
Exercises on General Relativity and Cosmology

Priv. Doz. Dr. S. Förste

–HOME EXERCISES–
DUE 29 JUNE 2011

Exercise 10.1: Einstein equation (2 credits)

- (a) In a local region of spacetime, an observer finds that the Ricci scalar is nearly constant, $R \approx +1/a^2$. What can you say about the equations of state of the matter sourcing this curvature? If the region of spacetime is filled only with electromagnetic energy, what is R ? (1 credit)
- (b) Is it possible to have a solution of the Einstein equations, in which space is empty to the past of some surface of constant time $t = 0$, but in which there is a nonvanishing $T_{\mu\nu}$ to the future of this surface? (1 credit)

Exercise 10.2 Motion in Schwarzschild geometry (18 credits)

- (a) A particle falls radially into a Schwarzschild metric.
- (i) As measured by proper time at infinity, what is its inward coordinate velocity (dr/dt) at a (curvature-) radius r ? (3 credits)
 - (ii) What is the locally-measured velocity relative to a stationary observer at the same radius? (2 credits)
- (b) Derive the equations of motion (ie. equations relating t , r and proper-time τ) for a particle falling radially in the Schwarzschild geometry. Consider the three cases:
- (i) particle released from rest at $r = R$ (3 credits)
 - (ii) particle released from rest at infinity (3 credits)
 - (iii) particle projected inward from infinity with velocity v_∞ . (2 credits)
- (c) Show that the trajectory of light rays in the Schwarzschild metric obeys:

$$\frac{d^2u}{d\phi^2} + u = 3u^2$$

where, $u = M/r$ and r is the Schwarzschild radial coordinate. Denote the minimum value of r along the trajectory by b , the *impact parameter*. In case of $(M/b) \ll 1$, what is the deflection of a photon as it passes a spherical gravitating body? Give a formula for the deflection angle to lowest nonvanishing order in (M/b) . (5 credits)