Dark Matter: a debate Cold/Warm Dark Matter is ruled out

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Bethe Colloquium, 18th November 2010

A historical perspective : 500 years ago learning from experience Pavel Kroupa: AIfA, University of Bonn

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Step II : Making the theory fit

Add epicycles to achieve high precision (Claudius Ptolemaeus in the 2nd century AD)

==> excellent description of the data.

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Assuming Einsteinian / Newtonian gravity (1915) to be valid

The Standard / Concordance Cosmological Model

can only be made consistent with the data by adding

inflation (1965 and 1980)

cold dark matter (around 1980)

dark energy (1999)

dark force (2010)

The Standard / Concordance Cosmological Model

Problems

dark energy: 70 %	the implied dark energy density is so small that it is unstable to quantum correction (Dvali et al. 2002); not seen by WMAP (Shanks); energy creation; may not be there at all (Wiltshire)
dark matter: 25 %	despite much search hitherto unknown stuff
baryons: 5 %	only 40% of these found - the missing baryon problem
dark force :	totally unknown (Peebles & Nusser 2010; Kroupa et al. 2010)
	Pavel Kroupa: AlfA, University of Bonn

That is, we are trying to describe / model the universe with essentially unknown physics.

This is like trying to construct stellar models based to 95% on completely unknown ingredients.

MOND and structure formation and galaxies

Combes & Tiret 2009, arXiv:0908.3289 :

"Their simulations yield comparable results between MOND and the standard model for the large-scale structure, with even more clustering than with the parallel approximation. "

MOND and structure formation and galaxies

Skordis et al. 2006, PhRvL

"We show that it may be possible to reproduce observations of the cosmic microwave background and galaxy distributions with Bekenstein's theory of MOND."

Angus 2010 on SciLogs: "State-of-the-art cosmology: the current status" Perfect fit to all CMB peaks with 11ev sterile neutrino.

One of the *primary arguments* for cold or warm dark matter comes from flat rotation curves . . .

But, flat rotation curves are explained much better by non-Einsteinian / Newtonian gravity:

For example :

Modified Newtonian Dynamics (MOND)

Milgrom 1983

Modified Gravity (MOG)

Moffat 2005

In fact, given an observed baryonic matter distribution, the rotation curve

can be precisely predicted using MOND

cannot be predicted using LCDM.

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can be precisely predicted using MOND

cannot be predicted using LCDM.

plus in MOND dark matter significantly reduced in galaxy clusters

 (e.g. Sanders 2009 (review):
"Modified Newtonian Dynamics: A Falsification of Cold Dark Matter")

In fact, galaxies are MONDian objects.

i.e., MOND is the correct dynamical description for galaxies.

two hypotheses

1982 Suggestion of massive, weakly interacting dark matter particles and their role in structure formation

(I)

Cold/Warm Dark Matter

-Bond, Szalay & Turner 1982, Phys. Rev. Lett. -Blumenthal, Pagels & Primack 1982, Nature -Peebles 1982, ApJL -Blumenthal, Faber, Primack, Rees, 1984, Nature

(II)

1983 Suggestion of a modification of Newton's force law

> MOND MOG

-Milgrom 1983, ApJ -Moffat 2005, JCAP

Progress is always linked to cultural pre-disposition and Sociology → Fanelli D. (2010) Do Pressures to Publish Increase Scientists' Bias? An Empirical Support from US States Data. PLoS ONE 5(4): e10271. doi:10.1371/journal.pone.0010271 Pavel Kroupa: AIfA, University of Bonn

Today and in the past young researchers are

1. afraid

2. discouraged to try alternatives

Some personal examples - statements by *well-known* and *very influential* scientists :

"Pavel, in 1997 you have written that paper on *dSph satellites without dark matter* - you are *unhirable*." (about 2004)

"It is not worth reading those papers on satellite galaxies by Pavel Kroupa." (2009)

"But everyone knows that MOND is crap!" (at STScI May, 2010)

"*I would be scared* (mostly because I am still in search of a permanent post) of being labelled once and for all as a "hardcore-MONDian" person." (July 2010)

... as if being labeled a "hardcore-LCDM" person were acceptable ... (my own note added Nov. 2010).

"I can't do any MOND work - the director would not appreciate it" (Garching August, 2010).

-Bond, Szalay & Turner 1982, Phys. Rev. Lett. -Blumenthal, Pagels & Primack 1982, Nature -Peebles 1982, ApJL -Blumenthal, Faber, Primack, Rees, 1984, Nature

Step IV

Decision made possible through technological advance

Assume the standard cosmological model (i.e. Newtonian gravity) is valid.

Two Zwicky Conjectures of fundamental importance :

1. Zwicky (1937): There must be dark matter.

2. Zwicky (1956): *Tidal-dwarf galaxies form* out of the collisional debris of other galaxies.

Both have to be true

This has two immediate implications :

1. There exist large numbers of *dark-matter dominated satellite galaxies*.

2. There exist large numbers of *newly formed (tidal-dwarf) satellite galaxies* (they do not contain dark matter).

This is OK, but are there two different types of dwarf galaxy?

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1. There exist large numbers of dark-matter dominated satellite galaxies

2. There exist large numbers of newly formed (tidal-dwarf) satellite galaxies.

NO, there is only one type of dwarf galaxy !

But, which one?

And *why* only one?

Lets consider first the dark-matter type satellite dwarf galaxy:

A vast amount of theoretical research has been done by countless LCDM research groups with the aim of explaining the observed satellite galaxy population naturally in LCDM.

... they all claim success

The DM mass-luminosity relation :

more gravitating mass, more luminous mass ?







The *lack* of an *observed* mass-luminosity relation $(\kappa \approx 0)$



nature apparently does not care about the existence of the putative dark matter halo.

Thus, the *concept* of dark-matter halos appears to be *unphysical* for dSph satellites



Individual dSph morphology:

DM gravitating potential : smooth luminous morphology ?

Significant isophote structure is present in many dSph satellites despite a large $\sigma \approx 700 \text{ pc}/100 \text{ Myr}$ (=7 km/s) Substructure should phase-mix away if σ is really due to a DM halo, unless it has a harmonic core. consistent with DM halo? Pavel Kroupa: AIfA, University of Bonn



UMi D=65kpc

(Martinez-Delgado et al., in prep)

S shape : strong evidence for extra-tidal stars

Massive CDM halo ?



The distortions apparent in many of the dSph satellites do not support the notion that they are shielded by $10^9 M_{\odot}$ dark-matter halos.



The spatial distribution of the MW satellites

... further clues





















The MW satellite *DoS* is defined mostly by the *outer satellites*; but the *angular momenta* of the *inner satellites* are aligned to the pole of the DoS



the satellites form a highly correlated phase-space population

This correlated phase-space population is *inconsistent* with the satellites being dark-matter sub-haloes that fell into the MW halo *in a group* or *individually*.



Within about 2 Mpc (Local Group):

Combined likelihood that the Λ CDM model accounts for the observed Local Group < 0.056 %.

(Kroupa et al. 2010)

Based on nearly 10⁹ particle - LCDM simulation. (Libeskind et al. 2009)

Within about 8 Mpc (Local Volume) :

Likelihood that the Λ CDM model accounts for the 3 massive galaxies above Local Sheet P < 1 %and far too few galaxies in the void ($P \sim 10^{-3} \%$).

(Peebles & Nusser 2010, Nature)

And, *disk galaxies far too similar*. (Peebles & Nusser 2010 *citing* Disney et al. 2008, Nature)



The dark-matter ansatz fails.

Lets consider now the *tidal-dwarf galaxy* (TDG):

(Zwicky's 2nd conjecture)



Relevance: The collision of two disks at high redshift



Wetzstein, Naab & Burkert 2007





Thus, by direct observation *new dwarf galaxies* with masses comparable to dE/dSph galaxies form.

They are baryon dominated (Barnes & Hernquist 1992).

Evolution of

TDGs


Evolve dwarf galaxies w/o dark matter in a computer









For TDGs we know today that

The early (<100Myr) star-formation and chemical enrichment evolution is similar to the observed dSph satellites. (Recchi et al. 2007)

Later dynamical evolution does not destroy the satellites. (Kroupa 1997)

The number of old TDGs amounts to the dE population observed.

(Okazaki & Taniguchi 2000)

dE galaxies are observed to contain *no Dark Matter*, consistent with them being *TDGs*.

(Toloba et al. 2010, arXiv:1011.2198v1)

Young (age<100Myr) TDGs have rotation curves showing the *missing mass syndrome*. But they cannot have DM !

(Gentile et al. 2007)

=> breakdown of Newtonian/Einsteinian dynamics

This appears to be direct evidence that the LCDM model is not realistic.



... and, the bulge mass vs number of satellites correlation ?





Both, the *Disk of Satellites* and the *bulge--satellite correlation* are easily understandable if the MW satellites are *ancient TDGs*.

(Kroupa et al. 2010)



(Weilbacher et al. 2000)

Phase-space correlated satellites form naturally in the same event as a *bulge* does.

Both, the Disk of Satellites and the bulge--satellite correlation follow trivially if the MW satellites are ancient TDGs.

And, the *DM mass--luminosity (non)correlation* $(\kappa \approx 0)$: there is no dark matter.

And, *TDGs are known to form*. They are the result of well understood standard-physical processes.

(Kroupa et al. 2010)

Thus,

(1) a *fully self-consistent TDG scenario* thus emerges which very naturally accounts for the properties of dE and satellite galaxies;

(2) no consistent, and in fact
a contradictory picture
emerges in the dark-matter framework;

(3) there is simply *no evidence for* the existence of *DM satellites*.

Actually,

every prediction of the standard cosmological model *failed*.

Some failures can be "resolved" by introducing unknown physics (Inflation, Dark Matter, Dark Energy) but on galaxy scales and beyond, failure remains the rule...

e.g. • the disk-halo conspiracy remains unsolved;

- invariant disk galaxies;
- preponderance of disk galaxies without bulges;
- the likelihood of having a Local Group is $< 5.6 \times 10^{-5}$. Pavel Kroupa: AIFA, University of Bonn

A comparison (galactic astrophysics)

Problem at hand	Standard Model	MOND
irregular dSph morphology	X	\checkmark
dSph dmass - luminosity relation (energy conservation)	X	\checkmark
phase-space correlation (Disk of Satellites)	X	\checkmark
bulge-mass vs satellite number correlation	X	\checkmark
too many tidal-dwarf galaxies due to hierarchical formation	X	?
invariant baryonic galaxies	X	\checkmark
Local Volume of galaxies	X	?
"missing mass" in young tidal dwarf galaxies	X	\checkmark
surface density of dark matter / baryonic matter = constant	X	\checkmark
PLUS (not covered here)		
core/cusp problem	X	\checkmark
many bulge-less disk galaxies	X	?
missing satellite problem	X	?
downsising	X	?





Remember this slide ?



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It is irrelevant to debate whether the LCMD model fits any data. Beauty or even "high-precision" of a model can misguide.

It is therefore *irrelevant* to argue whether *the CMB*

or

large-scale structure are consistent with the LCDM model





It is a pleasure to welcome to this debate :

Prof. R.H. (Bob) Sanders (Groningen)



Prof. Gerhardt Hensler (Vienna)



Dark Matter and MOND

Book on "The Dark Matter Problem", Cambridge University Press, 2010 Star formation in galaxies, chemo-dynamical evolution

Prof. Tom Shanks (Durham)



CMB, Dark Energy

The new baryonic structure formation scenario



