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# FERMI<sup>ONS</sup> IN WARPED SPACETIME AND THE SMALLNESS OF THE COSMOLOGICAL CONSTANT

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*RK et. al. Phys. Rev. D 79 (R), 041902 (2009)*

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## OUTLINE

- ◆ Warped spacetime :
  - Braneworlds : Randall-Sundrum (RS) model
  - Solution of gauge hierarchy problem
  - Generalised RS
- ◆ Standard model fields in warped spacetime
  - Fermions on brane
  - The mechanism of confinement
  - Localization in generalised RS
- ◆ Smallness of cosmological constant: linked with fermion confinement
- ◆ Summary

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## WARPED SPACETIME : BRANEWORLDS

- ◆ The basis of Einstein's theory of General Relativity
- ◆ Most familiar example : **FRW metric**

$$ds^2 = -dt^2 + a^2(t)(dr^2 + r^2d\theta^2 + r^2\sin^2\theta d\phi^2) \quad (1)$$

- ◆ Example of a five dimensional **warped** spacetime :

$$ds^2 = d\sigma^2 + e^{2f(\sigma)}(-dt^2 + dx^2 + dy^2 + dz^2) \quad (2)$$

⇒ The 4D spacetime is warped by the factor  $f(\sigma)$

- ◆ Phenomenological models of extra dimension  
motivated by string theory : **Randall-Sundrum model**

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- ◆ Phenomenological models of extra dimension  
motivated by string theory : **Randall-Sundrum model**

**Gives a novel solution of the  
Gauge Hierarchy problem**

## Island Universes in Warped Space-Time

According to string theory, our universe might consist of a three-dimensional "brane," embedded in higher dimensions. In the model developed by Lisa Randall and Raman Sundrum, gravity is much weaker on our brane than on another brane, separated from us by a fifth dimension. (Time is the unseen fourth dimension.)

GRAVITY BRANE  
(where gravity is concentrated)

### Fifth dimension

Space is warped by energy throughout five-dimensional space-time. As a result, gravity is much weaker on our brane.

### Gravitions,

which transmit gravity, are closed strings, which are not confined to either brane.



The ends of **open strings**, whose oscillations are particles and forces other than gravity, are stuck to our brane.

BRANE  
(our universe)

### Warped space-time

Because space-time is warped, things are exponentially bigger and lighter closer to our brane.

Figure courtesy : cosmicvariance by S. M. Carroll

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## RS MODEL : SOLUTION OF GAUGE HIERARCHY PROBLEM

Some intriguing question about our physical universe

- Why does it appear to have (3 +1) spacetime dimensions ? Are there additional unobserved dimensions?
  - Why the ratio of electroweak scale/Higgs mass ( $m$ ) to Planck mass ( $m_0$ ) so tiny ( $\simeq 10^{-16}$ ) ?  $\Rightarrow$  *Gauge Hierarchy Problem*
- ◆ **Randall-Sundrum** model : 5D spacetime

$$ds^2 = r_c^2 d\sigma^2 + e^{-2kr_c|\sigma|}(-dt^2 + dx^2 + dy^2 + dz^2) \quad (3)$$

- ◆ **RS prescription** : masses get scaled on the visible brane by the warp factor as

$$m = e^{-kr_c\pi} m_0 \quad \text{where} \quad kr_c \simeq 11.84$$

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## GENERALISED RS MODEL

RS → visible brane is flat → inconsistent with present value

Generalised RS model :

- ◊ visible brane → **non-flat**
- ◊ visible brane tension → can be both positive/negative
- ◊ The 5D spacetime metric :

$$ds^2 = e^{-2A(y)} g_{\mu\nu} dx^\mu dx^\nu + r^2 dy^2 \quad (4)$$

- ◊ The warp factor  $e^{-2A(\sigma)}$  obtained by extremising the action :

$$S = \int d^5x \sqrt{-G} (M^3 \mathcal{R} - \Lambda) + \int d^4x \sqrt{-g_i} \mathcal{V}_i \quad (5)$$

$\Lambda$  → bulk cosmological constant,  $\mathcal{R}$  → bulk Ricci scalar  
and  $\mathcal{V}_i$  → tension of the  $i^{th}$  brane.

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◊ For AdS bulk ( $\Lambda < 0$ ) & -ve induced curvature  $\Omega$  on visible brane

$$e^{-A} = \omega \cosh \left( \ln \frac{\omega}{c_1} + ky \right) \quad (6)$$

where  $k = \sqrt{\frac{-\Lambda}{12M^3}}$ ,  $\omega^2 \equiv -\Omega/3k^2$  and  $c_1 = 1 + \sqrt{1 - \omega^2}$

◊ The hierarchy problem has been solved for

$$k\pi r_1 \simeq 36.84 \times 10^{-93}, \quad k\pi r_2 = 250.07$$

◊ visible brane tension  $\rightarrow +$  ve for second solution.

◊ For AdS bulk ( $\Lambda < 0$ ) & +ve induced curvature  $\Omega$  on visible brane

$$e^{-A} = \omega \sinh \left( \ln \frac{c_2}{\omega} - ky \right) \quad (7)$$

where  $\omega^2 \equiv \Omega/3k^2$  and  $c_2 = 1 + \sqrt{1 + \omega^2}$ .

Curvature on the brane is parametrized by  $\omega^2$

Brane cosmological constant, can be of arbitrary magnitude.

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# STANDARD MODEL FIELDS IN WARPED SPACETIME

## STUDY LOCALIZATION : why?

Extra dimensions are hidden from us

⇒ Kaluza-Klein compactification

Assumption in warped braneworld models:

- ◆ Gravity can access the whole bulk spacetime
- ◆ But SM fields are localized on the brane

The idea is borrowed from string theory (D-branes)

Localization of fields ⇒ alternative to compactification

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**Are the fields really localized  
in different braneworld models ?**

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## THE MECHANISM : FOR FERMIONS

- ◆ Consider that the field can access the whole bulk

Consider fermion :  $\Psi(x^\mu, y) = \psi(x^\mu)\xi(y)$

- ◆ Express the action of the field in canonical form

$$\int \sqrt{-g} \mathcal{L}_{Dirac} d^5x = \int \sqrt{-g} (i\bar{\Psi}\Gamma^a \mathcal{D}_a \Psi) d^5x$$

- ◆ Extract out the standard four dimensional part

$$i\gamma^\mu \partial_\mu \psi(x^\mu) = m\psi(x^\mu)$$

- ◆ Rest of the integral gives the localization condition

$$\int e^{-3A(y)} \xi_m \xi_n dy = \delta_{mn}$$

- ◆ Find the functional dependence on higher dimension  $\rightarrow \xi(y)$

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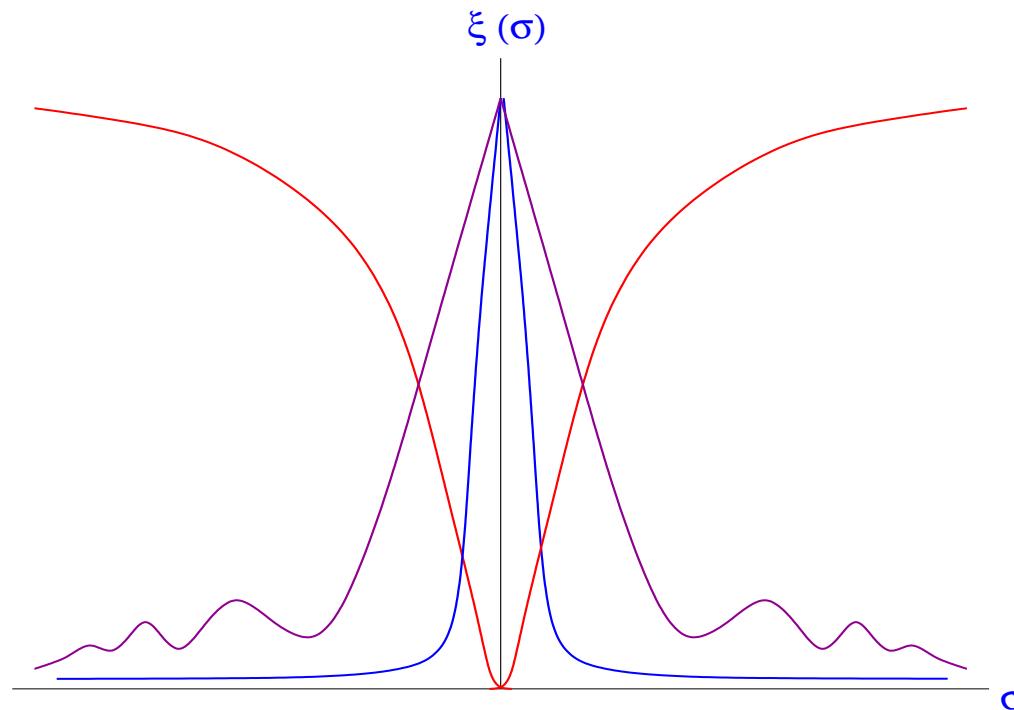
- ◆ Find the functional dependence on higher dimension  $\rightarrow \xi(y)$

**Whether this function,  $\xi(y)$ , satisfies  
the localization condition ?**

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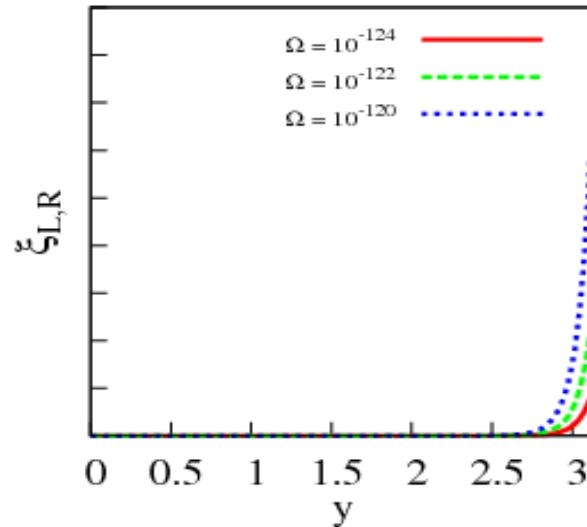
## Functional dependence & localization scenario

- ◆ Function sharply peaked at the brane  $\Rightarrow$  Localized
- ◆ Function not sharply peaked at the brane  $\Rightarrow$  Quasi localized
- ◆ Function grows along the extra dimension  $\Rightarrow$  Not localized



## FERMION LOCALISATION IN GENERALISED RS MODEL

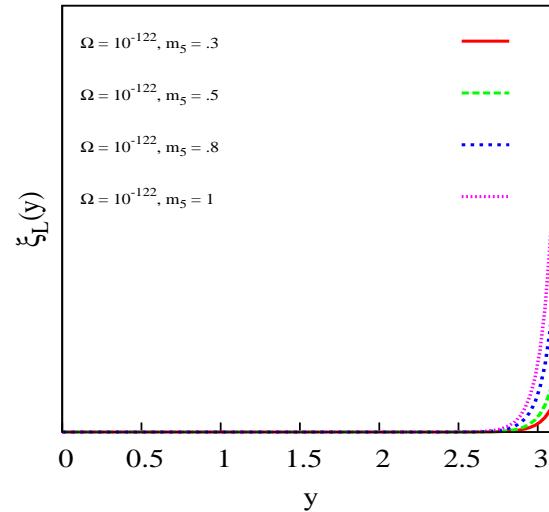
- ◆ Massless bulk fermion in AdS 3-brane model: left and right chiral massless modes turn out to be  $\xi_{L,R}(y) = N_1 \operatorname{sech}^2 \left( \ln \frac{\omega}{c_1} + ky \right)$



⇒ As magnitude of  $\Omega$  becomes close to  $10^{-124}$  the zero modes get sharply localized on the visible brane with -ve brane tension.

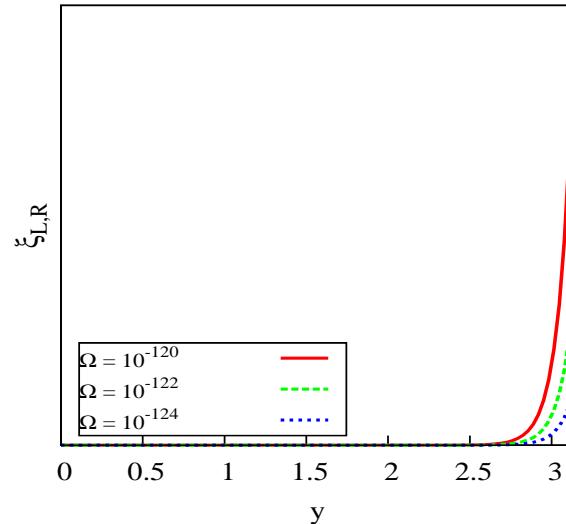
But, for positive tension brane fermions are clearly localized deep inside the bulk space-time

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- ◆ Massive bulk fermion in AdS 3-brane model: left and right chiral massless modes :  $\xi_{L,R}(y) = N_3 \operatorname{sech}^2 \left( \ln \frac{\omega}{c_1} + ky \right) e^{\pm m_5 y}$



- ▲ Degeneracy between two chiral modes lifted by the mass term.
- ▲ For a small value of  $\Omega$  as bulk fermion becomes more and more massive left chiral mode has higher peak values on visible brane
- ▲ The right chiral mode shows exactly the reverse nature.

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- ◆ Massless bulk fermion in dS 3-brane model: left and right chiral massless modes :  $\xi_{L,R}(y) = N_2 \operatorname{cosech}^2 \left( \ln \frac{c_2}{\omega} + ky \right)$



- ▲ As the value of  $\Omega$  becomes smaller the fermions become more sharply localized on the visible brane.
- ▲ Thus the fact that our universe has a very small and positive curvature  $\Rightarrow$  the bulk fermions are likely to get localized near the visible brane just as predicted in string theory.

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## SUMMARY

- Studied crucial features of a brane with non-zero brane curvature
- Profiles of massless and massive modes of bulk fermions studied in generalised RS model
- Fermions get localized near the visible brane for a very small positive value of the brane curvature.
- This is an important finding for the explanation of the smallness of the cosmological constant in connection with the brane localization of the fermion in phenomenological braneworld model.