# Exercises on General Relativity and Cosmology

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## -CLASS EXERCISES-

### Exercise 0.1: Lorentz boosts

Consider a boost in the x-direction with speed  $v_A = \tanh(\alpha)$  followed by a boost in the y-direction with a speed  $v_B = \tanh(\beta)$ . Show that the resulting Lorentz transformation is the same as doing a pure rotation followed by a pure boost, and determine the rotation and boost.

### Exercise 0.2: Electromagnetism

Maxwell's equations are given by

$$\overrightarrow{\nabla} \cdot \overrightarrow{E} = \epsilon, \quad \overrightarrow{\nabla} \times \overrightarrow{B} = \frac{\partial \overrightarrow{E}}{\partial t} + \overrightarrow{J}, \quad \overrightarrow{\nabla} \cdot \overrightarrow{B} = 0, \quad \overrightarrow{\nabla} \times \overrightarrow{E} = -\frac{\partial \overrightarrow{B}}{\partial t}$$

To make their properties under Lorentz transformation explicit, we can choose an antisymmetric  $4 \times 4$  tensor,  $F^{\mu\nu}$ , the electromagnetic field tensor and the (charge and current) source density four-vector,  $J^{\mu}$ , as:

$$F^{\mu\nu}(t, \mathbf{x}) = \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & B_3 & -B_2 \\ -E_2 & -B_3 & 0 & B_1 \\ -E_3 & B_2 & -B_1 & 0 \end{pmatrix}$$
$$J^{\mu}(t, \mathbf{x}) = (\epsilon, \mathbf{J}) = \left(e\delta^3(\mathbf{x} - \mathbf{x}(t)), \ e\delta^3(\mathbf{x} - \mathbf{x}(t))\frac{d\mathbf{x}(t)}{dt}\right)$$

- (a) Show that  $\partial_{\mu}F^{\mu\nu} = -J^{\nu}$  and  $\epsilon^{\mu\nu\rho\sigma}\partial_{\nu}F_{\rho\sigma} = 0$  reproduce Maxwell's equations. What are the Lorentz invariants that can be constructed?
- (b) Verify in the rest frame that

$$f^{\mu} \equiv \frac{dp^{\mu}}{d\tau} = eF^{\mu}_{\ \nu}\frac{dx^{\nu}}{d\tau}$$

is the correct equation for the electromagnetic four-force  $f^{\mu}$  acting on a charged particle.  $(p^{\mu} = m \ dx^{\mu}/d\tau)$ 

### Exercise 0.3: Energy-momentum tensor

In analogy to the electrical charge and current density vector above, we can define a "charge" and "current density" for the matter 4-momentum,  $p^{\mu}$ , the matter energy-momentum tensor:

$$T_m^{\mu\nu}(\mathbf{x},t) \equiv p^{\mu}(t) \frac{dx^{\nu}(t)}{dt} \delta^3(x - x(t))$$

(a) Show that the energy-momentum tensor is only conserved up to a force density,  $G^{\mu}$  which vanishes for free particles:

$$\partial_{\nu}T^{\mu\nu} = G^{\mu}$$

- (b) Check that for the (interacting) electromagnetic quantities given above, we get  $G^{\mu}$  to be  $F^{\mu}_{\ \nu}J^{\nu}$ .
- (c) The electromagnetic energy-momentum tensor was defined in the lecture as:

$$T^{\mu\nu}_{em} \equiv F^{\mu}_{\ \rho} F^{\nu\rho} - \frac{1}{4} \eta^{\mu\nu} F_{\rho\sigma} F^{\rho\sigma}$$

Show that the divergence of this cancels with that of  $G^{\mu}$  defined in (b), so that  $T_m^{\mu\nu} + T_{em}^{\mu\nu}$  is conserved.