2c) Perturbation theory, Feynman rules

Start with scalar field theory.

To make contact with non-relativistic scattering theory, write e.o.m.

 $(\partial_{\mu}\partial_{\mu}^{m}+m_{\psi}^{2})\psi = -V_{k6}\psi$  (2.25)

The transition matrix element (finallinitial) is in 1storder

(justification: QIT):

 $\begin{array}{c} (1) & - & - \\ H_{f_{1}}^{(1)} & = & - \\ & - & \int d^{4} \times d^{4} \int V_{k} d^{4} \cdot (2.26) \end{array}$ 

This holds for scattering on "external" polarial VKG; in our

i) & theory (implest cose)

 $(2.23) = 7 (\partial_{\mu}\partial^{r} + m_{\mu}^{2})d = -\frac{1}{6}d^{3}$ (2.27)

Want to describe "2 >2" scatting: K2 1/2

 $\phi(k_1) + \phi(k_1) \longrightarrow \phi(p_1) + \phi(p_1)$ 

let's assume (2.27) is used to describe ewlytion of \$ (K,) The \$ factor must the describe \$ (K2), \$ (P1), \$ (P2)

There are 
$$3! = 6$$
 ways to distribute these  $3$  fields over the  $d^3$   
factor => Explaints normalizations in (2, L1)  
 $A_{f}^{(1,04)}$  =  $-\frac{1}{2} \int d^4x \, d_x(x) \, d_y(x) \, d^4(x) \, d^4_p(x)$  (2, 28)

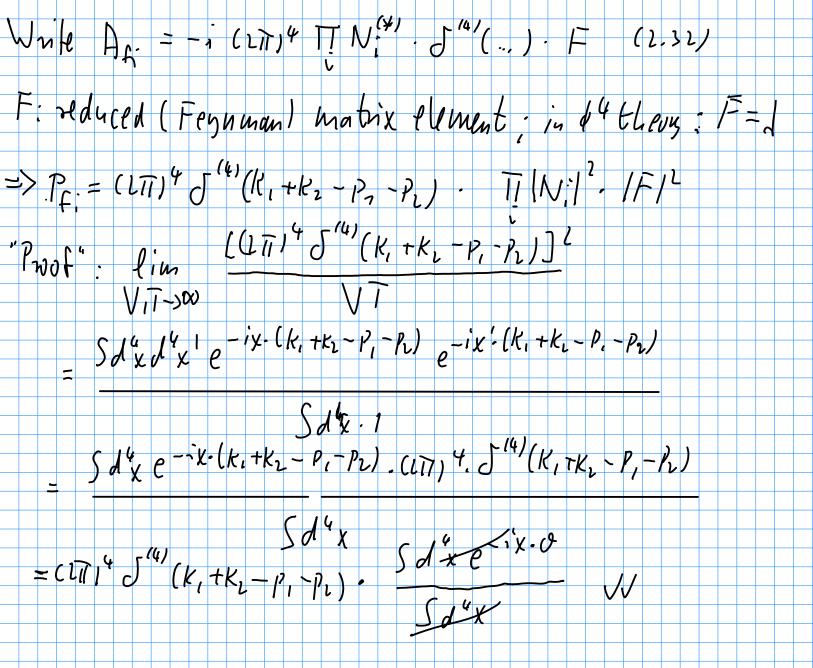
Perturbative expansion is an expansion in powers of the coupling constant! To 1st order, can use free fields on r.h.s. of (2.28):

$$\frac{d_{p}(x) = N(p) e^{-ip \cdot x} (1.6)}{\sum_{r=1}^{n} \frac{d_{r}}{d_{r}}} = -iA N(k_{r}) N(k_{r}) N(p_{r}) N(p_{r}) \int d^{4}x e^{i(p_{r} + p_{r}) - k_{r} - k_{r}} \cdot x$$

 $-i \lambda N(t_1) \cdots N^{*}(P_1) \cdot (1\pi)^{4} \int^{(4)} (K_1 + K_2 - P_1 - P_1) (2.2g)$ 

Note: up to "trivial" factors (normalization, 4-mom. conserv.) the transition amplitude is just given by the coupling!

Rate of transitions initial --> final ~ 
$$|H_{f_i}|^{-}$$
 more precisely:  
Wout to calculate the cross section!  
Normalization: Probability density  $(1.7): g = 2.1NI^2 E$   
Take norm: 2E paticles in why V, i.e.  $S g(i) dX = 2E$   
 $=> N = \frac{1}{VV}$  (2.30)  
Note: V-factors dwp out in the full => V ->  $\infty dK$ .  
Transition rate per final and why  $M_{f_i} = \frac{|A_{F_i}|^2}{VT}$  (2.31)  
T: dwafion of interaction.



To form a quantity which is independent of experimental set-up, i.e. of the flux of "incident" and "target" particles: have to divide by these factors

\*) Flux of "beam" particles on a stationary target (in rest frame of "particle 2") = number of particles reaching target per unit area and unit time:  $2E_1 + 2E_2$ 

flux = 1 <del>d</del> . 2 <del>E</del> V **R** dlusites of particle <u>1</u> Setween Seam & tanget

Finally, need to multiply with final state phase space volume element. For standard normalization (2E particles in V), this is for a 2-particle final state:

$$\frac{\sqrt{d^{2}p_{1}}}{c_{L}\pi)^{5}} \frac{\sqrt{d^{2}p_{2}}}{2E_{1}} \frac{\sqrt{d^{2}p_{2}}}{C_{L}\pi)^{5}} \frac{P_{1} = (E_{1}, P_{1}); P_{2} = (E_{2}, P_{1})}{2E_{2}}$$

=> differential cross section:

$$d \overline{b} = P_{F} \cdot \frac{V}{2\overline{E}_{1}} \frac{V}{V} \cdot \frac{V}{2\overline{E}_{2}} \cdot \frac{d^{3}}{2\overline{E}_{1}} \frac{V}{(2\overline{n})^{3}} \frac{d^{3}}{2\overline{E}_{1}} \frac{V}{(2\overline{n})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{n})^{3}} \frac{2\overline{E}_{2}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{n})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{n})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{E}_{1})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{E}_{1})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{E}_{1})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{E}_{1})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{E}_{1})^{3}} \frac{d^{3}}{2\overline{E}_{2}} \frac{d^{3}}{(2\overline{E}_{2})^{3}} \frac{d^{3}}{(2\overline{E}_{2}$$

## Note:

- (4)

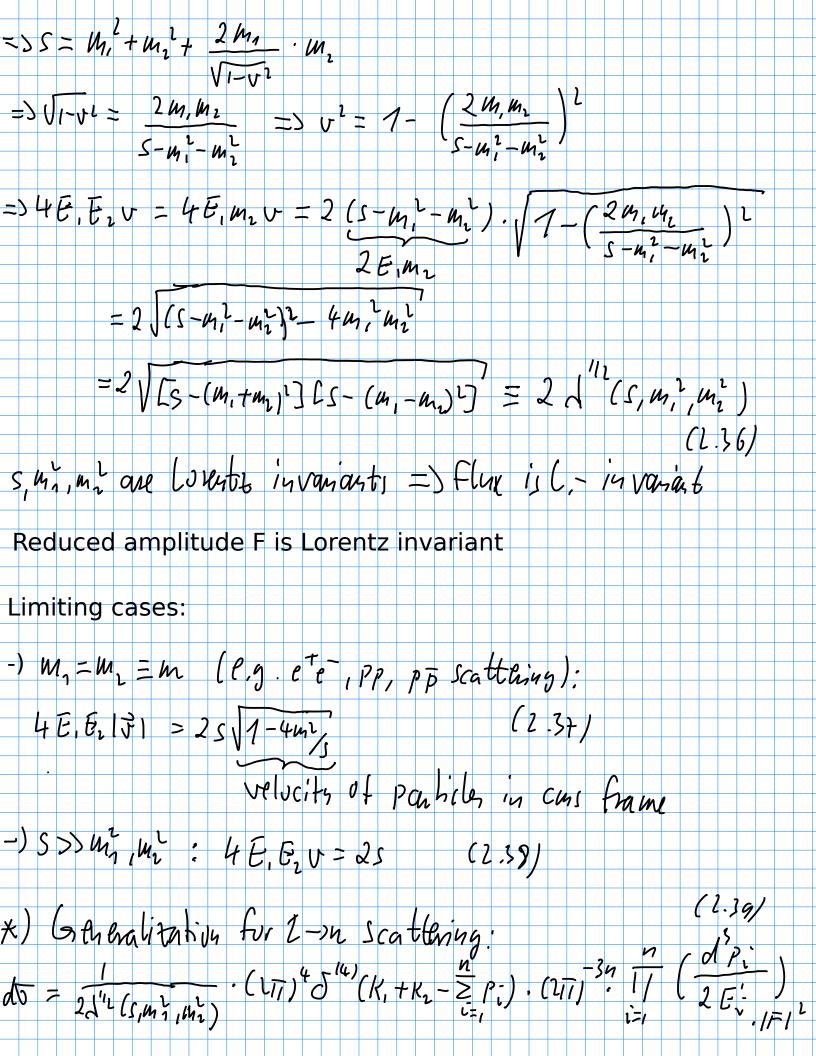
\*) V has dropped out, as advertized

\*) Expression is Lorentz invariant! 30

 $k_{1} = (E_{1}, \vec{k}_{1}); k_{2} = (u_{1}, \vec{0})$ 

$$K = E^{-1}(0, 0, v)$$
 in  $+z$  direction

Mondelstam-s =  $(K_1 + k_1)^2 = K_1^2 + K_1^2 + 2K_1 \cdot K_2 = M_1^2 + M_1^2 + 2E_1 M_2$ 



For experimentally useful quantity: need to integrate over sufficiently many final state momenta (components) to get rid of the delta-functions. For 2 --> 2 scattering:

\*) In cms frame:

 $\frac{1}{\sqrt{5}} = \frac{1}{\sqrt{5}} = \frac{1}{\sqrt{5}} = \frac{1}{\sqrt{5}} \frac{1$ (1.40) $\begin{array}{c|c} & & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ &$  $\frac{1}{P_{n}^{*}} = \frac{1}{P_{n}^{*}} \frac{1}{P_{n}$ 2-32 scattering cross sections have non-trivial of dependence outs if initial particles are turs venely polariold; Need Lobrak Symmetry unde what wis around beau ak's! \*) Using Lorentz invariants: Mandelstam t =  $(K_1 - \rho_1)^{t} = (K_1 - \rho_1)^{2}$  $\frac{d \overline{5}(1-51)}{dt} = \frac{1}{10 \overline{11} d (5, m_1^2, m_2^2)} \cdot (F_1^2) \cdot (1-42)$ ; f IFI has ho of dependence (=) Sd of # = 2 TT) \*) If final state particles are identical: not allowed to integrate over entire phase space, or divide by a symmetry factor. Reason:

(P1 P2) and (P2, P1) counst be dishinginght! Part. 1 Part. 2

For n identical particles in final state. need factor 4 (2.43) \*) 1->n trustions (particle decarp); Very similar reasoning, but normalization ("Flux") is how simply -= in rest frame of decaying particle:  $\frac{1}{2m} = \frac{1}{2m} \left( \frac{1}{1-3m} + \frac{1}{2m} + \frac{1}{1} + \frac{1}{3m} + \frac{1}{3m$ (In rest finne, not lovents invariant.)