(7 points)

(5 points)

## Exercises on General Relativity and Cosmology

Dr. Hans Jockers, Prof. Dr. Hans-Peter Nilles

http://www.th.physik.uni-bonn.de/klemm/gr\_15/

## -Home Exercises-

## Due on July 6, 2015

**H 12.1 Falling freely inside a black hole** (15 points) Consider a particle that has crossed the event horizon of a Schwarzschild black hole at  $r = R_S$  and is then falling freely within the event horizon. Work with the standard Schwarzschild coordinates  $(t, r, \theta, \phi)$  and denote the particle's proper time by  $\tau$ .

a) Show that

$$\left|\frac{\mathrm{d}r}{\mathrm{d}\tau}\right| \ge \sqrt{\frac{R_S}{r} - 1} \ . \tag{1}$$

When is this bound saturated?

- b) Calculate the maximum lifetime of a particle falling freely from  $r = R_S$  to r = 0. Plug in numbers to express this in seconds for a black hole with one solar mass,  $M = 2 \cdot 10^{30}$  kg. (4 points)
- c) Can an observer that also travels from  $r = R_S$  to r = 0 but not on a geodesic have a longer lifetime than the one calculated in item b)? Justify your answer. (4 points)

H 12.2 Particle motion in the presence of electromagnetic fields (15 points)As you have shown in H 9.1 b), in the presence of an electromagnetic field a particle of charge q and mass m moves along a curve C that obeys

$$\nabla_{\frac{\mathrm{d}\mathcal{C}}{\mathrm{d}\tau}} \frac{\mathrm{d}\mathcal{C}}{\mathrm{d}\tau} = \frac{q}{m} F^{\mu}_{\ \nu}(\mathcal{C}(\tau)) \frac{\mathrm{d}\mathcal{C}^{\nu}}{\mathrm{d}\tau} \partial_{\mu} , \qquad (2)$$

in terms of the particle's proper time  $\tau$ .

a) Given a Killing vector field K, show that

$$\nabla_{\frac{\mathrm{d}\mathcal{C}}{\mathrm{d}\tau}} g\left(\frac{\mathrm{d}\mathcal{C}}{\mathrm{d}\tau}, K\right) = g\left(\frac{q}{m} F^{\mu}_{\ \nu}(\mathcal{C}(\tau)) \frac{\mathrm{d}\mathcal{C}^{\nu}}{\mathrm{d}\tau} \partial_{\mu}, K\right) \ . \tag{3}$$

Now imagine that such a particle is moving in the field of a Reissner-Norström black hole of electric charge Q, magnetic charge P and mass M. From exercise H 10.1 recall the Reissner-Norström metric,

$$ds^{2} = -\Delta dt^{2} + \Delta^{-1} dr^{2} + r^{2} \left( d\theta^{2} + \sin^{2} \theta d\phi^{2} \right) , \quad \Delta = 1 - \frac{2GM}{r} + \frac{G\left(Q^{2} + P^{2}\right)}{r^{2}} , \quad (4)$$

and the non-vanishing components of the field strength tensor,

$$F_{tr} = -F_{rt} = -\frac{Q}{r^2} , \quad F_{\theta\phi} = -F_{\phi\theta} = P\sin\theta .$$
(5)

b) Use the Killing vector field corresponding to time-translation symmetry and eq. (3) to show that the energy

$$E = m\left(1 - \frac{2GM}{r} + \frac{G\left(Q^2 + P^2\right)}{r^2}\right)\frac{\mathrm{d}t}{\mathrm{d}\tau} + \frac{qQ}{r} \tag{6}$$

is conserved.

c) Use the Killing vector field corresponding to the symmetry under rotations along the angle  $\phi$  and eq. (3) to show that the angular momentum

$$L_{\phi} = m r^2 \sin^2 \theta \, \frac{\mathrm{d}\theta}{\mathrm{d}\tau} - q P \cos \theta \tag{7}$$

is conserved.

 $(5 \ points)$ 

 $(5 \ points)$