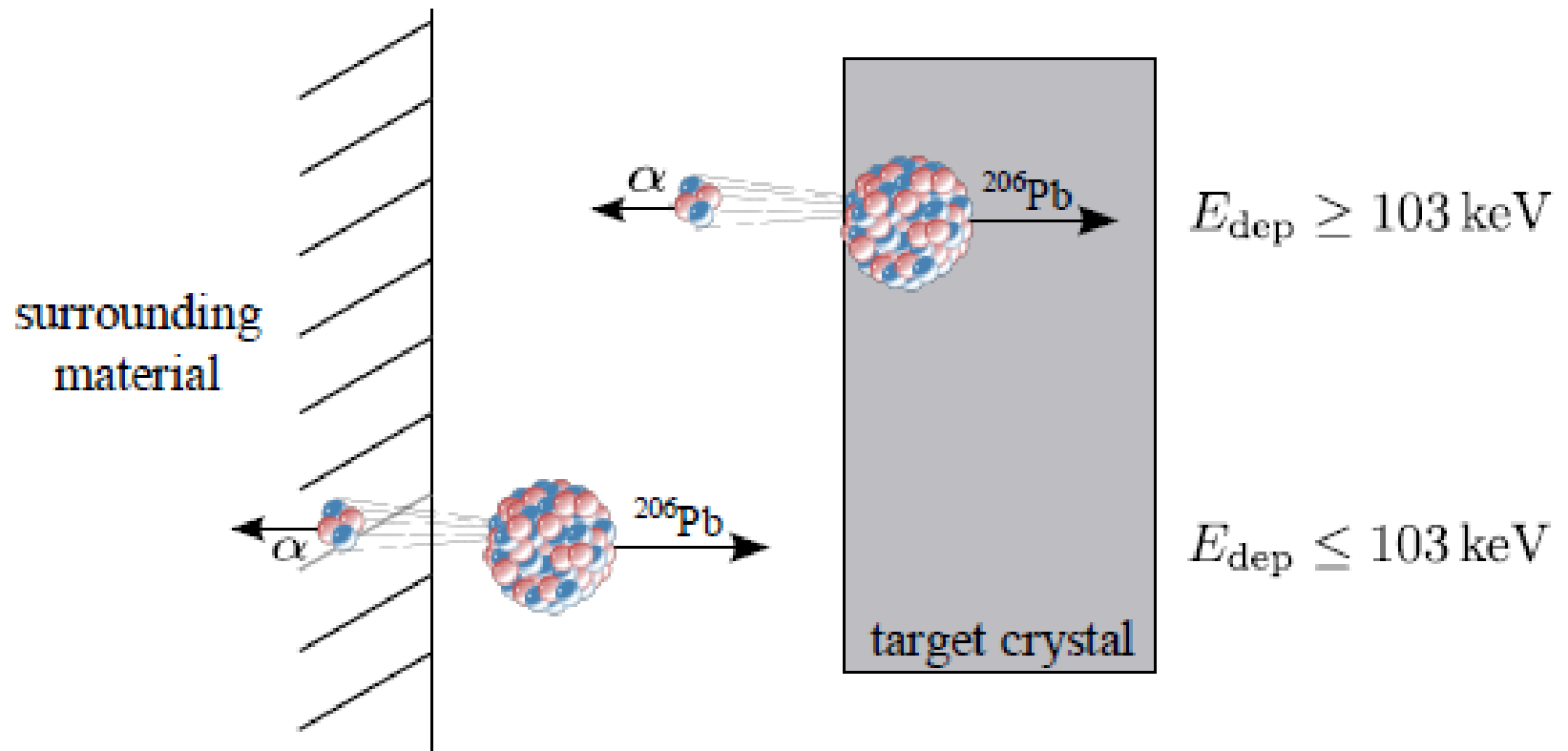
Background Run Example



Possible Background in Signal Region

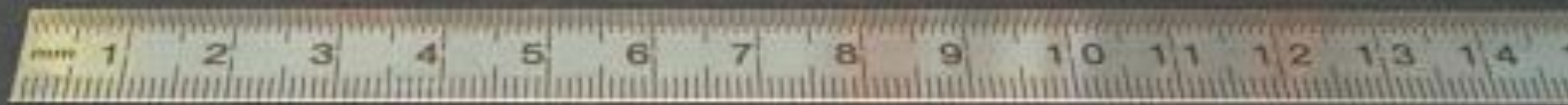
- Gamma leakage (definition of analysis threshold: 1 leaked event per detector module in whole data set)
- Alpha decays (detailed in next slides)
 - Alpha particle interacts with target crystal
 - Heavy nucleus (Pb) interacts with target crystal
- Neutron Background
 - Neutrons produced by natural radioactivity
 - Muon induced neutrons (can be estimated by multiplicity of muon induced events, see below)

Alpha Background



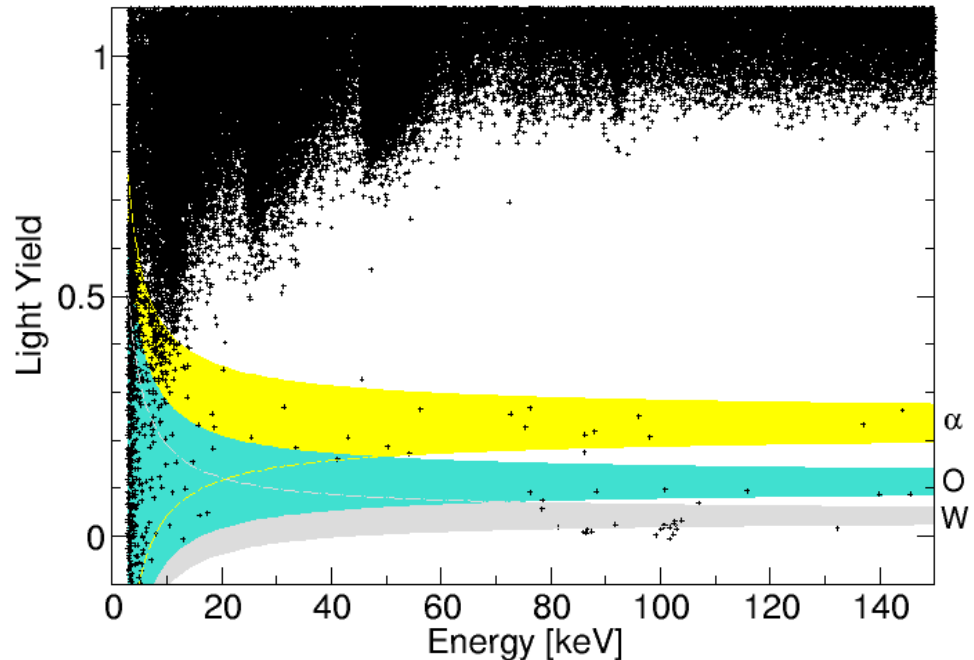
Alpha Background

- Possible alpha decay scenarios:
 - Internal alpha decay
 - several MeV energy deposition in target crystal, no leakage in signal region
 - Surface alpha decay
 - alpha hits crystal, heavy nucleus (Pb) hits scintillating /reflective foil
 - several MeV energy deposition in target crystal
 - Heavy nucleus hits crystal, alpha hits scintillating/reflective foil
 - high light yield due to scintillating foil
 - Heavy nucleus hits crystal, alpha hits passive detector material (i.e. clamp)
 - ≈ 100 keV deposited in crystal by heavy nucleus (signal region: up to 40 keV)
 - Nucleus implanted in passive detector material (i.e. clamps)
 - Alpha hits target crystal, previous loss of arbitrary amount of energy in clamp
 - degenerated alpha
 - Heavy nucleus hits target crystal
 - leakage of lead recoils in signal region



Overview: Observed Event Classes

- Bulk of events are electron recoils
- Degenerated alphas
- Pb recoils at 103 keV and lower
- Nuclear recoil events on O, Ca and W



Complete data for whole run of one exemplary module

Parametrization of Background Events

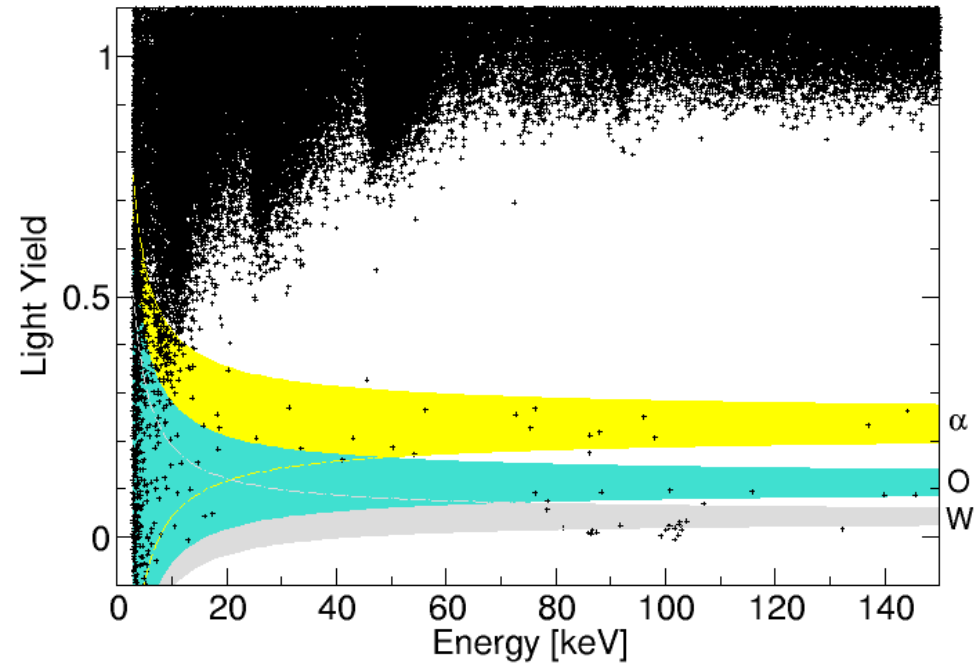
I – Electron Recoils

- Leakage expected at **high light yield** and **low energy threshold**
- Over complete run: High statistics of electron recoils
 - → very good description of electron recoil band at all energies
 - → definition of signal region:
 - $E_{\min}$: 1 leakage event for whole run and detector module.
 $E_{\min}$ thus individual for each detector
 - $E_{\max} = 40$ keV (contribution of WIMP signal above 40 keV negligible.)

Parametrization of Background Events

II – Alpha Recoils

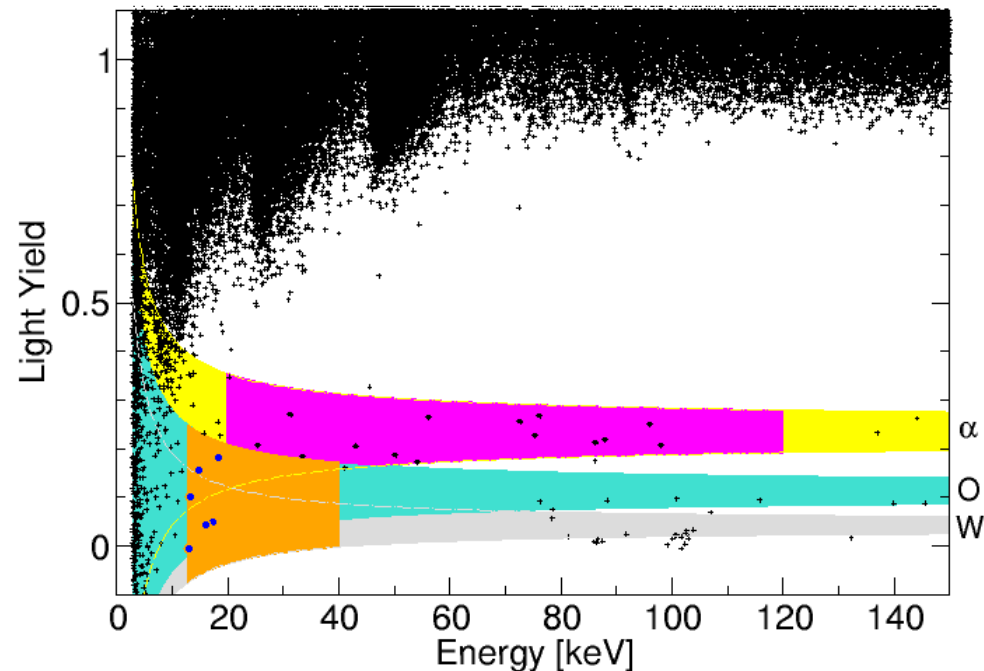
- Degenerated alphas observed at all energies



Parametrization of Background Events

II – Alpha Recoils

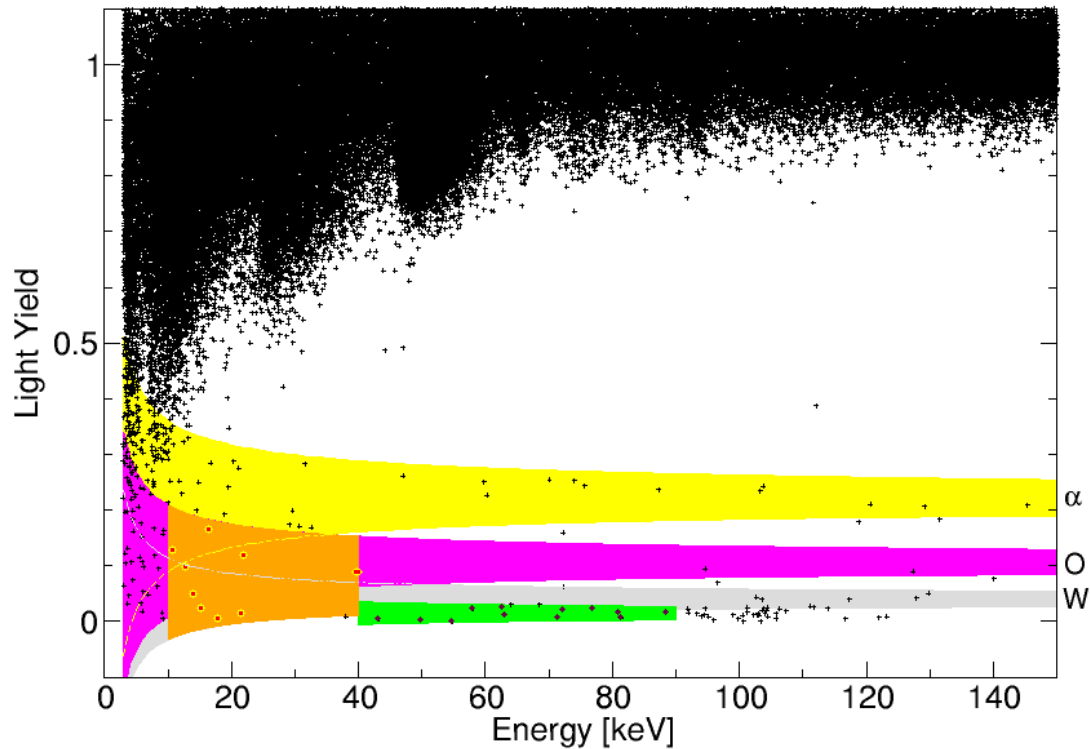
- Degenerated alphas observed at all energies
- Define **overlap free reference region** for alpha recoils:
Observed energy spectrum of these events: flat
 $dN_{\alpha}/dE = \text{const}$
(but individual for each detector)
- Based on the resolution of the light detector, the total background contribution of alpha recoils can be estimated for all detectors



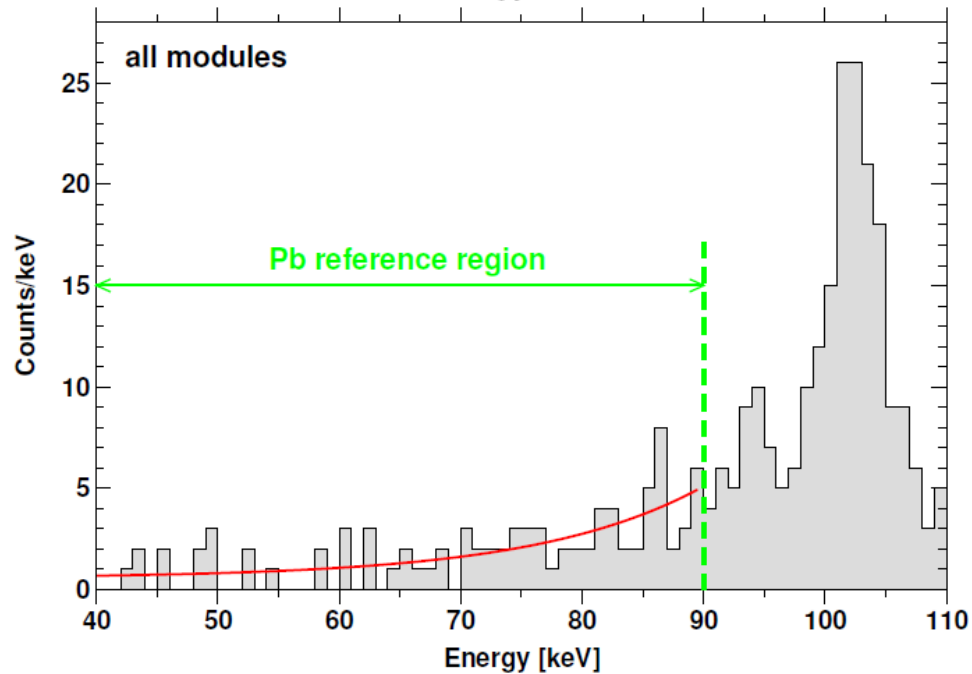
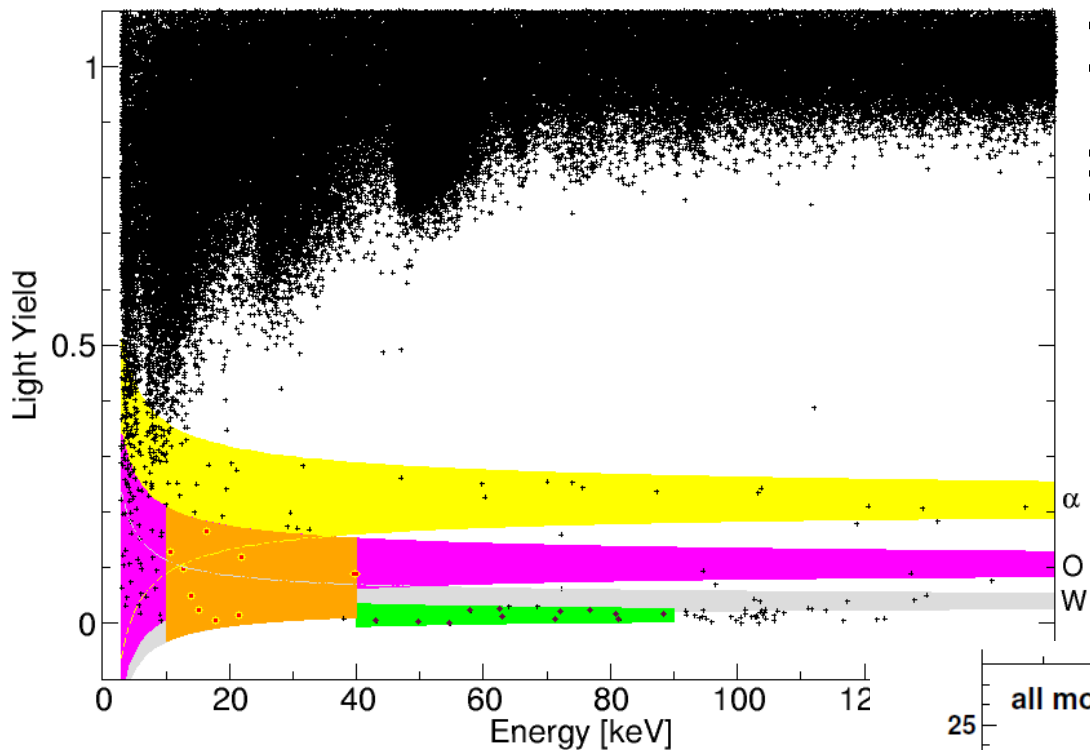
→ Simple estimate:
9.2 events for all detectors and complete run

Parametrization of Background Events

III – Lead Recoils



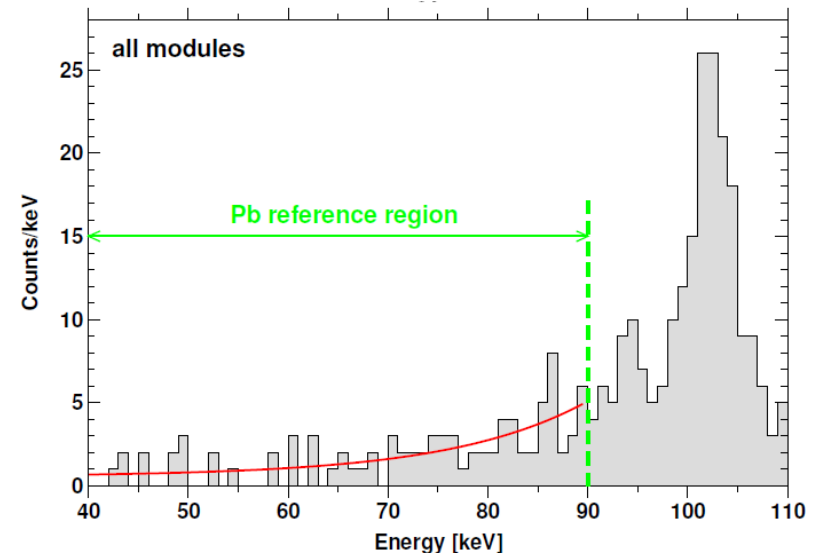
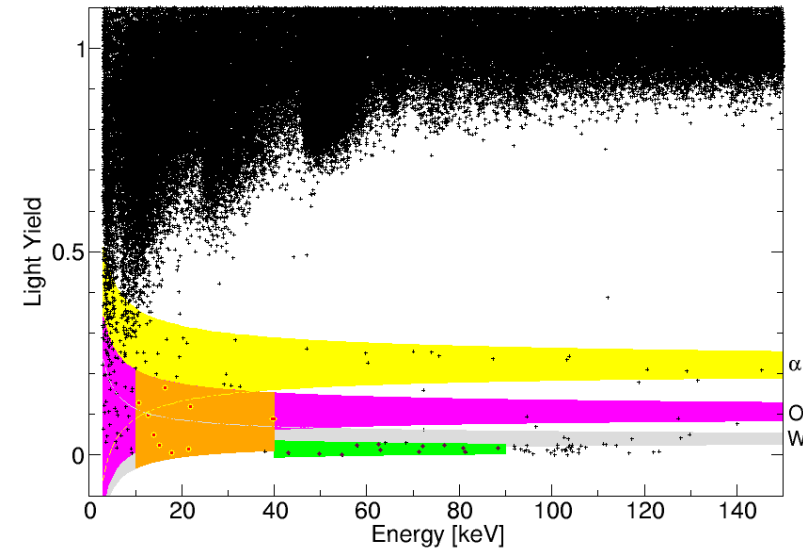
Background Events Decoils



Parametrization of Background Events

III – Pb Recoils

- Distribution of Pb recoils can have tail to low energies, they behave similar to W-recoils
- Define **overlap free reference region**
- Parametrize observed spectrum
- Extrapolate to signal region
- Simple estimate:
17 events for all detectors and complete run



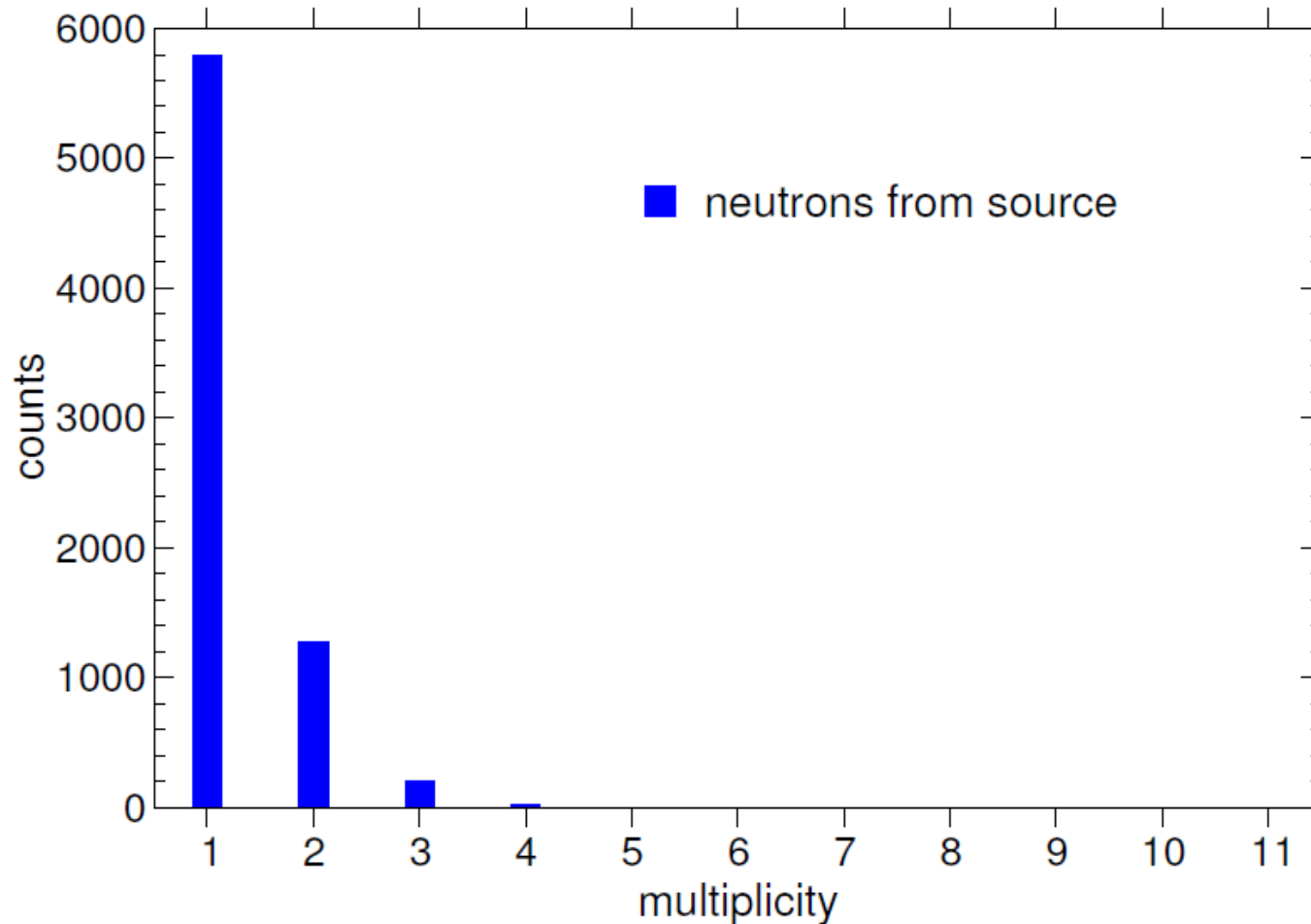
Parametrization of Background Events

IV – Neutrons

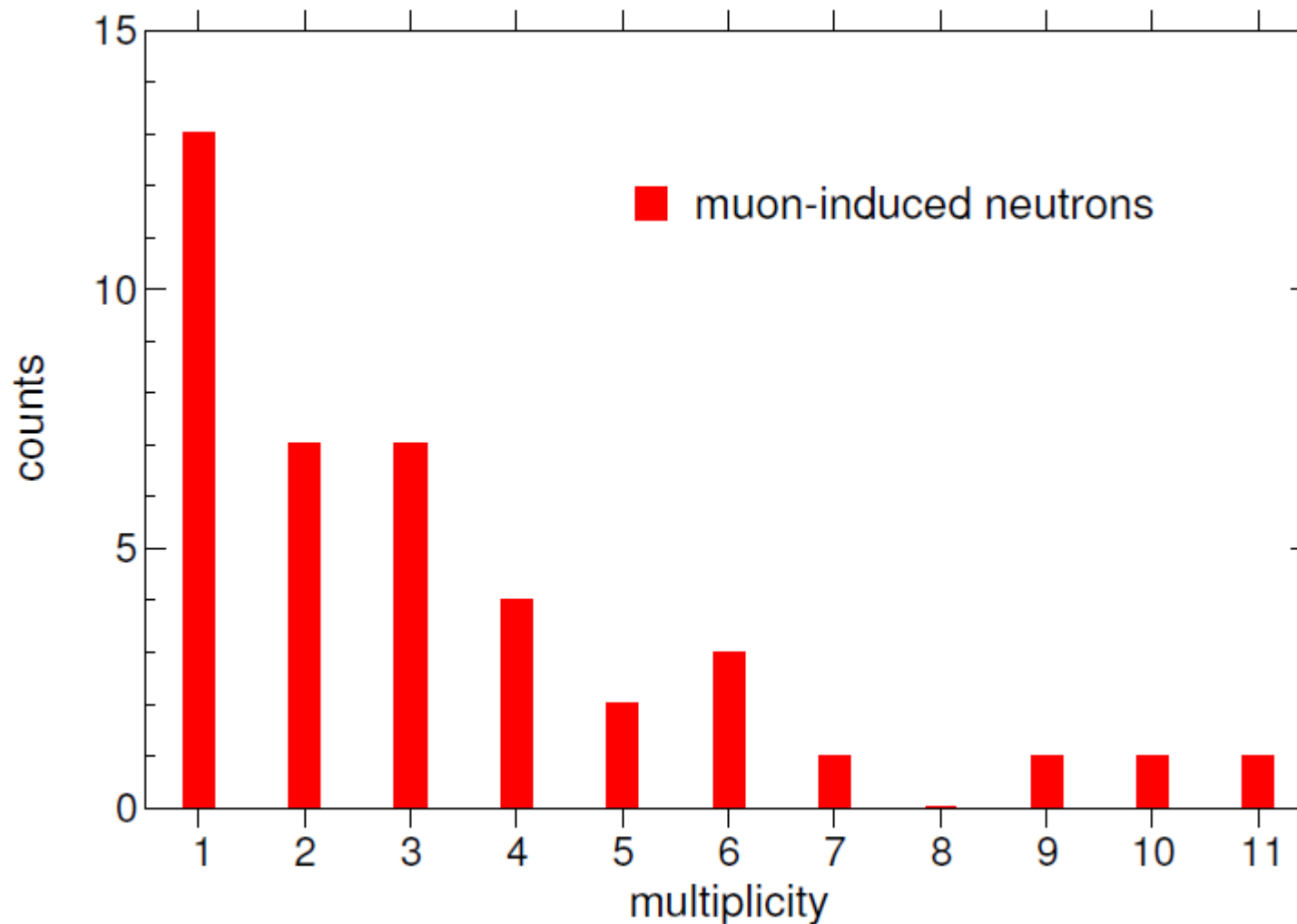
- Based on stoichiometric and kinematic arguments, neutrons are mainly visible as oxygen recoils (similar to light WIMPs)
- Event multiplicity: neutrons can interact in more than one detector module, leading to coincident events in the detector array
- In the data set, 3 coincident events were observed in the acceptance region:
 - Two with multiplicity 3
 - One with multiplicity 5
- Causes:
 - a) radioactive processes inside neutron shielding
 - b) neutrons are muon induced
- To estimate neutron background: Calculate the expected number of events with multiplicity = 1, if 3 events with multiplicity > 1 were observed in the signal region

Observed Event Multiplicity

Neutron Calibration AmBe source

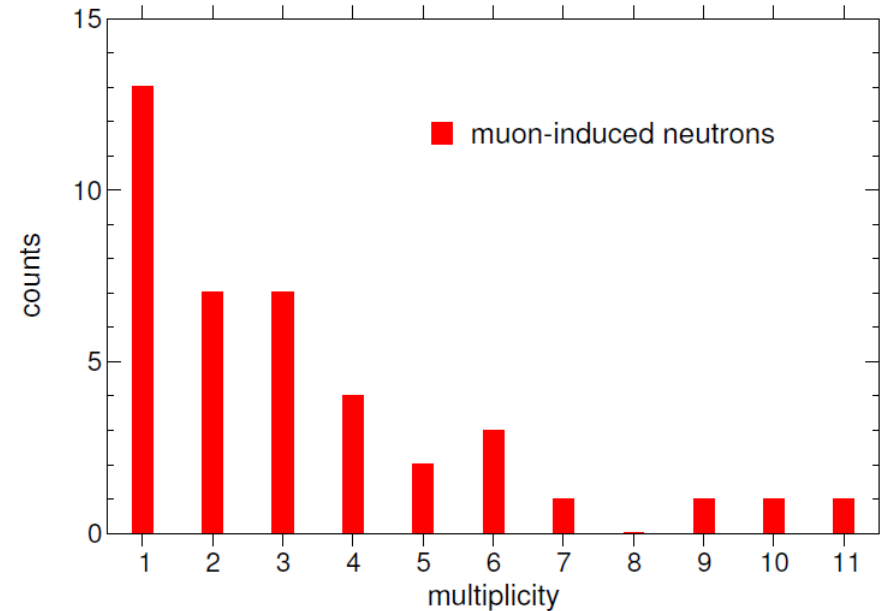
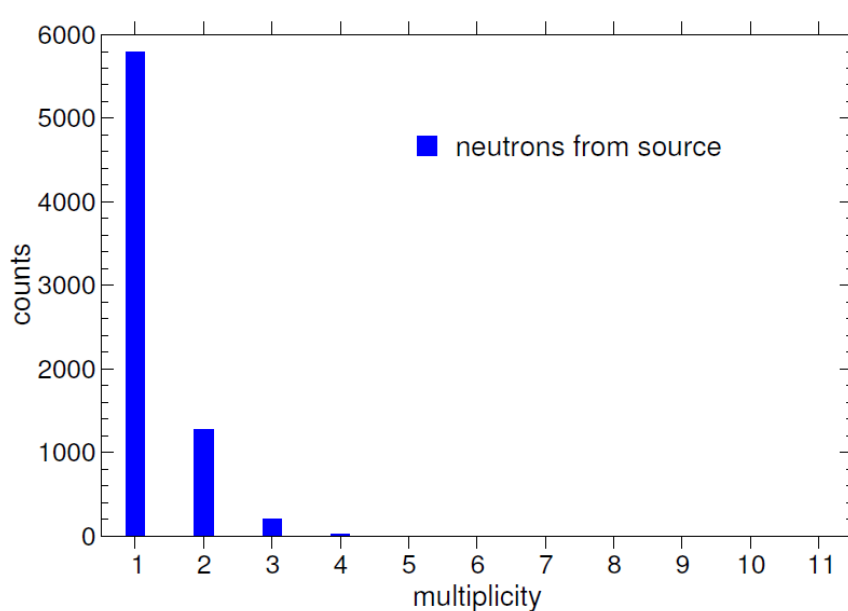


Observed Event Multiplicity Muon Coincident Neutrons



Parametrization of Background Events

IV – Neutrons



- Neutron source: single/coincident ≈ 3.8
 - $\rightarrow$ if 3 coincident events are observed, 11.4 single events are estimated
- Muon coincident: single/coincident ≈ 0.5
 - $\rightarrow$ if 3 (unvetod by the muon veto) coincident events are observed, 1.5 single events are estimated

Simple Estimates Summary

- Electron Recoils: 8 events
 - Degenerated alphas: 9.2 events
 - Pb recoils: 17 events
 - Neutron background: 1.5 - 11.4 events
 - Total: 35.7 – 45.6 events
 - Total observed events: 67
-
- Excess above considered background events observed

Full Likelihood Analysis

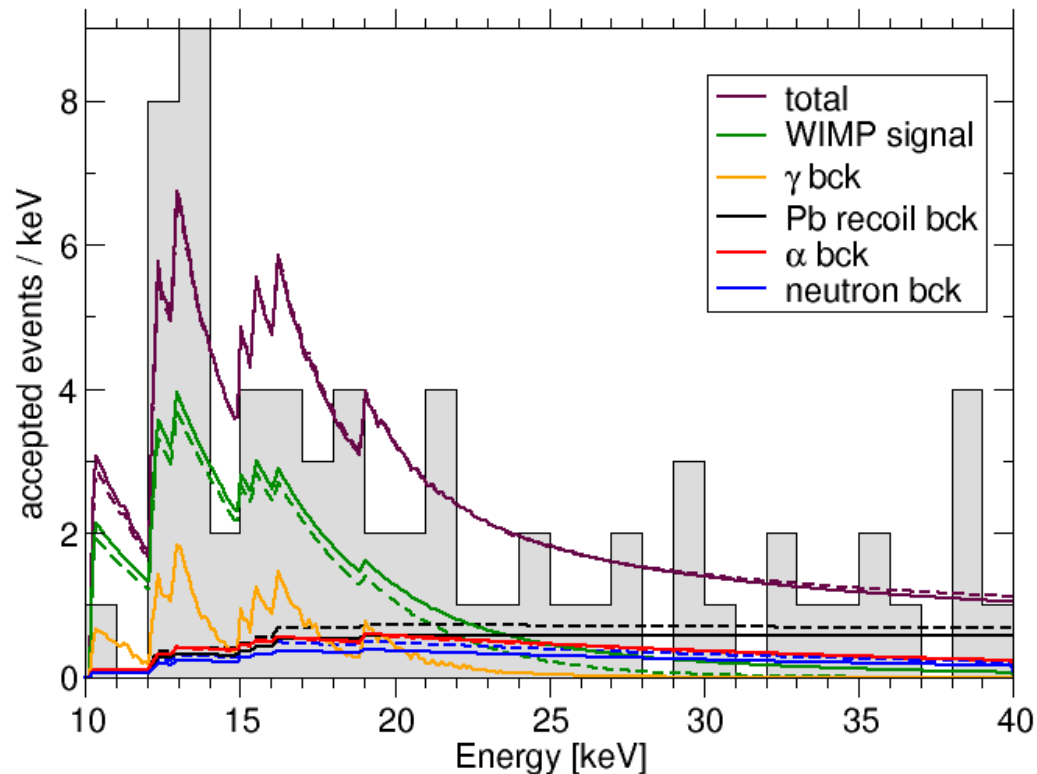
- The full analysis takes into account
 - All spectra in the reference regions
 - The light yield of each event
 - Spectral distribution of neutron calibration
 - Observed and expected multiplicities
 - the individual resolution and signal region of each detector module
 - A possible WIMP signal to explain the excess events
- Full likelihood analysis yields similar results to simple estimates (as is expected)

Full Likelihood Analysis Results

- Likelihood function shows two maxima for the WIMP parameters (M1 and M2)
 - M1 is the global maximum
 - M2 only slightly disfavoured with respect to M1
- Best fit values:

	M1	M2
e/γ -events	8.00 ± 0.05	8.00 ± 0.05
α -events	$11.5^{+2.6}_{-2.3}$	$11.2^{+2.5}_{-2.3}$
neutron events	$7.5^{+6.3}_{-5.5}$	$9.7^{+6.1}_{-5.1}$
Pb recoils	$15.0^{+5.2}_{-5.1}$	$18.7^{+4.9}_{-4.7}$
signal events	$29.4^{+8.6}_{-7.7}$	$24.2^{+8.1}_{-7.2}$
m_χ [GeV]	25.3	11.6
σ_{WN} [pb]	$1.6 \cdot 10^{-6}$	$3.7 \cdot 10^{-5}$

The Two Maxima: Phonon Channel

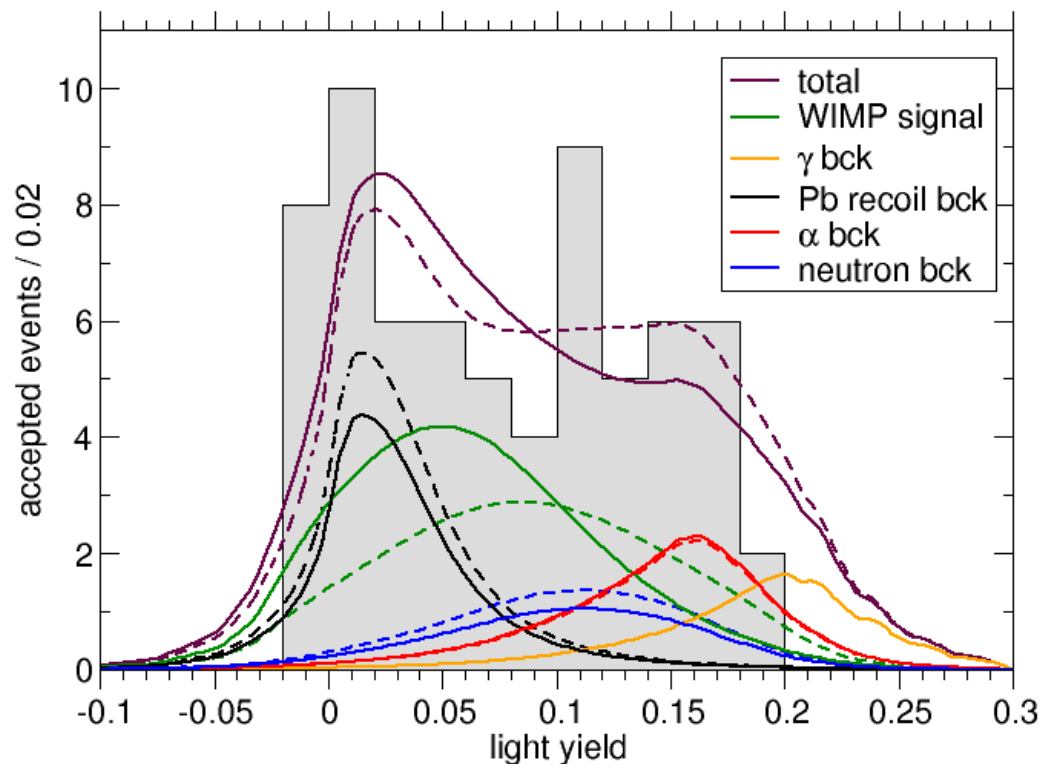


Solid lines: M1
Dashed lines: M2

Spectral shape: Pb recoils and alpha background **do not** reproduce observed excess events.

Gamma events?

The Two Maxima: Light Channel



Solid lines: M1
Dashed lines: M2

Gamma events?

Gamma events are at high light yield compared to excess events.

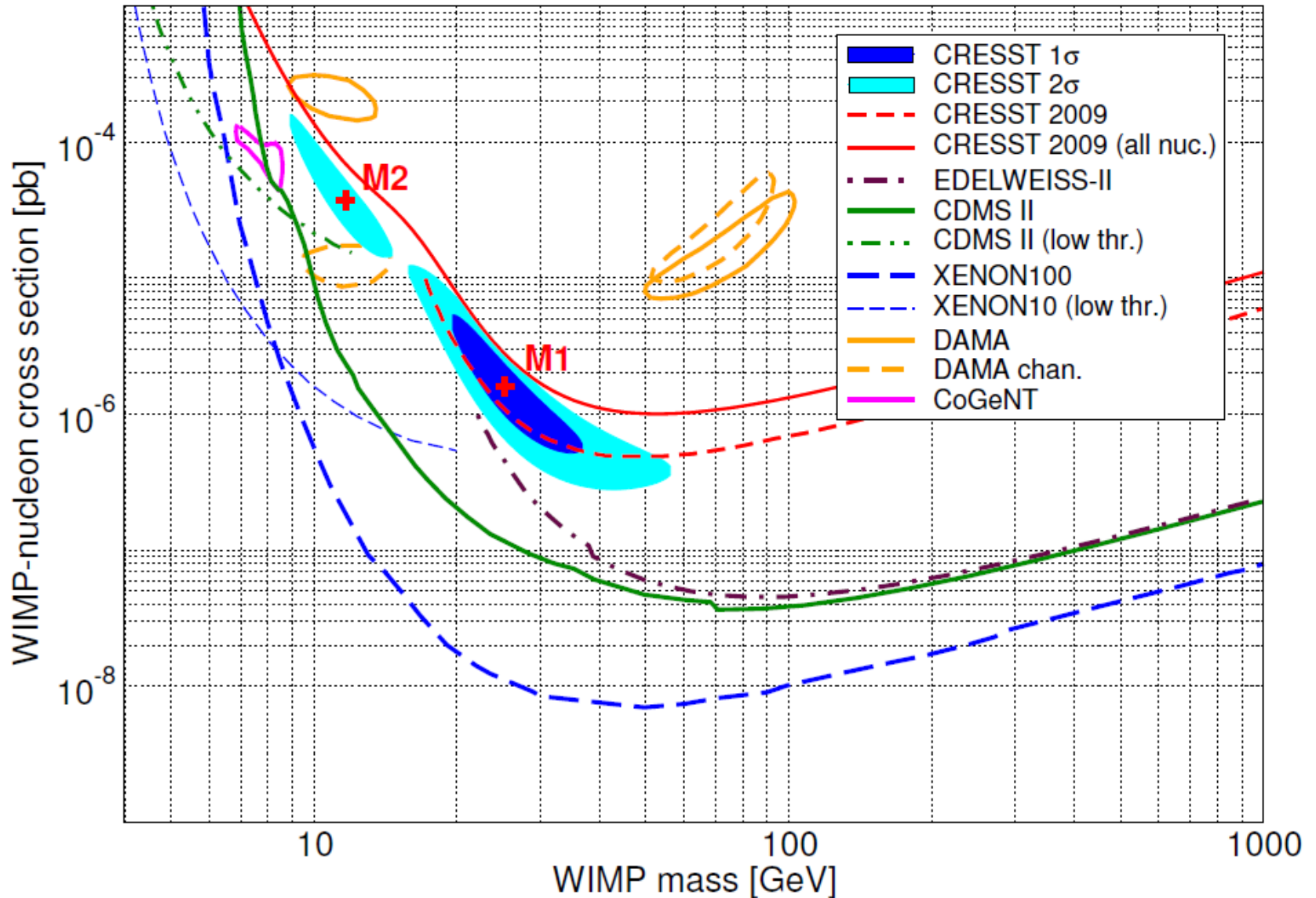
The Two Maxima: Composition

- M1 and M2 give nearly identical spectra in the phonon channel
- Composition of recoiling nuclei is different, thus leading to a different light yield distribution of the signal events:
 - M1: Mass ≈ 25.3 GeV
 7% oxygen, 25% calcium, 69% tungsten
 - M2: Mass ≈ 11.6 GeV
 52% oxygen, 48% calcium, 0% tungsten

Significance

- Fitting the observed events once with a WIMP contribution and once without
- Calculate the likelihood ratio for both maxima:
M1: 4.7σ
M2: 4.2σ
- Interpretation: Considered background is not sufficient to explain observation
 - **additional source of events needed**
 - WIMPs are a **possible** source for the excess events
- However: Background is still relatively large
 - Main background source coming from non-scintillating clamps
 - goal for next run: clarify possible signal by reducing the observed background

WIMP Parameter Space



Concluding Remarks

- Successful physics run of CRESST:
 - Total exposure after cuts: 730 kg days
 - 67 events observed in signal region
- Detailed background analysis does not explain the observed events. An additional source for events is needed.
 - The expected properties of WIMPs can explain the signal excess
- Paper submitted to European Physical Journal C, arXiv: 1109.0702
- Outlook:
 - New clamps to reduce alpha and Pb background significantly currently in production
 - Additional shielding currently under discussion
 - Increase of target mass planned
 - Next run scheduled to start early next year