

Supersymmetry at the LHC

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Based on the work done with Peter Skands [arXiv:1109.5852] and Biswarup Mukhopadhyaya [*in preparation*]

Part I: Implementation of SUSY in Pythia 8

Part II: Interpretation of ATLAS SUSY limits for third-
generation squarks

Part I: Implementation of SUSY in Pythia 8



SUSY in Pythia 8

Read in new masses and couplings

- Can read in files SUSY Les Houches v2 (SLHA2) format
- Completely flavour general couplings [Bozzi et. al. 2007]

$$\begin{pmatrix} \tilde{u}_1 \\ \tilde{u}_2 \\ \tilde{u}_3 \\ \tilde{u}_4 \\ \tilde{u}_5 \\ \tilde{u}_6 \end{pmatrix} = R^u \begin{pmatrix} \tilde{u}_L \\ \tilde{c}_L \\ \tilde{t}_L \\ \tilde{u}_R \\ \tilde{c}_R \\ \tilde{t}_R \end{pmatrix} ; \begin{pmatrix} \tilde{d}_1 \\ \tilde{d}_2 \\ \tilde{d}_3 \\ \tilde{d}_4 \\ \tilde{d}_5 \\ \tilde{d}_6 \end{pmatrix} = R^d \begin{pmatrix} \tilde{d}_L \\ \tilde{s}_L \\ \tilde{b}_L \\ \tilde{d}_R \\ \tilde{s}_R \\ \tilde{b}_R \end{pmatrix}$$

SUSY in Pythia 8

Allows processes with

- CP violation
- Flavour violation
- R-parity violation

Calculates all decay widths for 2-body decays of squarks, sleptons, gluino, neutralinos and charginos.

Cross sections

- Pair production of all strongly charged superparticles
- Pair production of Neutralinos and Charginos

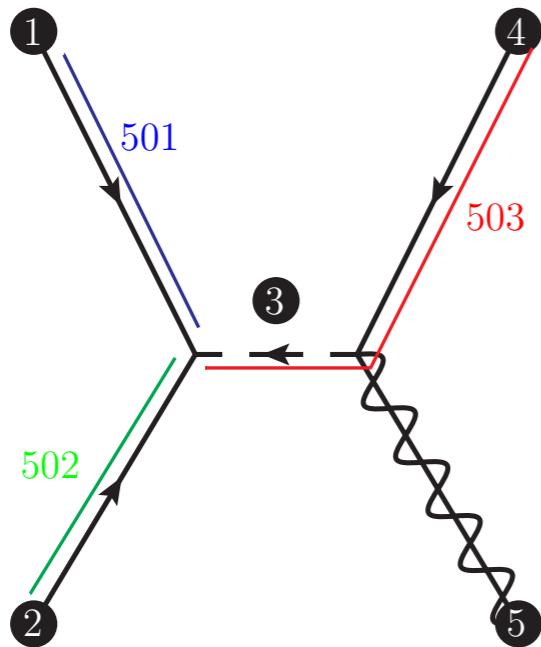
Production processes

Chargino and neutralino production	$q\bar{q}2\chi_0\chi_0$, $q\bar{q}2\chi_+\chi_0$, $q\bar{q}2\chi_+\chi_-$.
Gaugino squark production	$qg2\chi_0\text{squark}$, $qg2\chi_+\text{squark}$.
Gluino production	$gg2\text{gluino}\text{gluino}$, $q\bar{q}2\text{gluino}\text{gluino}$.
Squark-gluino production	$qg2\text{squark}\text{gluino}$
Squark-pair production	$gg2\text{squark}\text{antisquark}$, $q\bar{q}2\text{squark}\text{antisquark}$  $qq2\text{squarksquark}$ 
RPV resonant squark production	$qq2\text{antisquark}$

R-parity violating production

Three types of RPV couplings: LLE, LQD, UDD

$(\lambda_{ijk}, \lambda'_{ijk} \text{ and } \lambda''_{ijk})$



$q_i q_j \rightarrow \tilde{q}_k^* \text{ via } \lambda''_{ijk}$

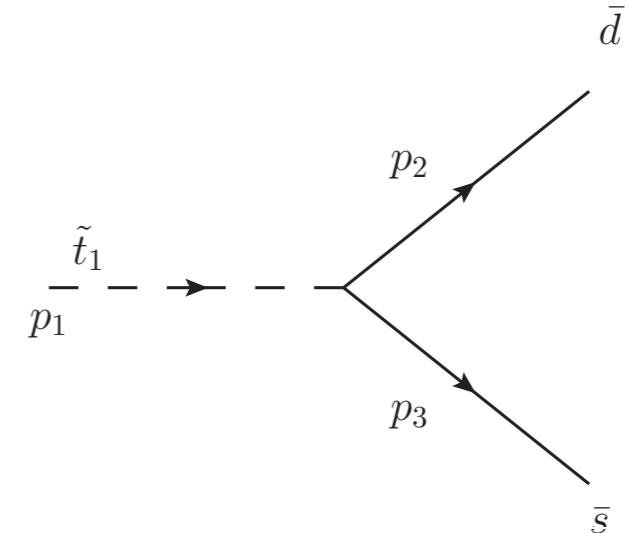
$$\sigma_{\tilde{u}_i^*} = \frac{2\pi}{3m_i^2} \sum_{jk} \sum_{i'} |\lambda''_{i'jk} (R^u)_{ii'}|^2$$

Sparticle Decays

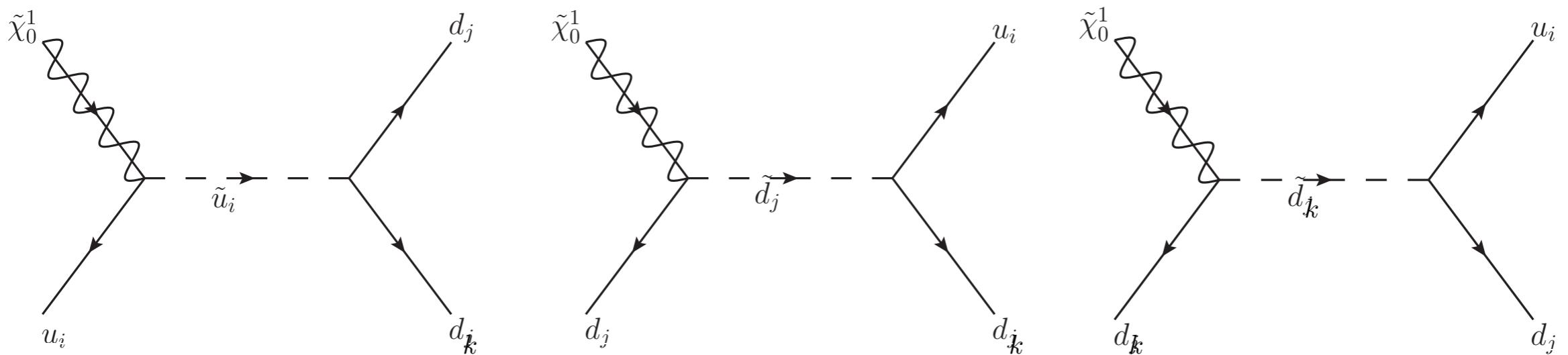
- $\tilde{g} \rightarrow \tilde{q}_i q_j$
- $\tilde{\chi}_i^0 \rightarrow \tilde{q}_i q_j, \tilde{l}_i l_j, \tilde{\chi}_j^0 Z, \tilde{\chi}_j^+ W^-, \tilde{\chi}_j^0 H_k, \tilde{\chi}_j^+ H^-$
- $\tilde{\chi}_i^+ \rightarrow \tilde{q}_i q_j, \tilde{l}_i l_j, \tilde{\chi}_j^+ Z, \tilde{\chi}_j^0 W^+, \tilde{\chi}_j^+ H_k, \tilde{\chi}_j^0 H^+$
- $\tilde{q}_i \rightarrow q_j \tilde{\chi}_k^0, q_j \tilde{\chi}_k^+, \tilde{q}_j Z, \tilde{q}_j W^+, \tilde{q}_j H_k, \tilde{q}_j' H^+$

Decays via UDD

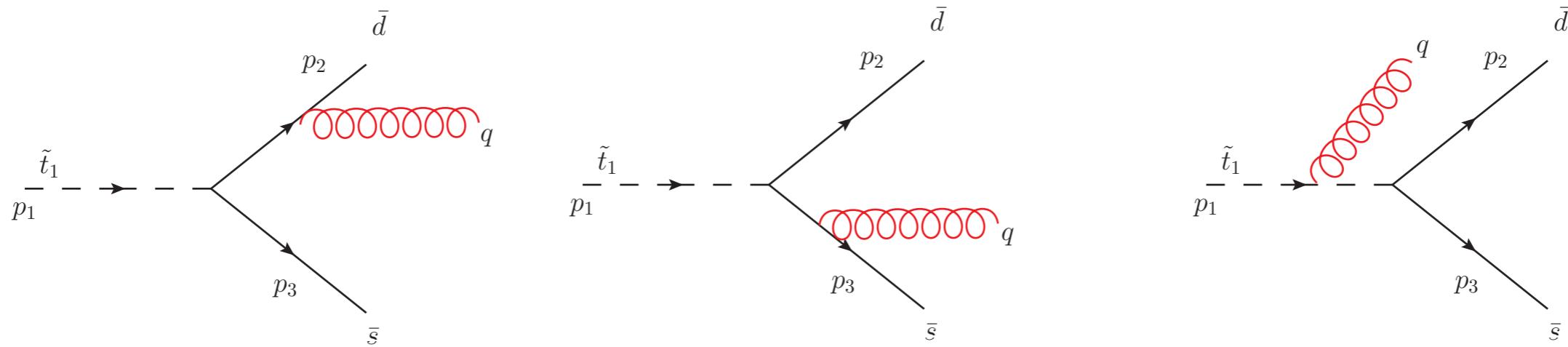
Squark decays: $\tilde{q}_i \rightarrow \bar{q}_j \bar{q}_k$



Neutralino Decays: $\tilde{\chi}_i^0 \rightarrow u_i d_j d_k$



Showering in the presence of baryon-number violation

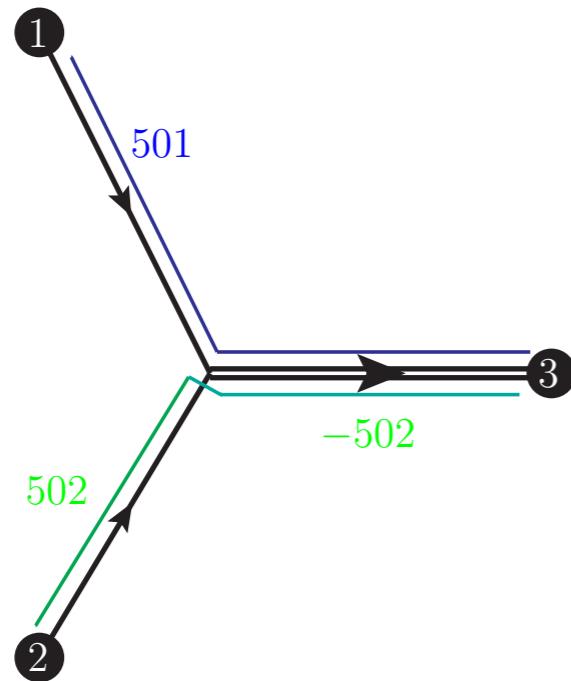


$$\frac{|M_1|^2}{|M_0|^2} = 4\pi\alpha_s C_F \left[\frac{1}{(N_c - 1)} \left(\frac{2s_{23}}{s_{2q}s_{3q}} + \frac{2s_{12}}{s_{1q}s_{2q}} + \frac{2s_{13}}{s_{1q}s_{3q}} \right) + \frac{s_{2q}}{s s_{3q}} + \frac{s_{3q}}{s s_{2q}} \right]$$

+ finite terms

Case of $\tilde{\chi}_i^0 \rightarrow u_i d_j d_k$ is similar, with **three half-strength dipoles** between the quarks

Other exotic colour structures: Sextets



Needs two color indices

We use a negative anti-colour index to denote the extra colour

Implementing completely new BSM models

Method 1: Use QNUMBERS and read in LHEF events

```
SLHA:file = fileName

BLOCK QNUMBERS 8765432 # yup yupbar
    1    2 # 3 times electric charge
    2    2 # number of spin states (2S+1)
    3    3 # colour rep (1: singlet, 3: triplet, 6: sextet, 8: octet)
    4    1 # Particle/Antiparticle distinction (0=own anti)

BLOCK MASS
#      ID code  pole mass in GeV
8765432  600.0 # m(yup)
```

Use LHEF format to pass events to Pythia8

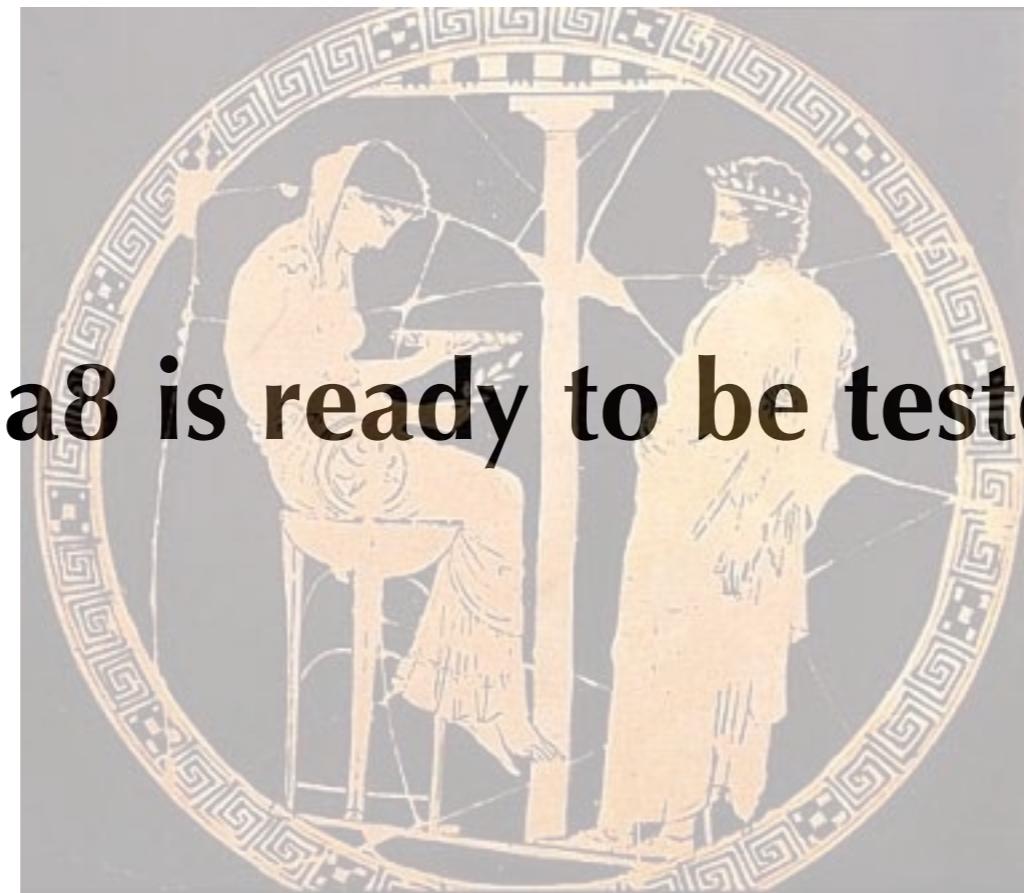
Method 2: Use Pythia's semi-internal process machinery

Write the process as an inherited class of Pythia's `SigmaProcess`

Use the SLHA file to pass any user-defined blocks. This can be used to pass couplings etc. for a new model which can be accessed from within the process using

```
bool slhaPtr->getEntry(string blockName, double& val);
bool slhaPtr->getEntry(string blockName, int idx, double& val);
bool slhaPtr->getEntry(string blockName, int idx, int jndx, double& val);
bool slhaPtr->getEntry(string blockName, int idx, int jndx, int kndx, double& val);
```

SUSY in Pythia8 is ready to be tested in the field!

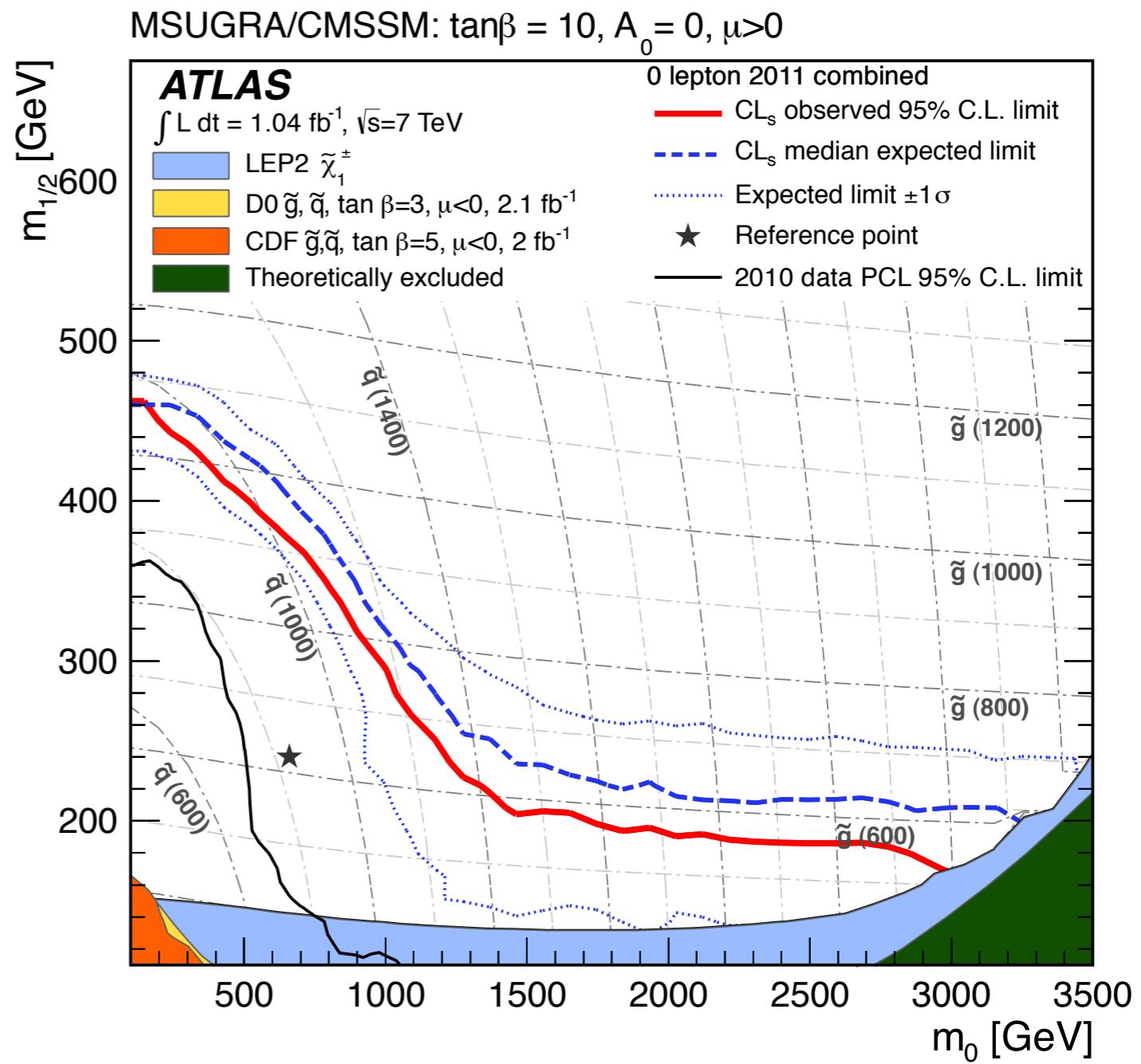


Work done with Peter Skands [arXiv:1109.5852]

Part II: Interpretation of ATLAS SUSY limits for third-generation squarks

ATLAS limits with 1 fb^{-1} (jets+MET) and 0.833 fb^{-1} (bjets+MET)

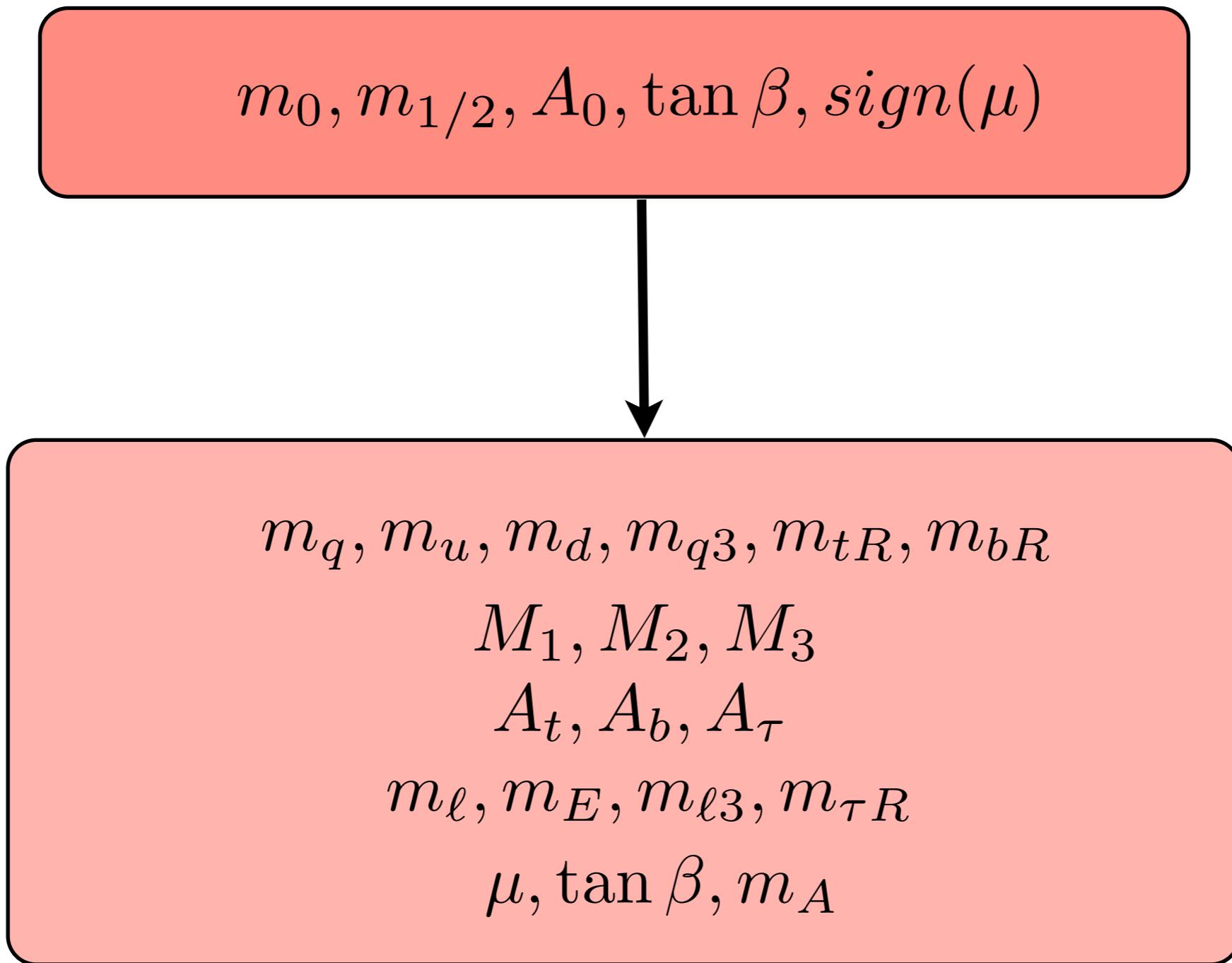
Channel	$\sigma \times acc (\text{fb})$
2 jets + MET	24
3 jets + MET	30
4 jets + MET ($M_{eff} = 1\text{ TeV}$)	32
1 btag + $M_{eff} > 500$ (3JA)	288
1 btag + $M_{eff} > 700$ (3JB)	61
2 btag + $M_{eff} > 500$ (3JC)	78
2 btag + $M_{eff} > 700$ (3JD)	17



Why separate limits for third generation squarks?

- Motivation for TeV-scale SUSY: hierarchy problem
i.e. \tilde{t}_1 mass should be ~ 500 GeV or less.
- Strongest limits currently from jets+MET signals
which are tailored to detect squarks of first two
generations
- The limits on third generation squarks are
indirect.
- We are already on the brink of un-naturalness if
we use limits from jets+MET searches.

Breaking free (somewhat) from the cMSSM



Direct limits on third generation squarks

Consider the case where the **first two generations of squarks and all sleptons are heavy** enough that they cannot be probed at the 7 TeV run of the LHC.

Only third generation squarks are accessible.

$$m_q, m_u, m_d, m_\ell, m_e, m_{\ell 3}, m_{\tau R} \sim 2 \text{ TeV}$$

We also retain the cMSSM gaugino mass pattern

$$M_1 : M_2 : M_3 \simeq 1 : 2 : 6$$

(hence the “somewhat” breaking free)

For parametrising the third generation, consider the stop mass matrix

$$\begin{pmatrix} M_{\tilde{t}_1} & 0 \\ 0 & M_{\tilde{t}_2} \end{pmatrix} = \mathcal{R} \begin{pmatrix} m_{q3} & m_t X_t \\ m_t X_t & m_{tR} \end{pmatrix} \mathcal{R}^{-1}; \quad \mathcal{R} = \begin{pmatrix} \cos \theta_{\tilde{t}} & \sin \theta_{\tilde{t}} \\ -\sin \theta_{\tilde{t}} & \cos \theta_{\tilde{t}} \end{pmatrix}$$

$$X_t = A_t - \mu \cot \beta$$

For fixed values of μ and $\tan \beta$, the independent parameters are m_{q3}, m_{tR}, A_t

Equivalently, we can invert the equation to have

$$M_{\tilde{t}_1}, M_{\tilde{t}_2}, \sin \theta_{\tilde{t}}$$

as the parameters of the scan

Of course, one has to be careful while applying the same inversion to the sbottom sector because it shares the parameter m_{q3}

Case A: \tilde{t}_1 lighter than other third generation squarks (closest to the cMSSM case)

$$\tilde{t}_1 \sim \tilde{t}_R (\sin \theta_{\tilde{t}} = 0.99)$$

Step 1: Choose $M_{\tilde{t}_1}$

Step 2: Set $M_{\tilde{t}_2} = M_{\tilde{t}_1} + 500$ GeV

Step 3: Calculate m_{q3}, m_{tR}, A_t

Step 4: Set $m_{bR} = m_{q3} (\sin \theta_{\tilde{b}} = 0.707)$

We can now scan over a $M_{\tilde{t}_1} - M_2$ plane for fixed values of $\mu, \tan \beta, m_A$

What does Case A probe?

Dominant production processes: $\tilde{g}\tilde{g}, \tilde{t}_1\tilde{t}_1^*$

$$\begin{aligned} M_2 > 150 \text{ GeV} &\Rightarrow M_{\tilde{g}} > 450 \text{ GeV} \\ &\Rightarrow \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_i^0 \end{aligned}$$

$$M_{\tilde{t}_1} < M_{\tilde{\chi}_1^0} \Rightarrow \text{Stop LSP!}$$

$$M_{\tilde{\chi}_1^0} < M_{\tilde{t}_1} < M_{\tilde{\chi}_1^0} + M_t \Rightarrow \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$$

$$M_{\tilde{t}_1} > M_{\tilde{\chi}_1^0} + M_t \Rightarrow \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$$

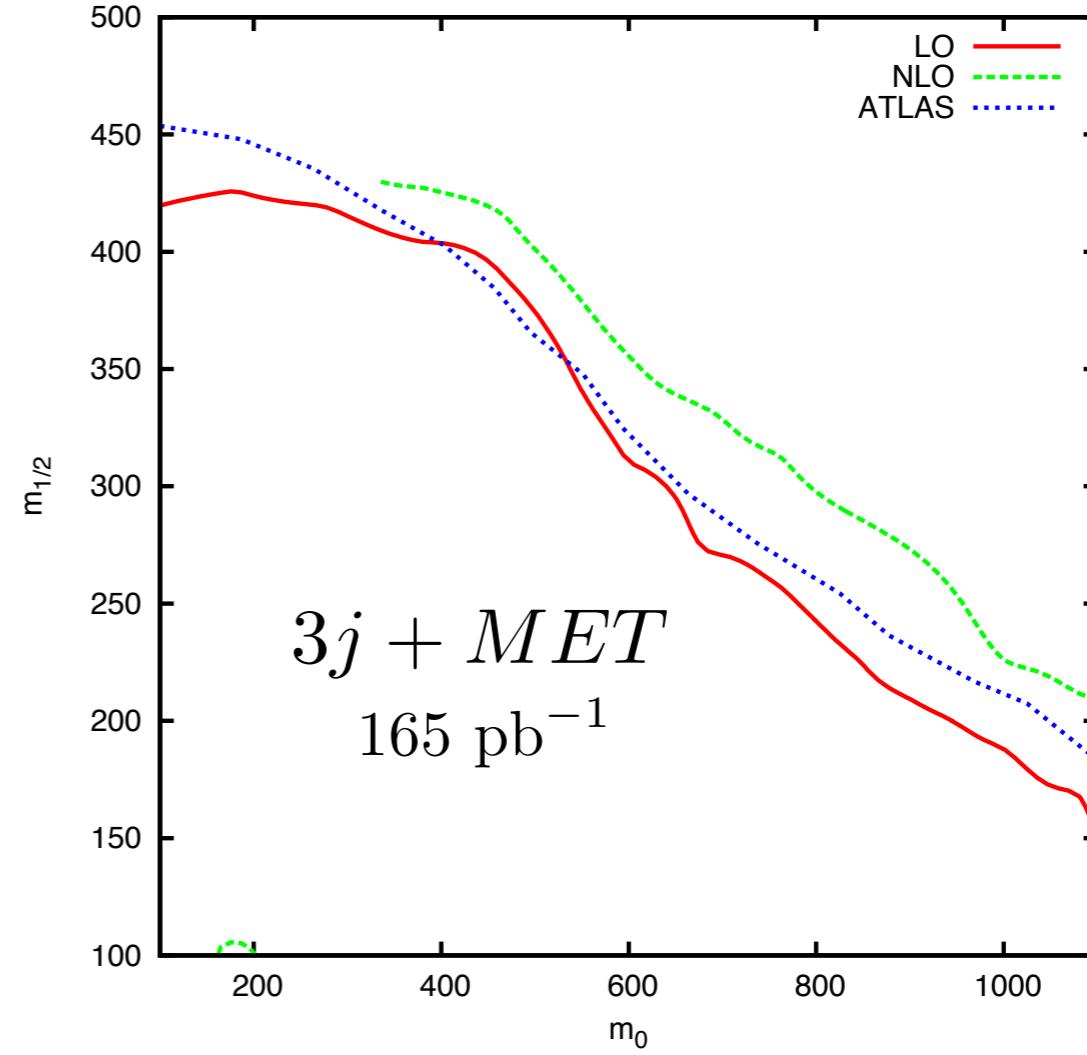
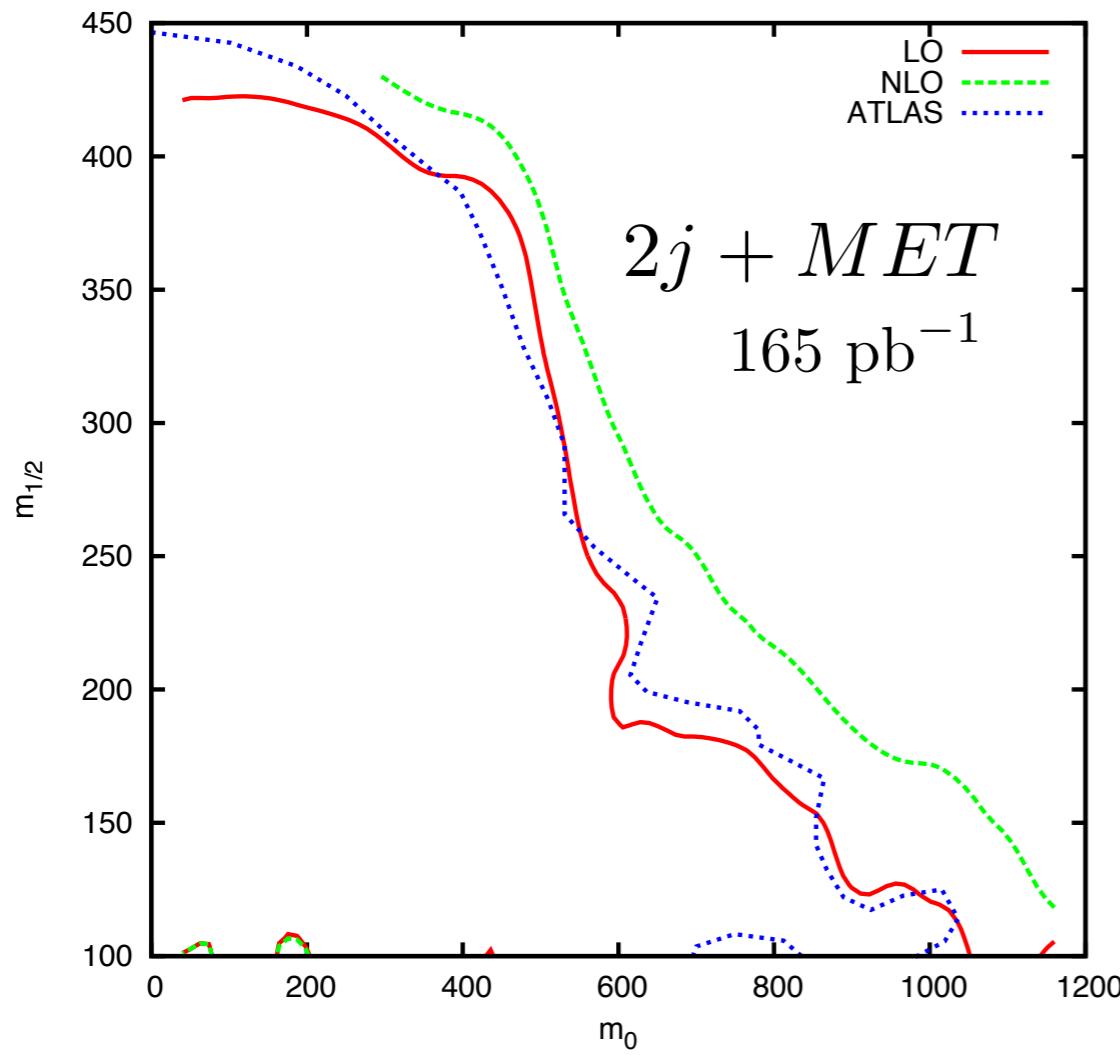
$$M_{\tilde{t}_1} > M_{\tilde{\chi}_1^+} + M_b \Rightarrow \tilde{t}_1 \rightarrow b\tilde{\chi}_1^+$$

Simulation of Signal

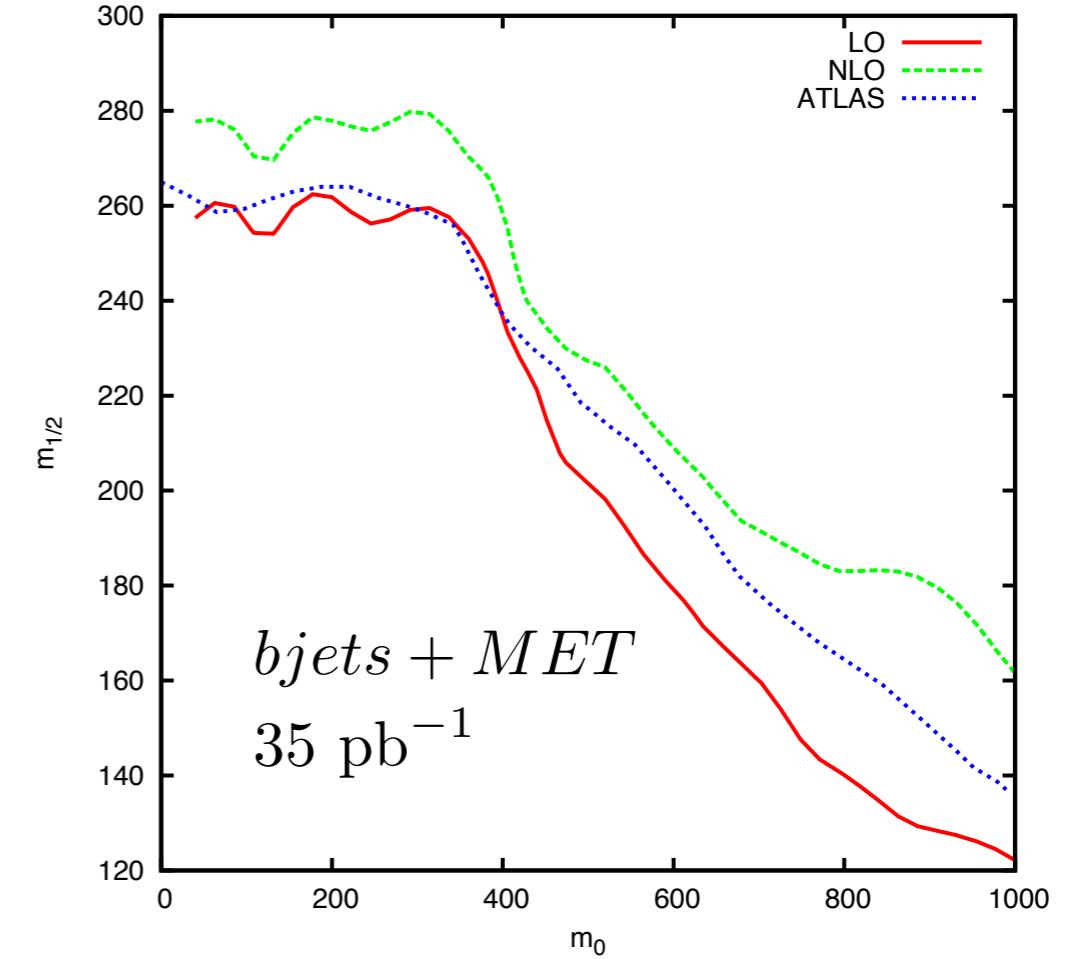
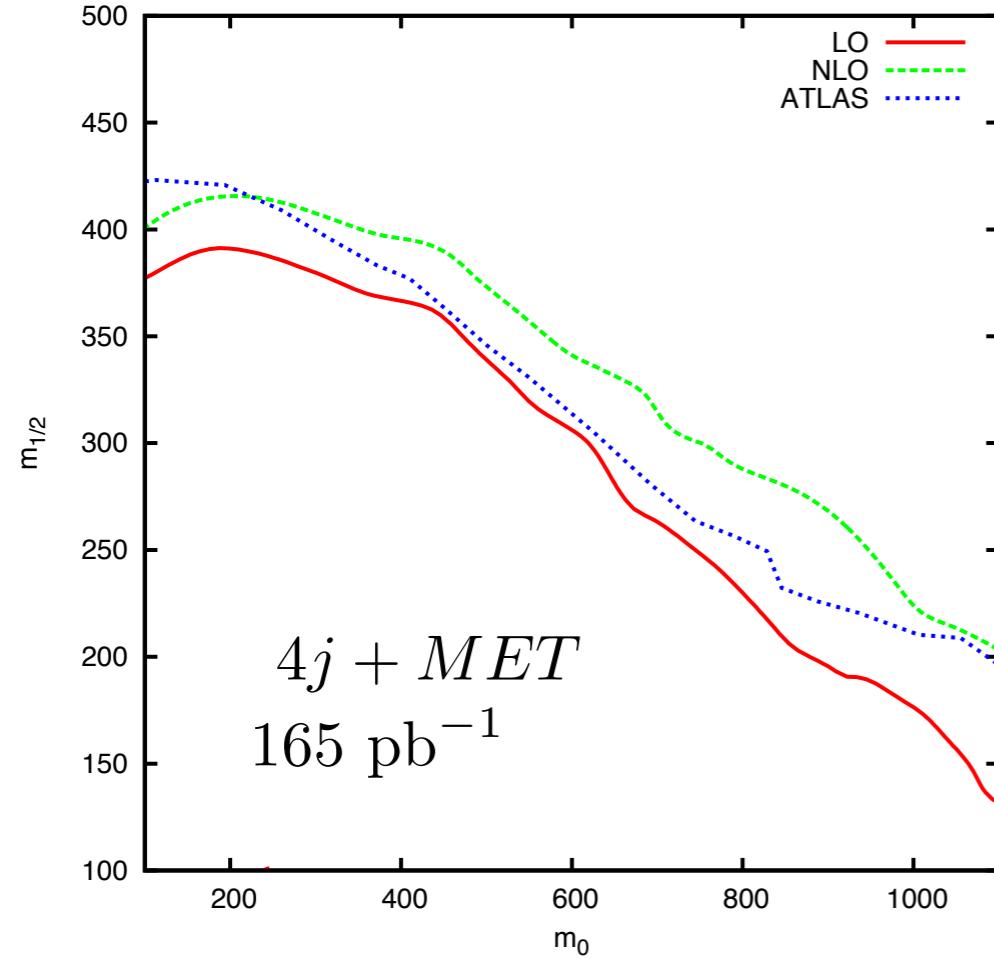
- Generate spectrum using SUSPect
- Use Pythia 6.4 for signal, Prospino for NLO normalisation
- Anti-kt algorithm for forming jets (FastJet)
- Smear momenta of all objects
- Apply cuts as mentioned in the paper
(jets+MET: [ATLAS-CONF-2011-086; arXiv:1109.6572];
bjets+MET: ATLAS-CONF-2011-098)

Ask how well do we reproduce ATLAS's exclusion curve?

Comparison to ATLAS contours



Comparison to ATLAS contours



LO (and NLO) contours agree within 20% in all cases!

Signal x Acceptance used for contours

Channel	$\sigma \times acc$ (fb)
2 jets + MET	24
3 jets + MET	30
4 jets + MET ($M_{eff} = 1$ TeV)	32
1 btag + $M_{eff} > 500$ (3JA)	288
1 btag + $M_{eff} > 700$ (3JB)	61
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Case A: \tilde{t}_1 lighter than other third generation squarks (closest to the cMSSM case)

$$\tilde{t}_1 \sim \tilde{t}_R (\sin \theta_{\tilde{t}} = 0.99)$$

Step 1: Choose $M_{\tilde{t}_1}$

Step 2: Set $M_{\tilde{t}_2} = M_{\tilde{t}_1} + 500$ GeV

Step 3: Calculate m_{q3}, m_{tR}, A_t

Step 4: Set $m_{bR} = m_{q3} (\sin \theta_{\tilde{b}} = 0.707)$

Case B: $\tilde{t}_1 \sim \tilde{t}_L$; $M_{\tilde{t}_1} \sim M_{\tilde{b}_1}$

$$\sin \theta_{\tilde{t}} = \sin \theta_{\tilde{b}} = 0.1$$

Step 1: Choose $M_{\tilde{t}_1}$

Step 2: Set $M_{\tilde{t}_2} = M_{\tilde{t}_1} + 500$ GeV

Step 3: Calculate m_{q3}, m_{tR}, A_t

Step 4: Set $m_{bR} = m_{tR}$

Degenerate $\tilde{t}_1, \tilde{b}_1; \tilde{t}_2, \tilde{b}_2$

This case is interesting because of the enhanced squark
couplings to the wino-like neutralino.

Case C: $\tilde{b}_1 \sim \tilde{b}_R$; $M_{\tilde{t}_1} = M_{\tilde{t}_2} = M_{\tilde{b}_2}$

This case is never realised in cMSSM-type high scale models but is closest to the model considered in the paper

Step 1: Choose $M_{\tilde{b}_1}$

Step 2: Set $M_{\tilde{b}_2} = M_{\tilde{b}_1} + 500$ GeV

Step 3: Calculate m_{q3}, m_{bR}, A_b

Step 4: Set $m_{tR} = m_{q3}$; $\sin \theta_{\tilde{t}} = 0.707$

Gluino decays via: $\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$

Sbottom decays via: $\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$

Case D: All third generation squarks nearly degenerate

$$\sin \theta_{\tilde{t}} = \sin \theta_{\tilde{b}} = 0.707$$

Step 1: Choose $M_{\tilde{t}_1}$

Step 2: Set $M_{\tilde{t}_1} = M_{\tilde{t}_2}$

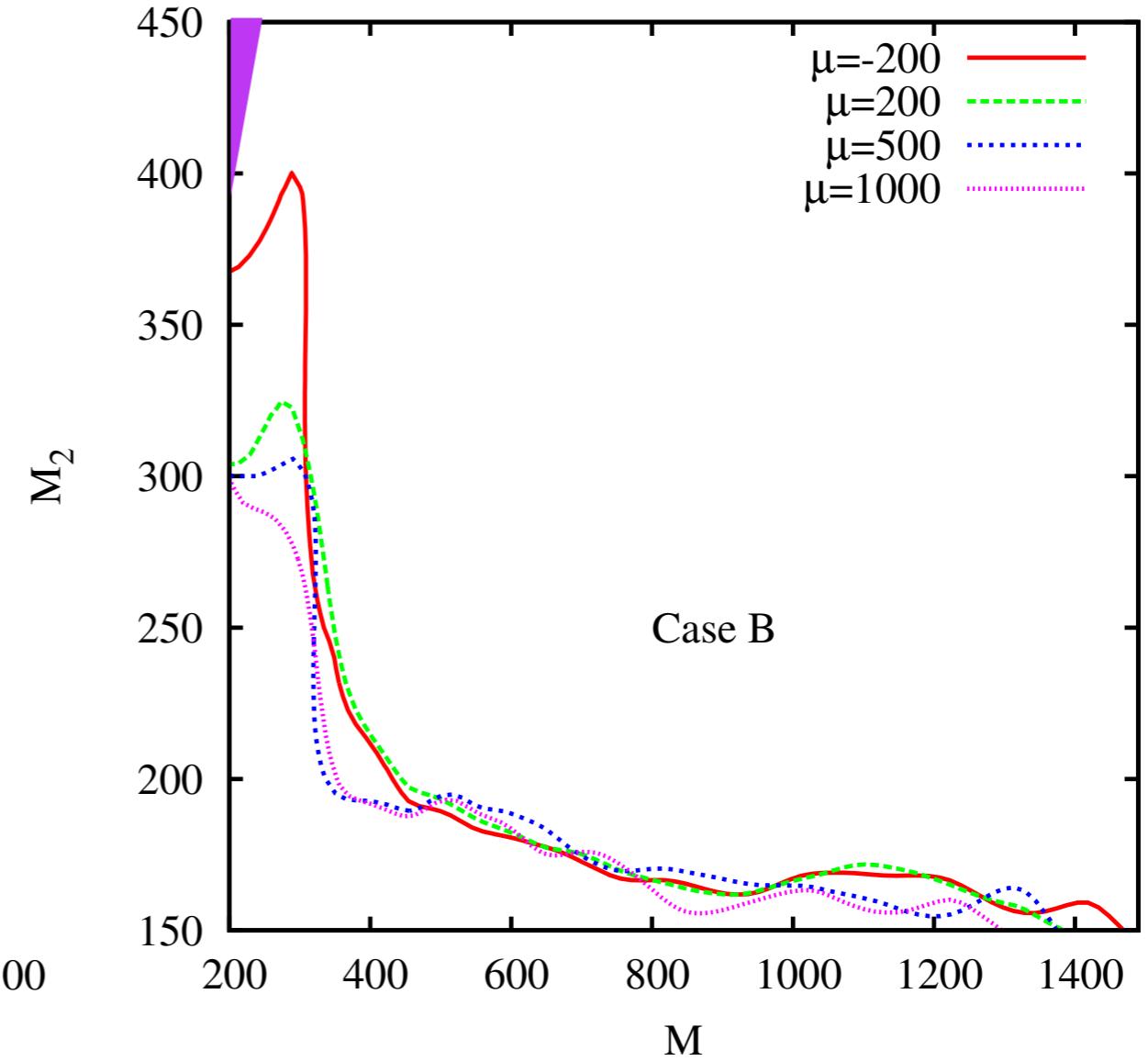
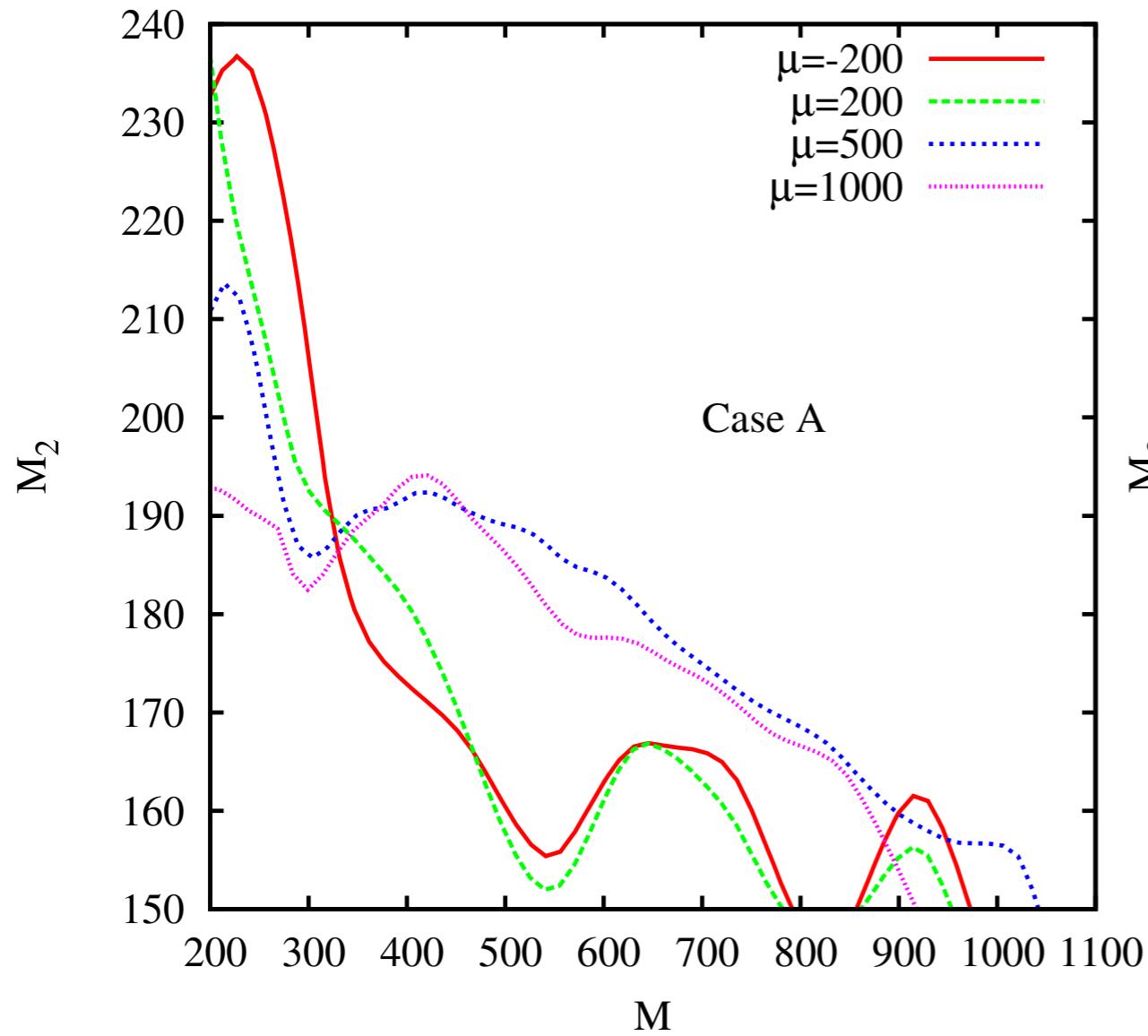
Step 3: Calculate m_{q3}, m_{tR}, A_t

Step 4: Set $m_{bR} = m_{tR}$

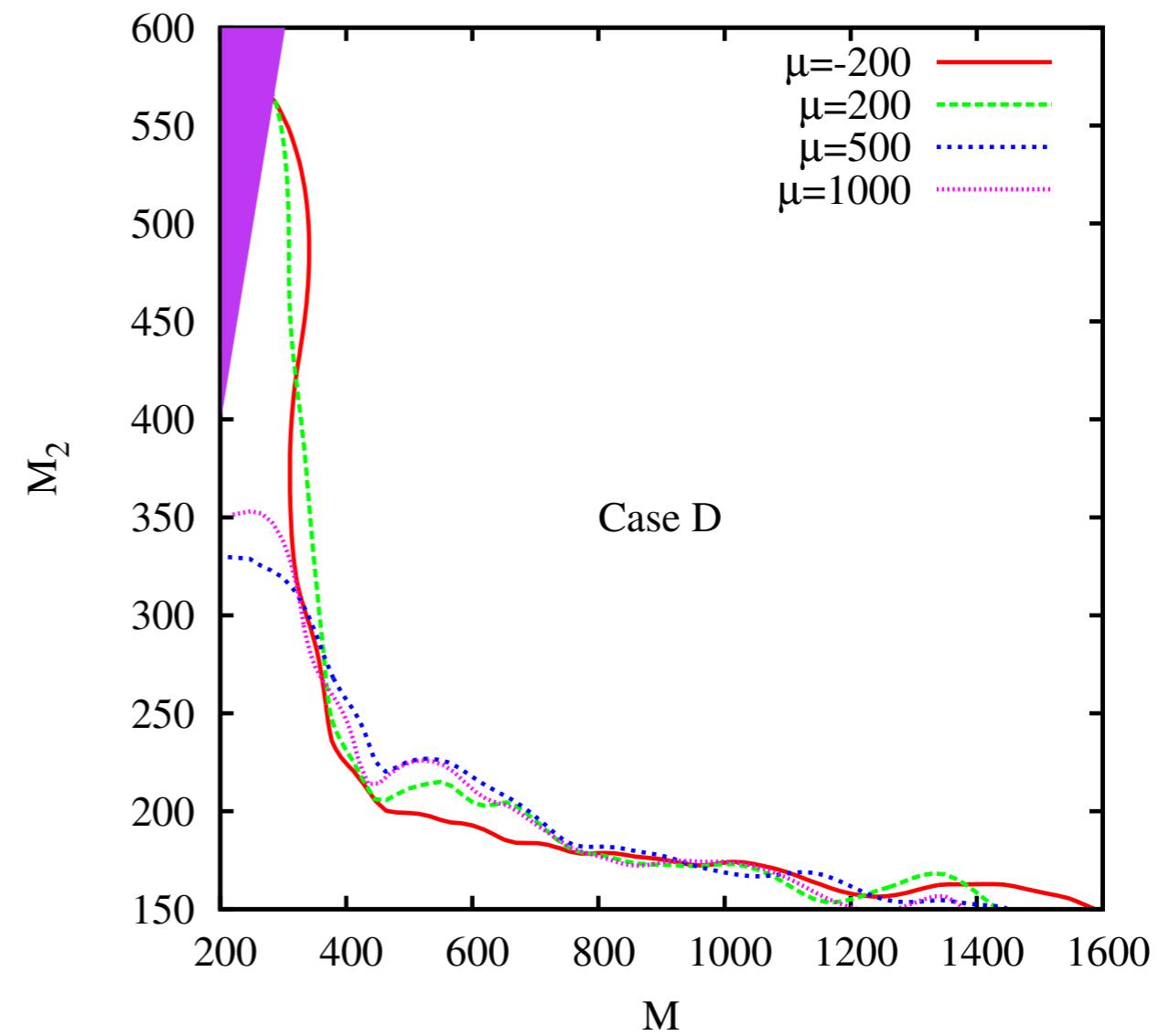
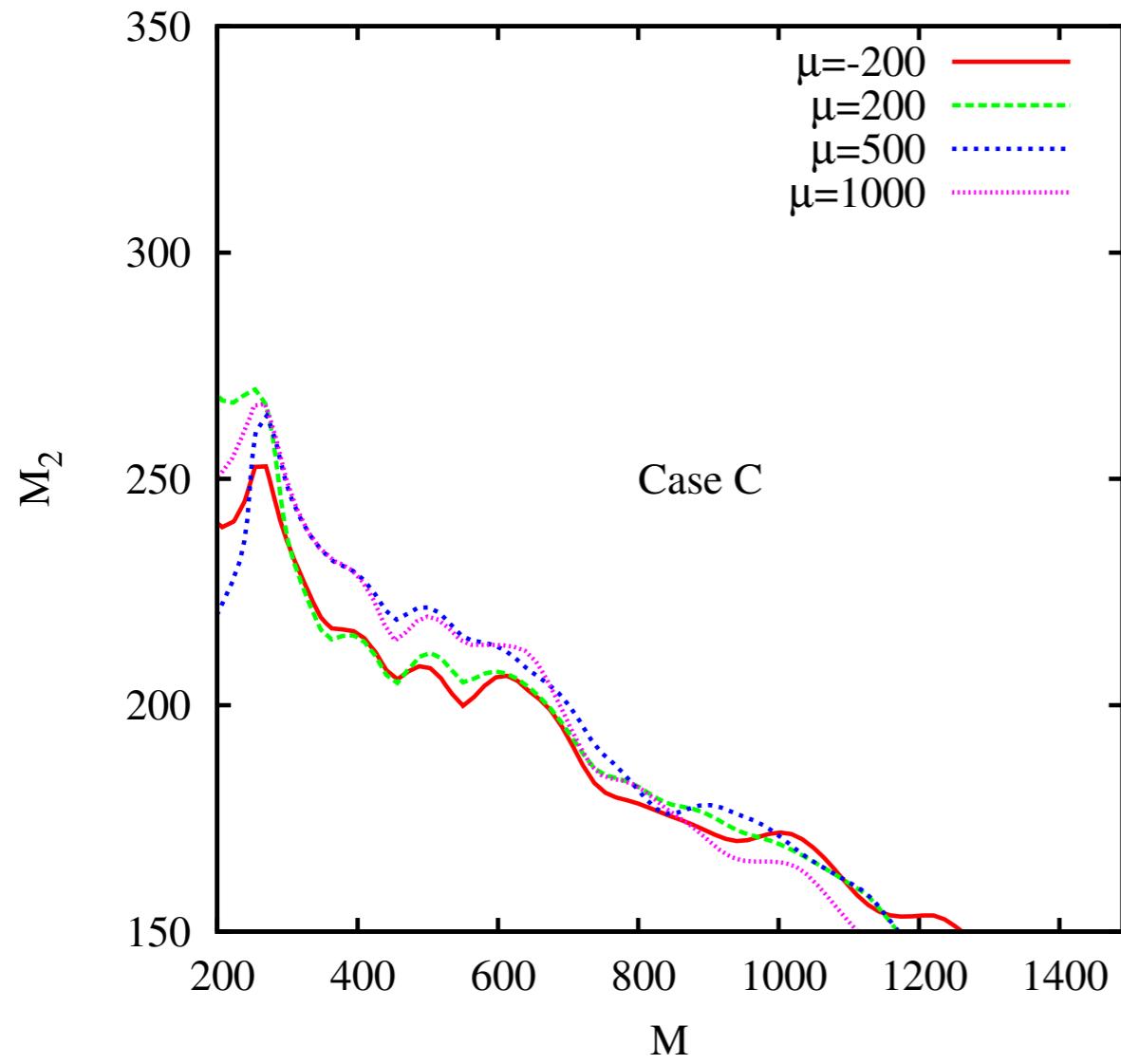
We scan over the following parameters

Parameter	Scan range
$M_{\tilde{t}_1}$	100 - 2000 GeV
M_2	150 - 600 GeV
$\tan \beta$	5, 10, 40
μ	-200, 200, 500, 1000 GeV

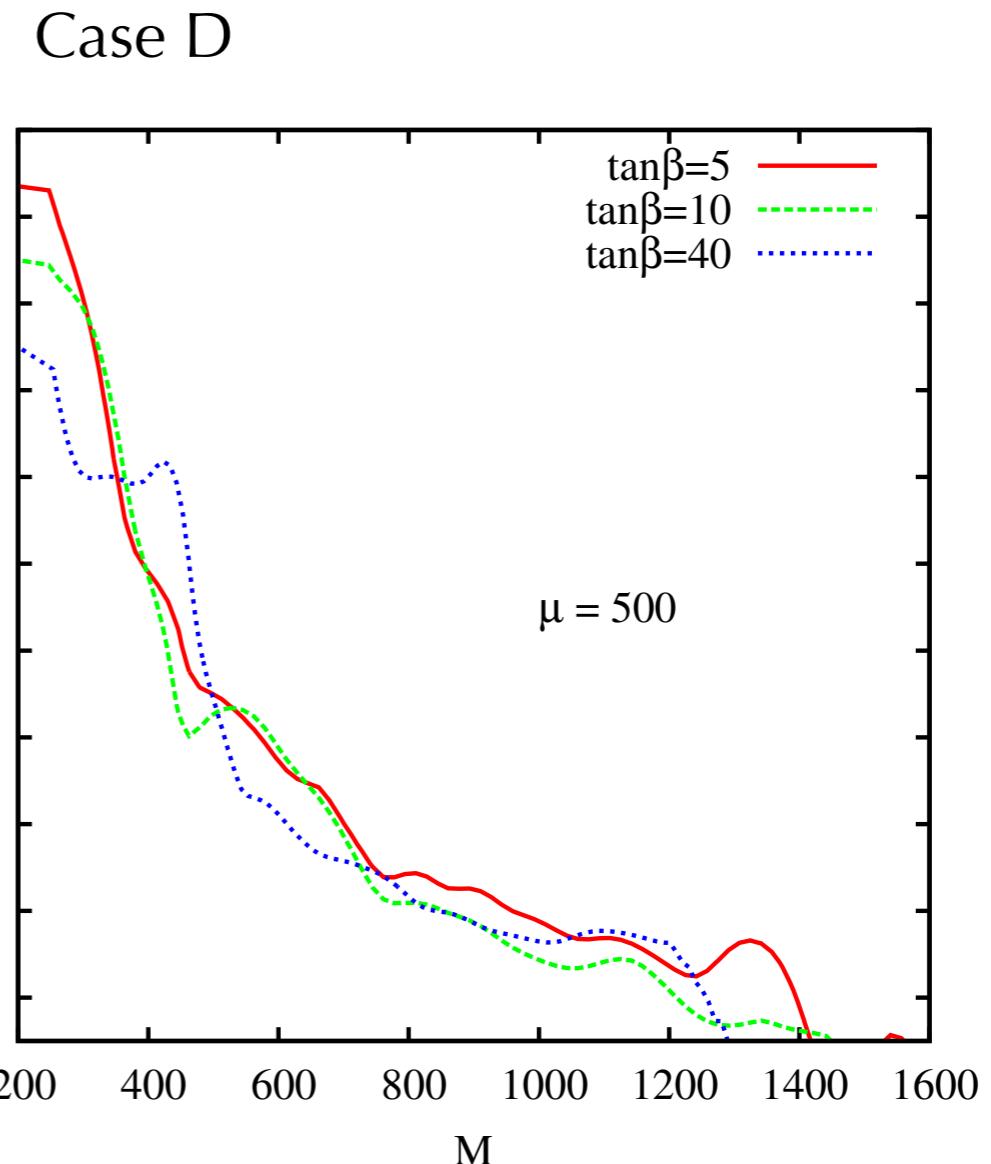
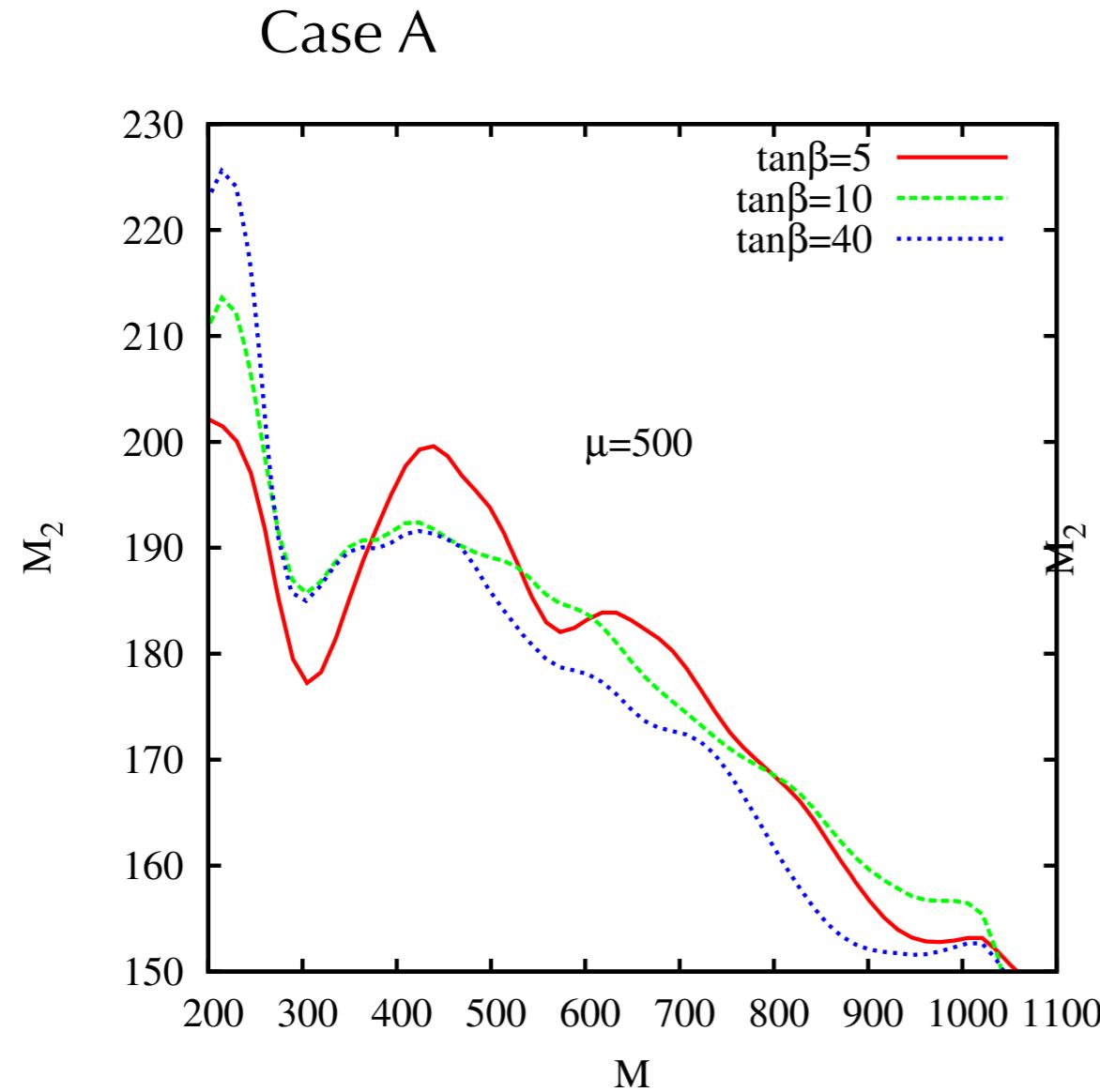
Effect of μ



Effect of μ



Effect of $\tan \beta$



High-Scale Non-Universality

We now address the question of what happens if sleptons are also possibly light

Two scales in the scalar sector: $M_{\text{heavy}}, M_{\text{light}}$

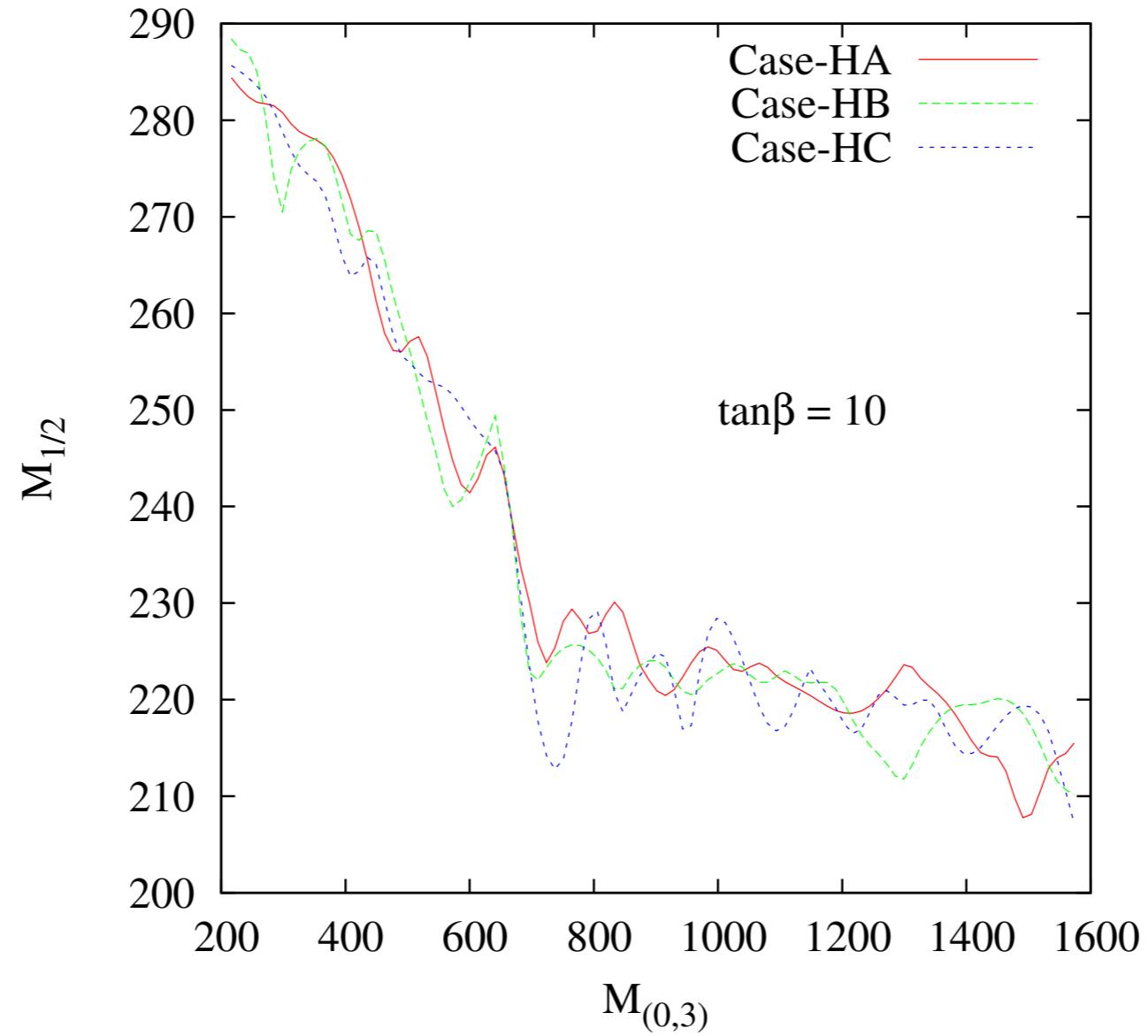
We still want the first two generations of squarks to be heavy

Case HA: All sleptons heavy

Case HB: First two generations of sleptons heavy

Case HC: All sleptons light

High-Scale Non-Universality



Conclusions

If one considers the limit on gluino mass to be 600 (700) GeV, that translates to a limit of

$$M_{\tilde{t}_1} > 300 - 350 \text{ (300) GeV (Case A)}$$

$$M_{\tilde{t}_1, \tilde{b}_1} > 340 - 450 \text{ (300 - 400) GeV (Case B)}$$

$$M_{\tilde{b}_1} > 750 \text{ (350 - 450) GeV (Case C)}$$

$$M_{\tilde{t}_i, \tilde{b}_i} > 450 - 700 \text{ (400 - 500) GeV (Case D)}$$